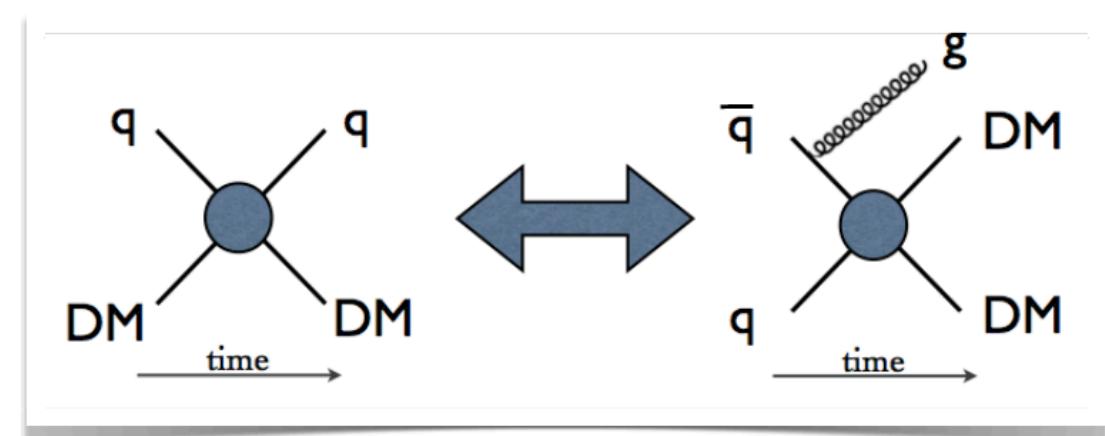
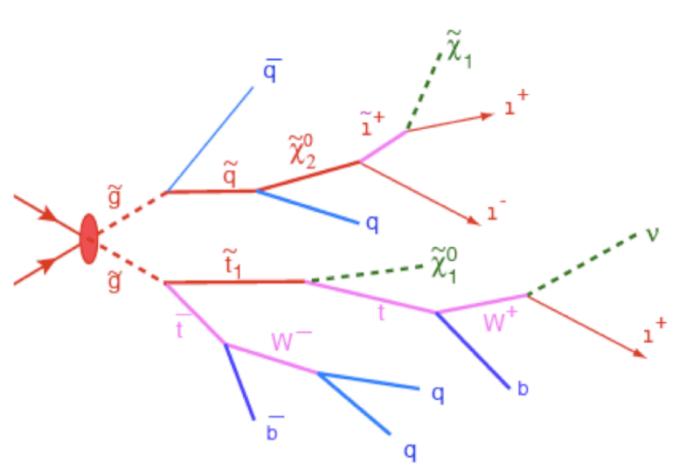


Oliver Buchmueller, Imperial College London

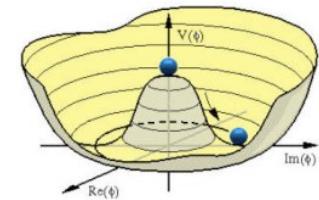
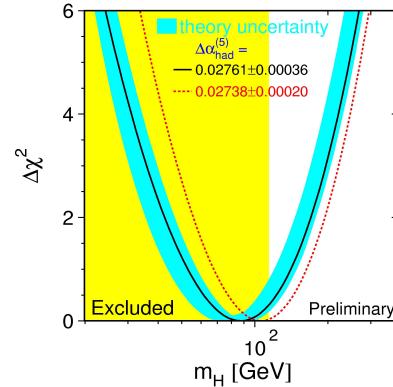
# SEARCHES FOR DARK MATTER PRODUCTION AT COLLIDER & DIRECT DETECTION EXPERIMENTS



# Fundamental Open Questions in Particle Physics

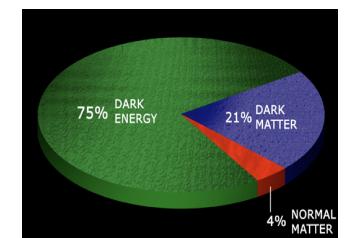
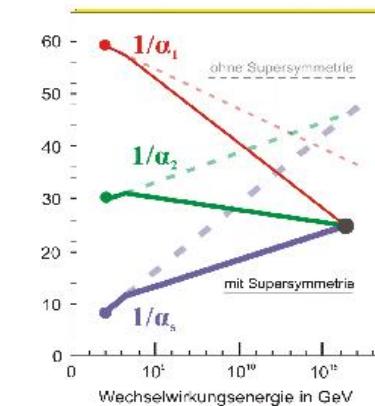
## I. What is the origin of mass?

- Why are the vector bosons  $Z$  and  $W$  massive whereas the photon is massless?
- Is there a Higgs boson - or even more of them ?



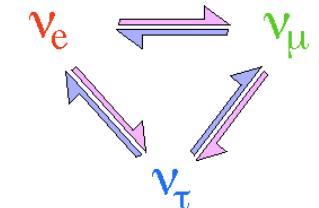
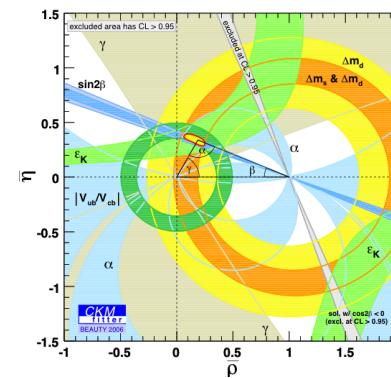
## II. What is the origin of Dark matter in our Universe ?

- Is a fundamental particle responsible for it?
- Is there a new symmetry in nature?  
=> Does Supersymmetry exist and can it explain DM?



## III. What is the origin of the matter-anti-matter asymmetry in our Universe?

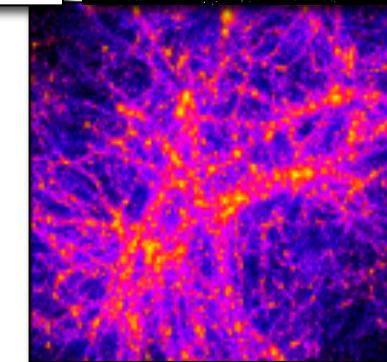
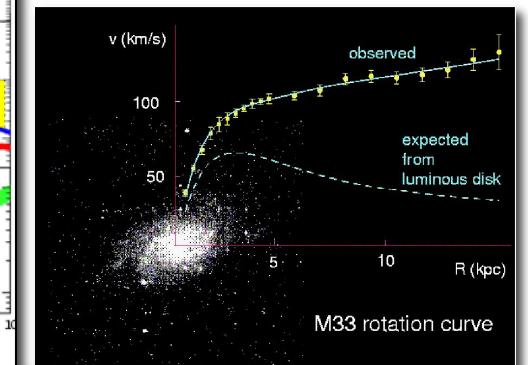
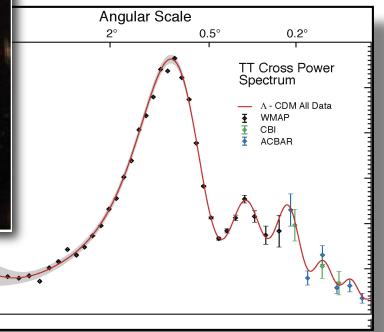
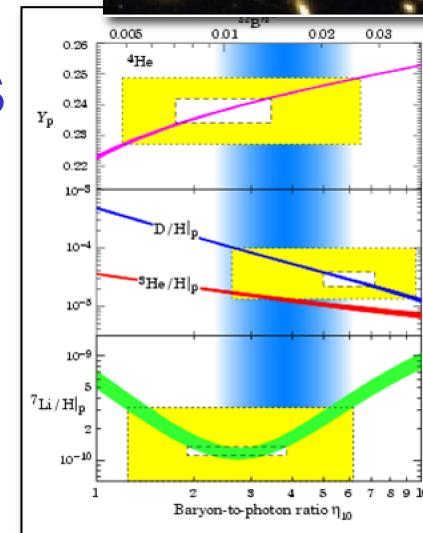
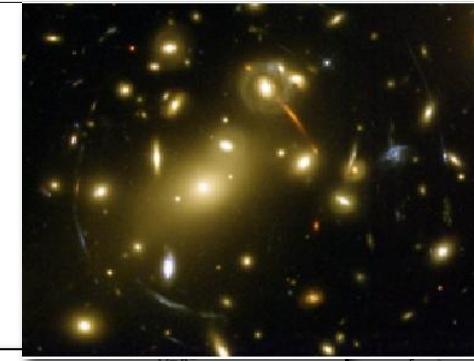
- Does the answer lie in CP violation?
- Neutrino masses and mixing - how do they fit in the picture?



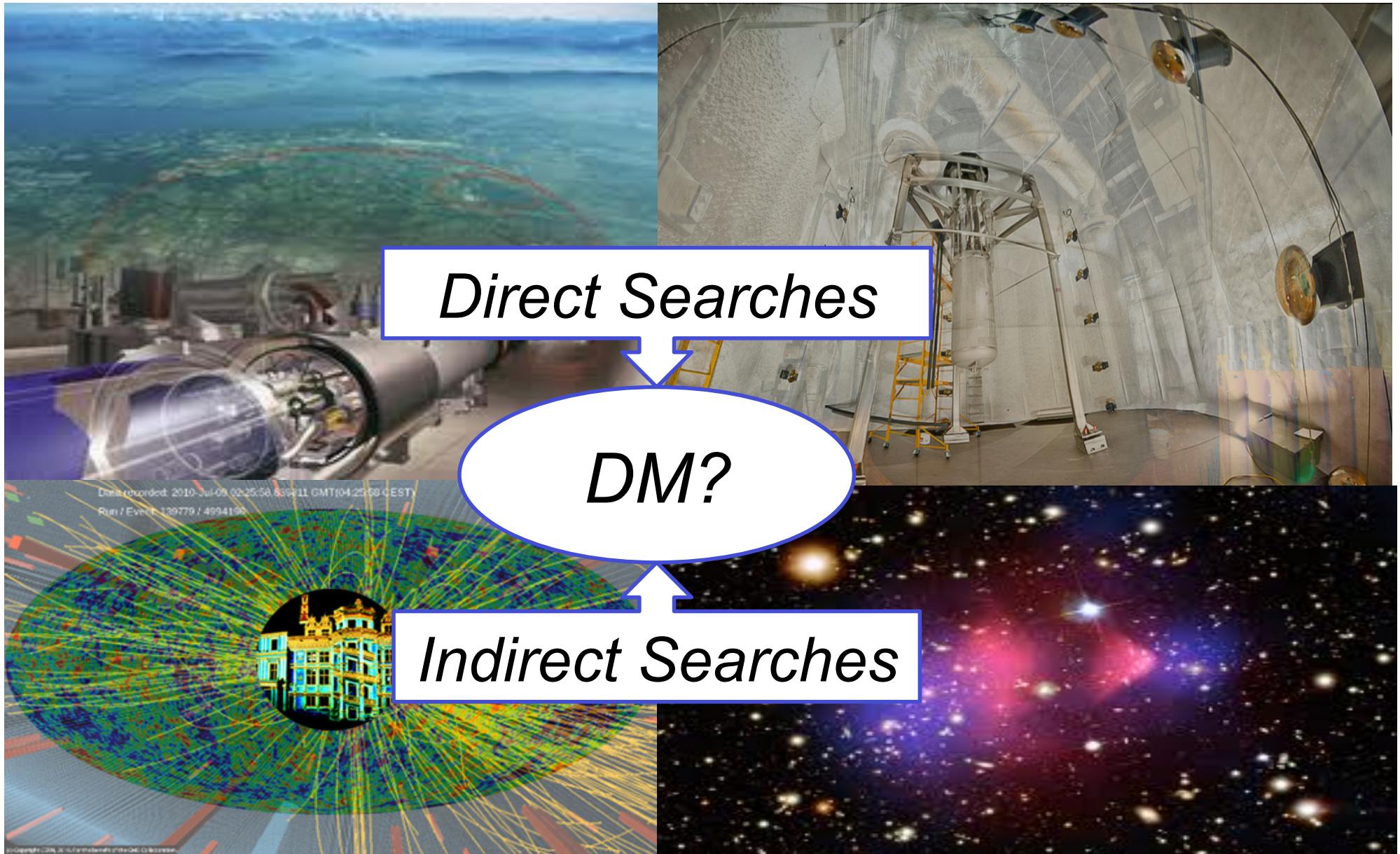
# Very Strong Evidence for Dark Matter Today!

DM Searches @ collider & Direct Detection O. Buchmüller

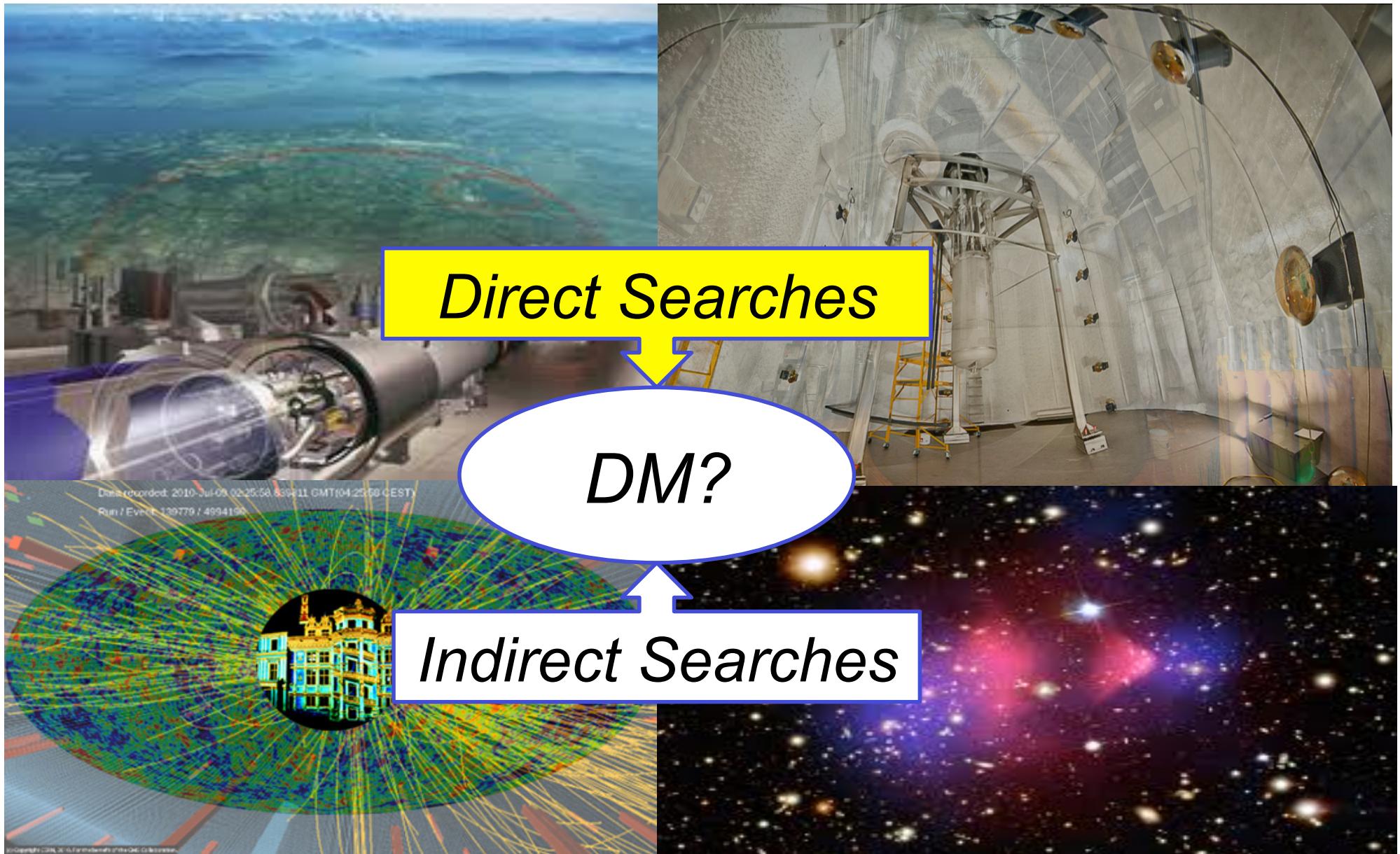
- Galactic rotation curves
- Gravitational lensing
- Light element abundances
- CMB anisotropies
- Large scale structure
- Etc...



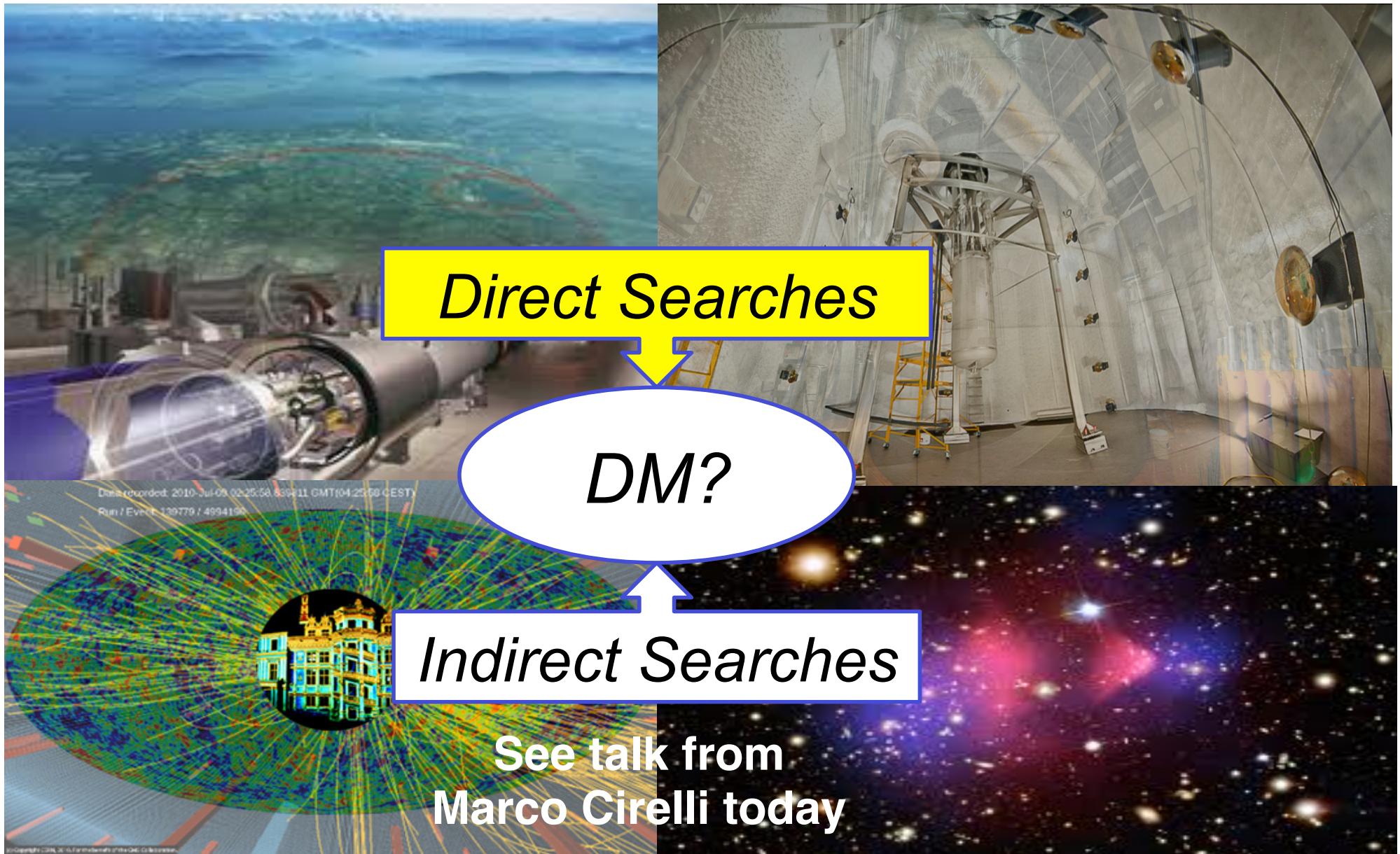
# Dark Matter Searches



# Dark Matter Searches



# Dark Matter Searches



# Dark Matter Searches at Colliders

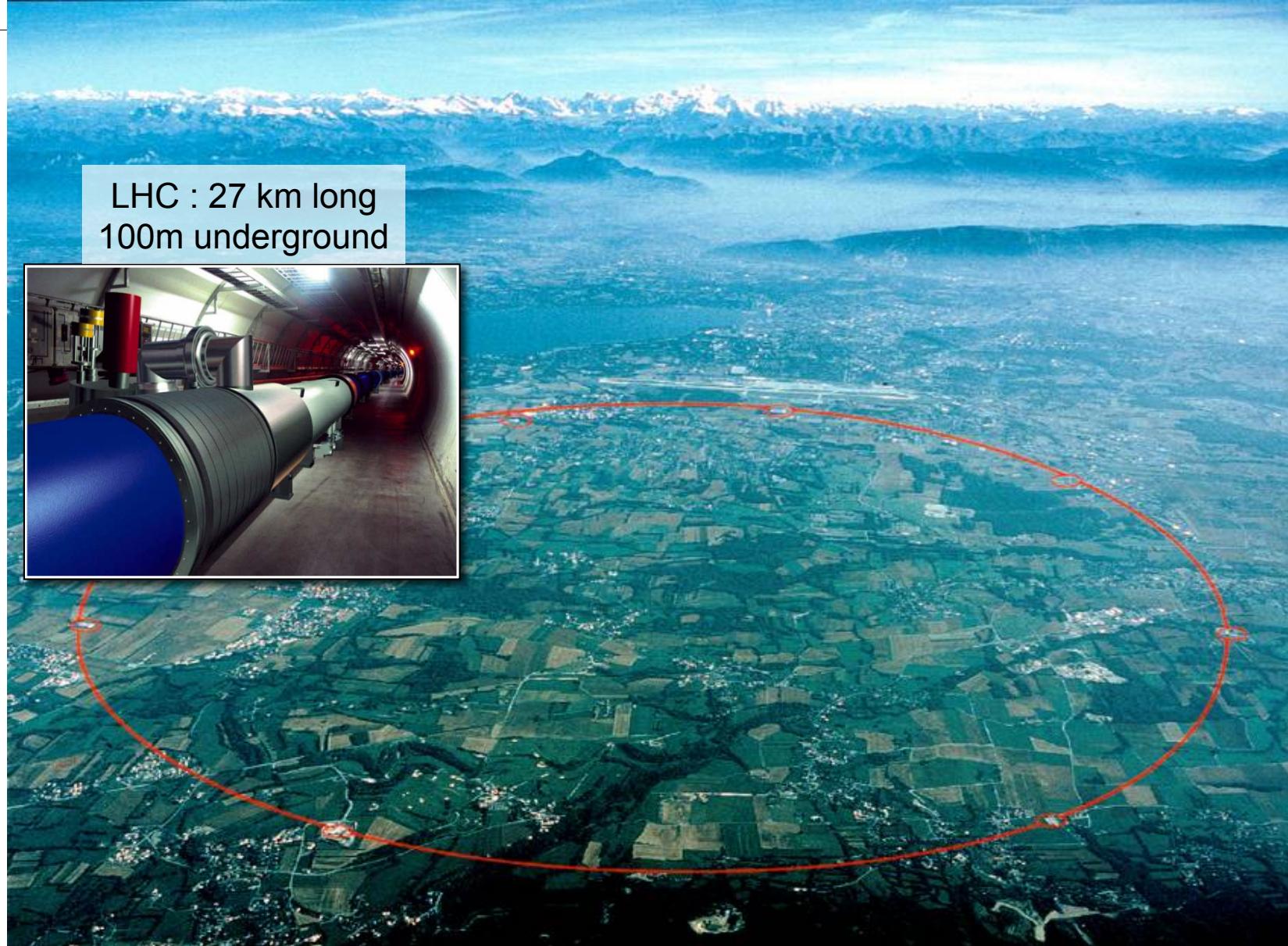


# *The Large Hadron Collider at CERN*

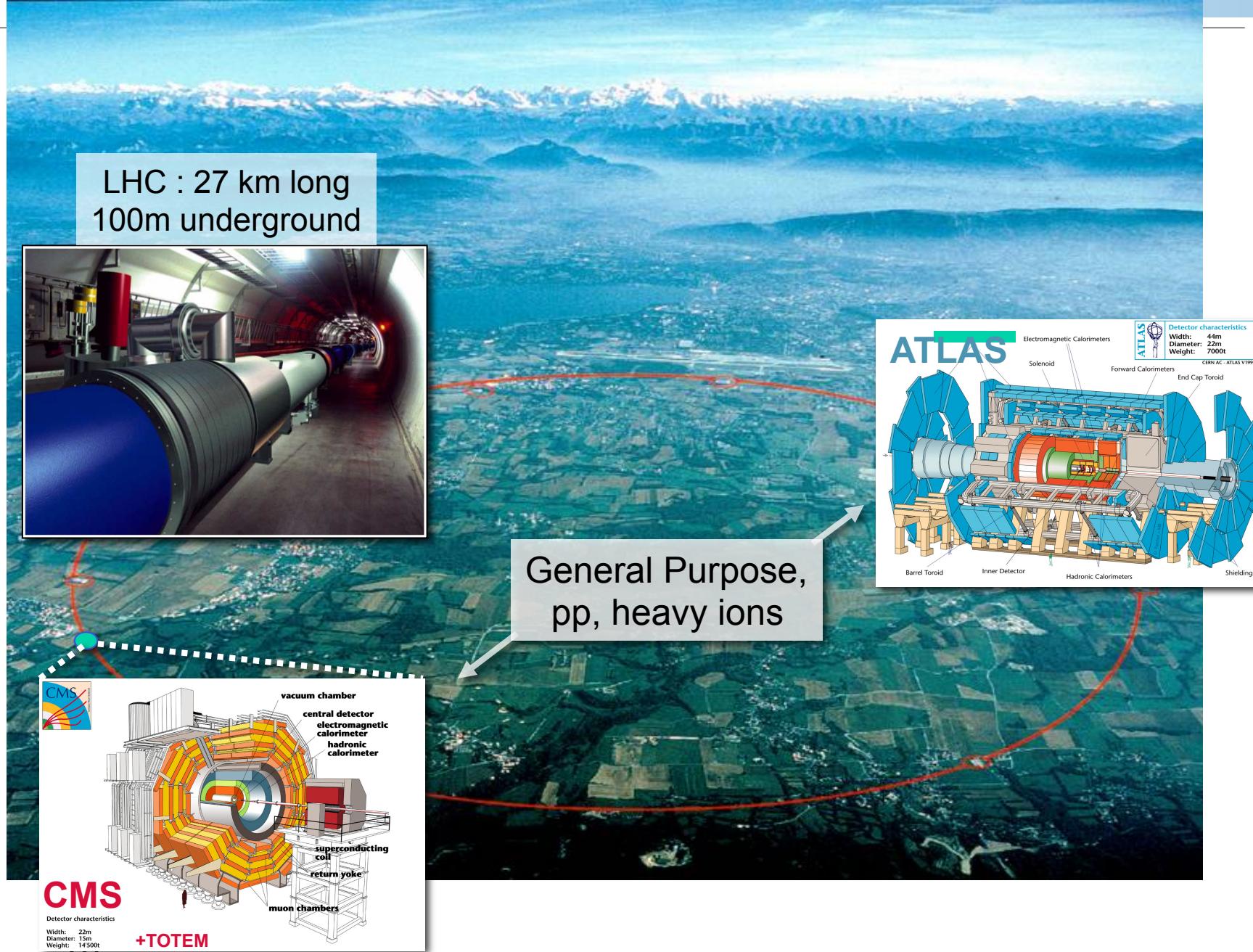
DM Searches @ collider & Direct Detection O. Buchmüller



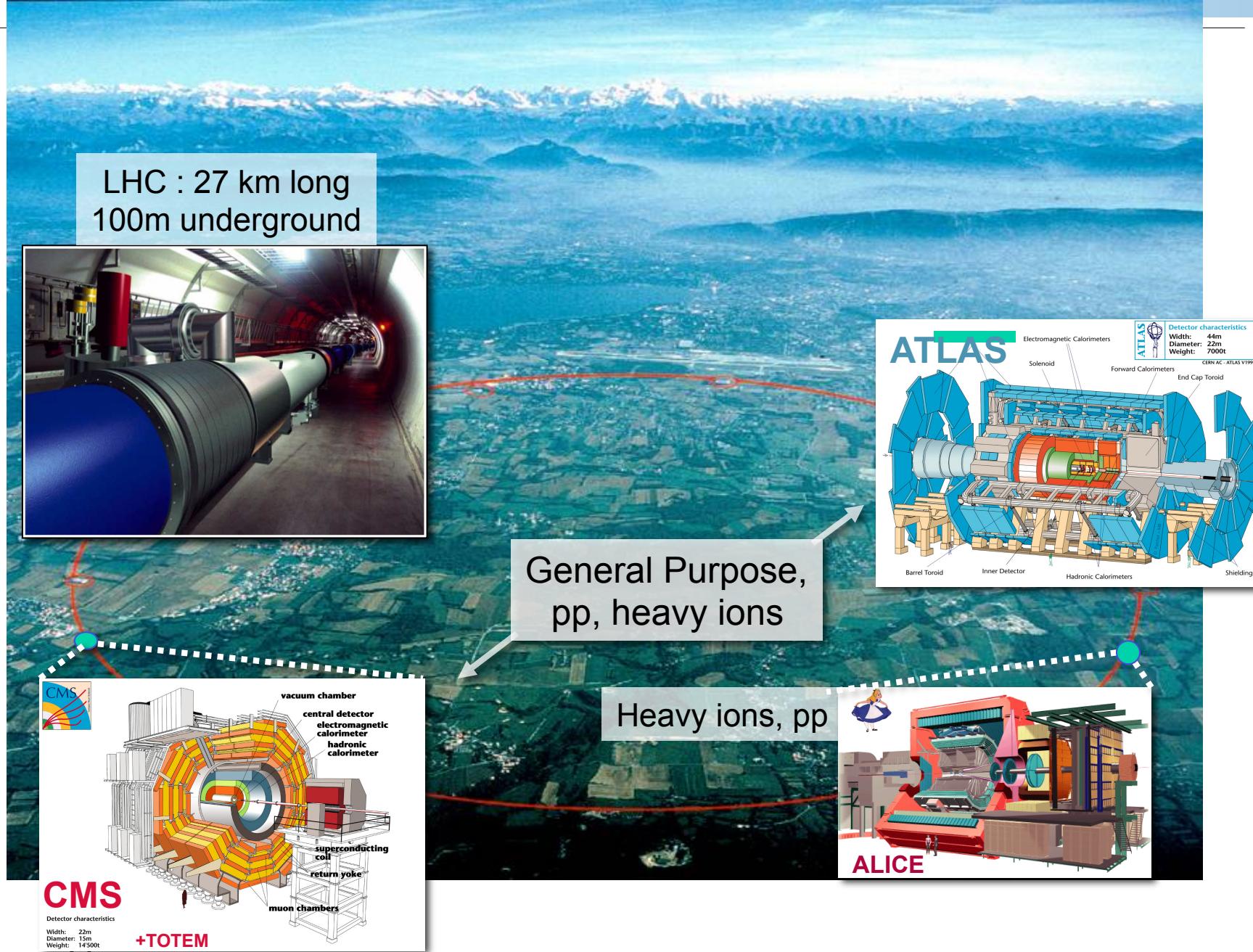
# The Large Hadron Collider at CERN



# The Large Hadron Collider at CERN

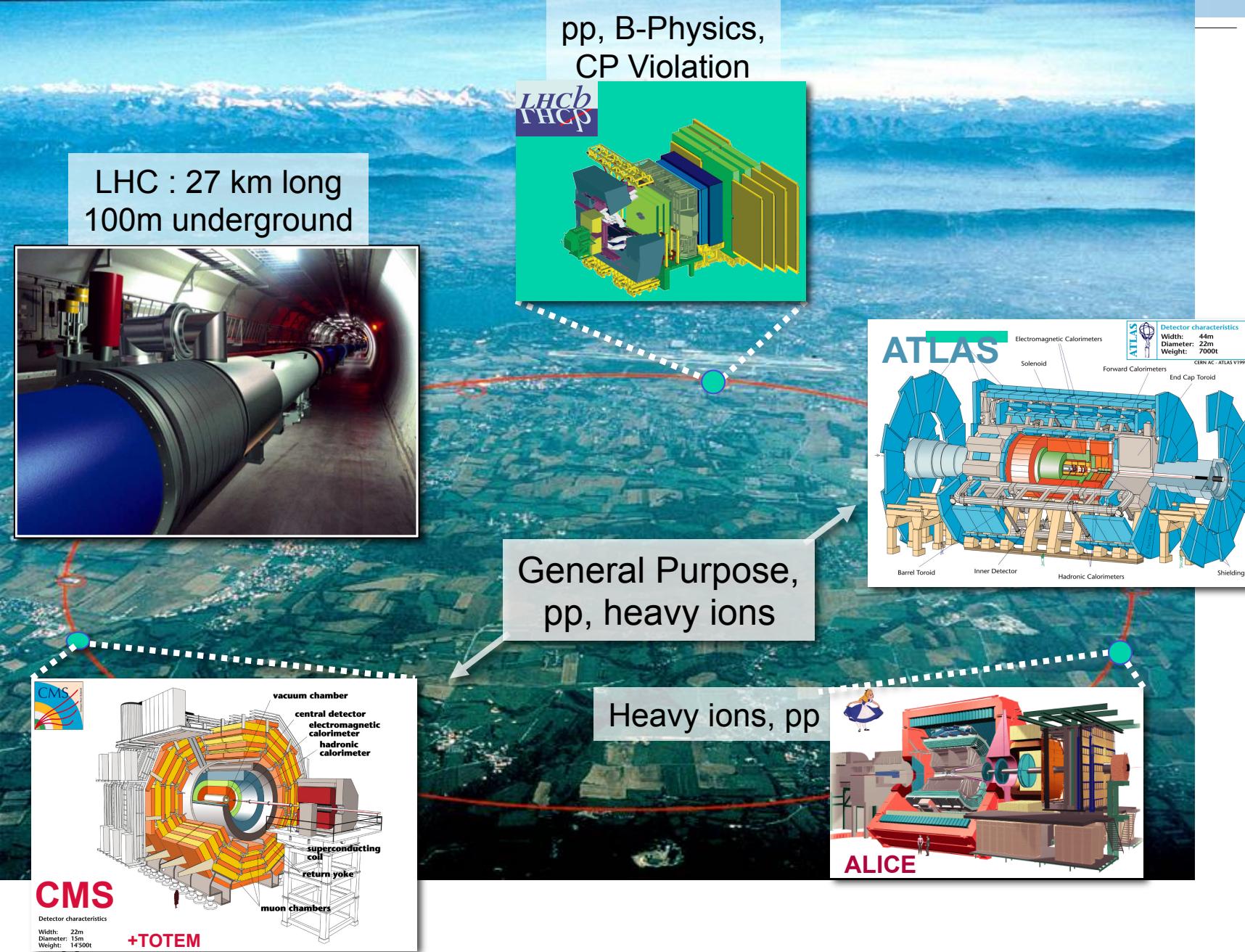


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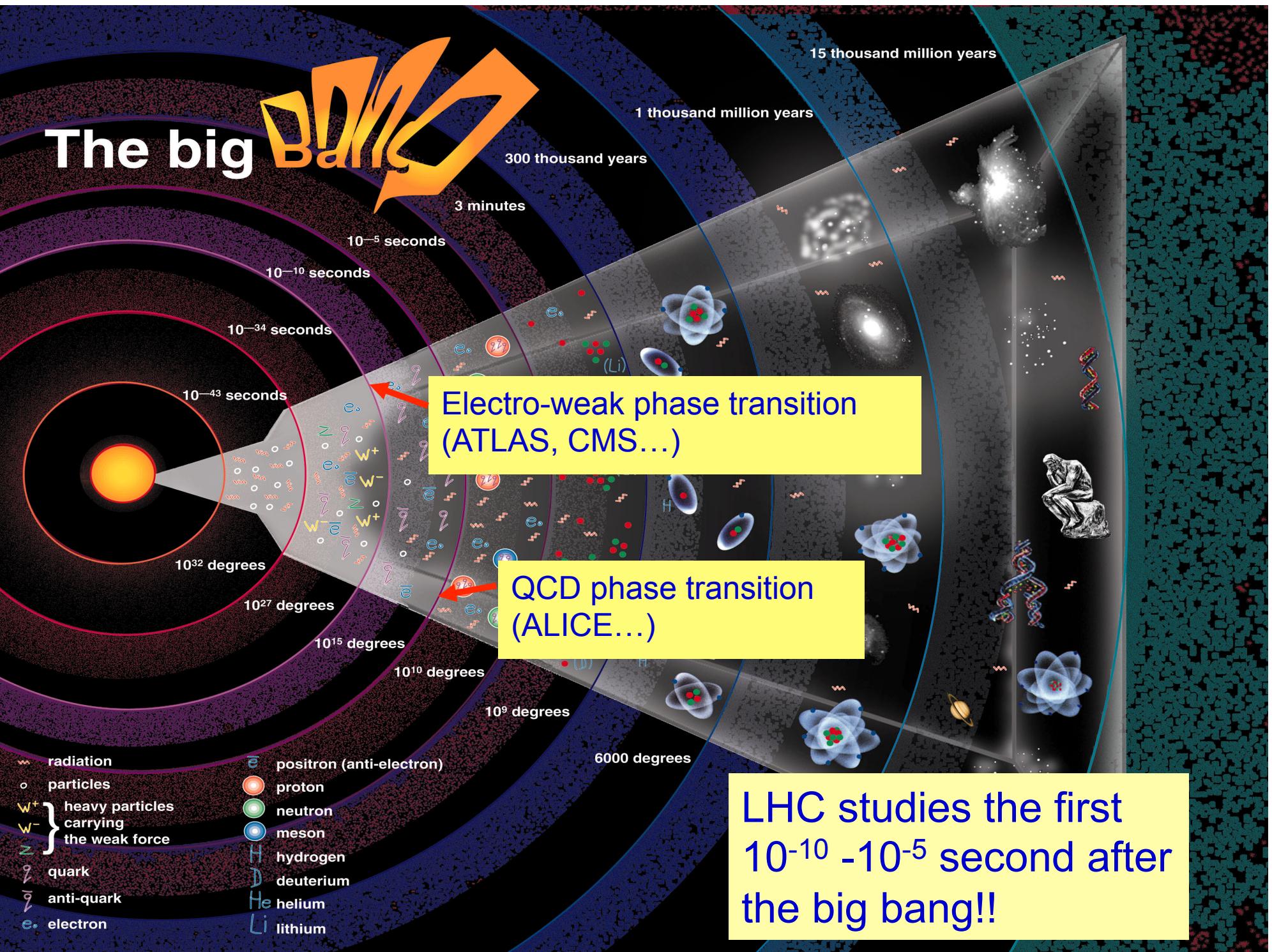


# The Large Hadron Collider at CERN

DM Searches @ collider & Direct Detection O. Buchmüller



# The big Bang

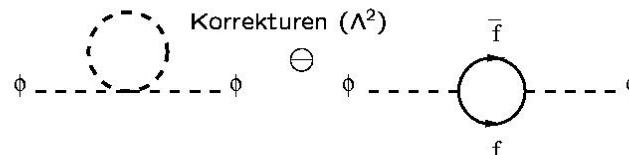


# In Supersymmetry at colliders

(i.e. a famous model...but only  
one model)

# Why is SUSY so attractive?

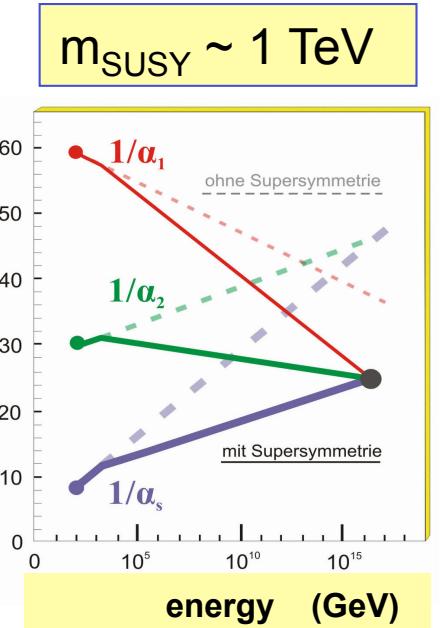
1. Quadratically divergent quantum corrections to the Higgs boson mass are avoided



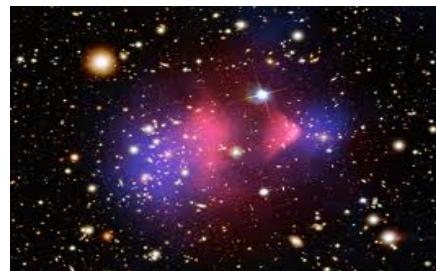
$$\Delta m_H = f(m_B^2 - m_f^2)$$

(Hierarchy or naturalness problem)

2. Unification of coupling constants of the three interactions seems possible

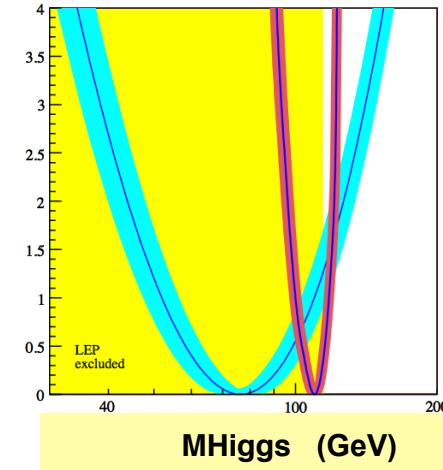


3. SUSY provides a candidate for dark matter,



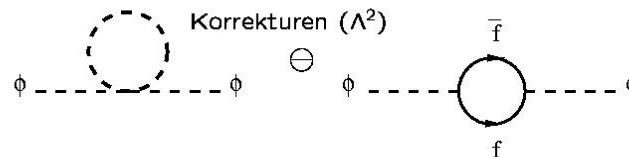
The lightest  
SUSY particle  
(LSP)

4. A SUSY extension is a small perturbation, consistent with the electroweak precision data



# Why is SUSY so attractive?

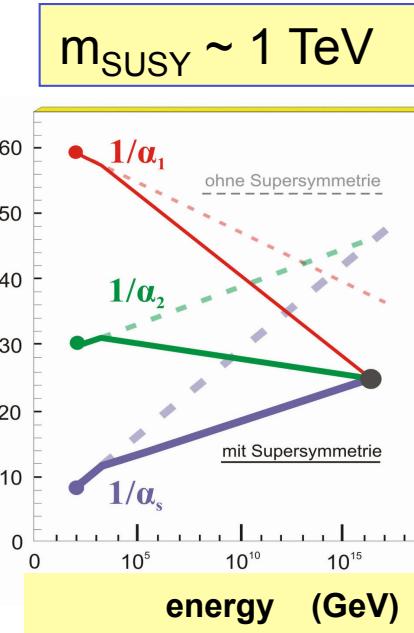
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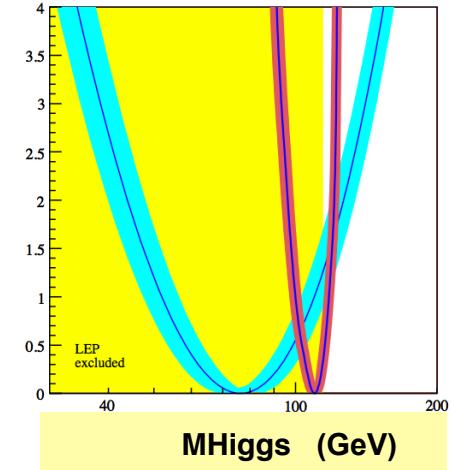


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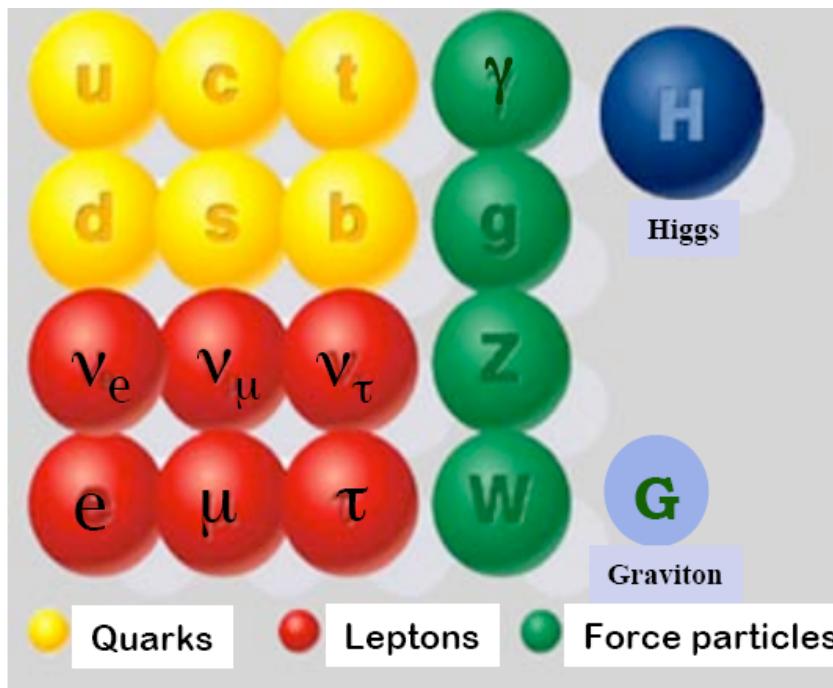
4. A SUSY extension is a small perturbation, consistent with the electroweak precision data



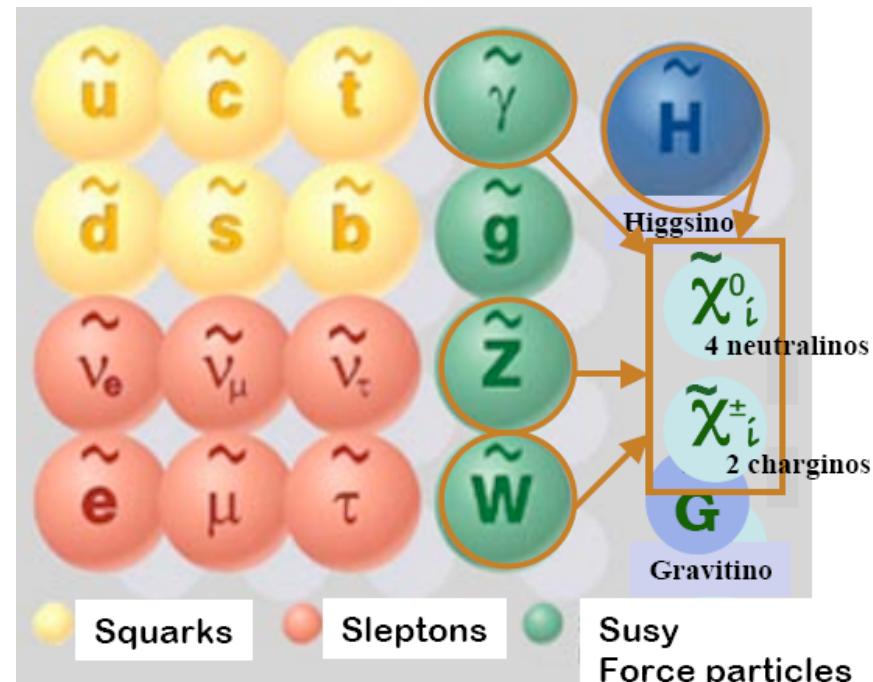
# Supersymmetry

Extension of the Standard Model: Introduce a new symmetry  
 Spin  $\frac{1}{2}$  matter particles (fermions)  $\Leftrightarrow$  Spin 1 force carriers (bosons)

## Standard Model particles



## SUSY particles



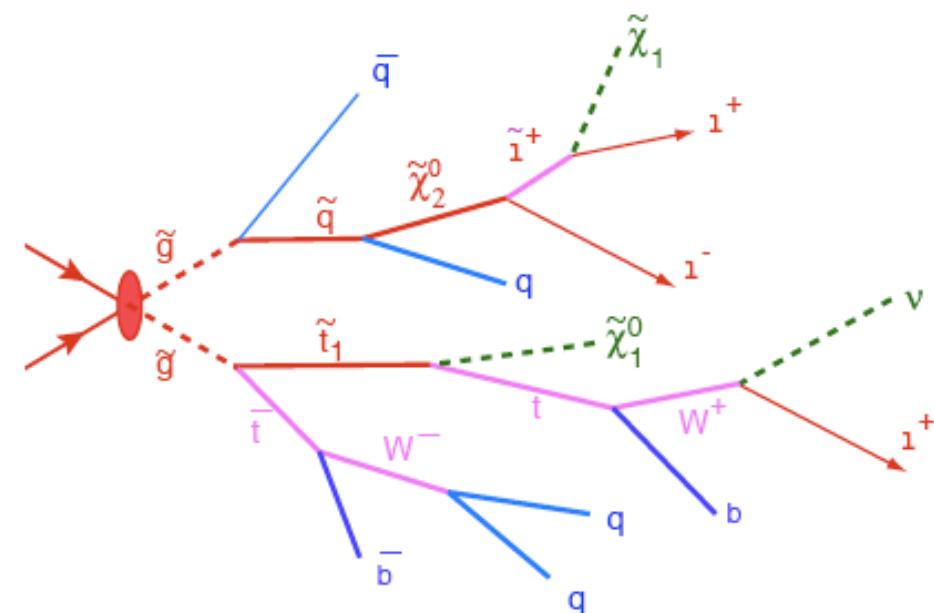
New Quantum number: R-parity:

R-parity conservation:

- SUSY particles are produced in pairs
- The lightest SUSY particle (LSP) is stable

$$R_p = (-1)^{B+L+2s} = \begin{cases} +1 & \text{SM particles} \\ -1 & \text{SUSY particles} \end{cases}$$

# What do we call a “SUSY search”?



RP-Conserving SUSY is a very prominent example predicting this famous signature but ...

## Missing Energy:

- from LSP

## Multi-Jet:

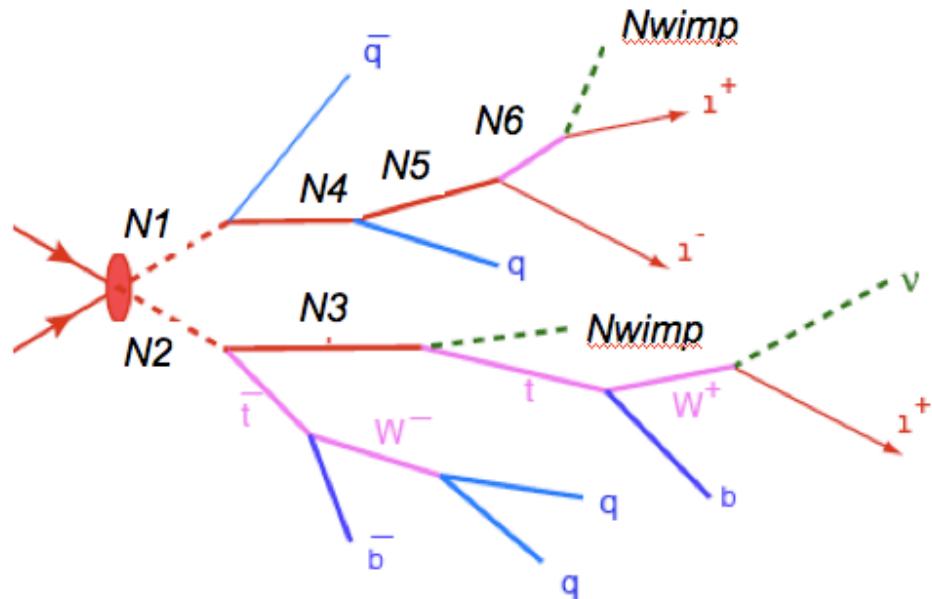
- from cascade decay (gaugino)

## Multi-Leptons:

- from decay of charginos/neutralinos

# What is its experimental signature?

... by no means is it the only New Physics model predicting this experimental pattern. Many other NP models predict this genuine signature



### Missing Energy:

- Nwimp - end of the cascade

### Multi-Jet:

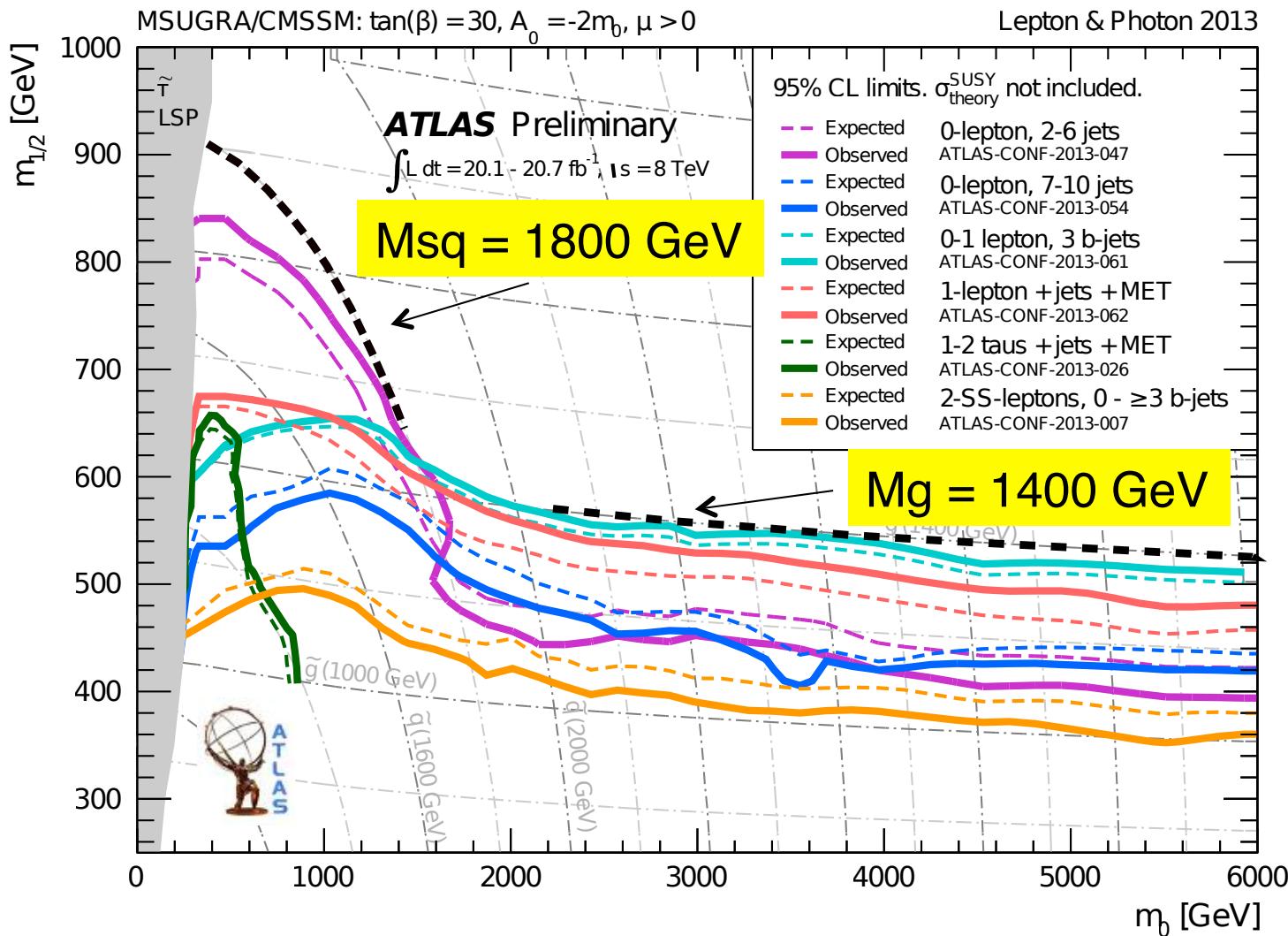
- from decay of the Ns (possibly via heavy SM particles like top, W/Z)

### Multi-Leptons:

- from decay of the N's

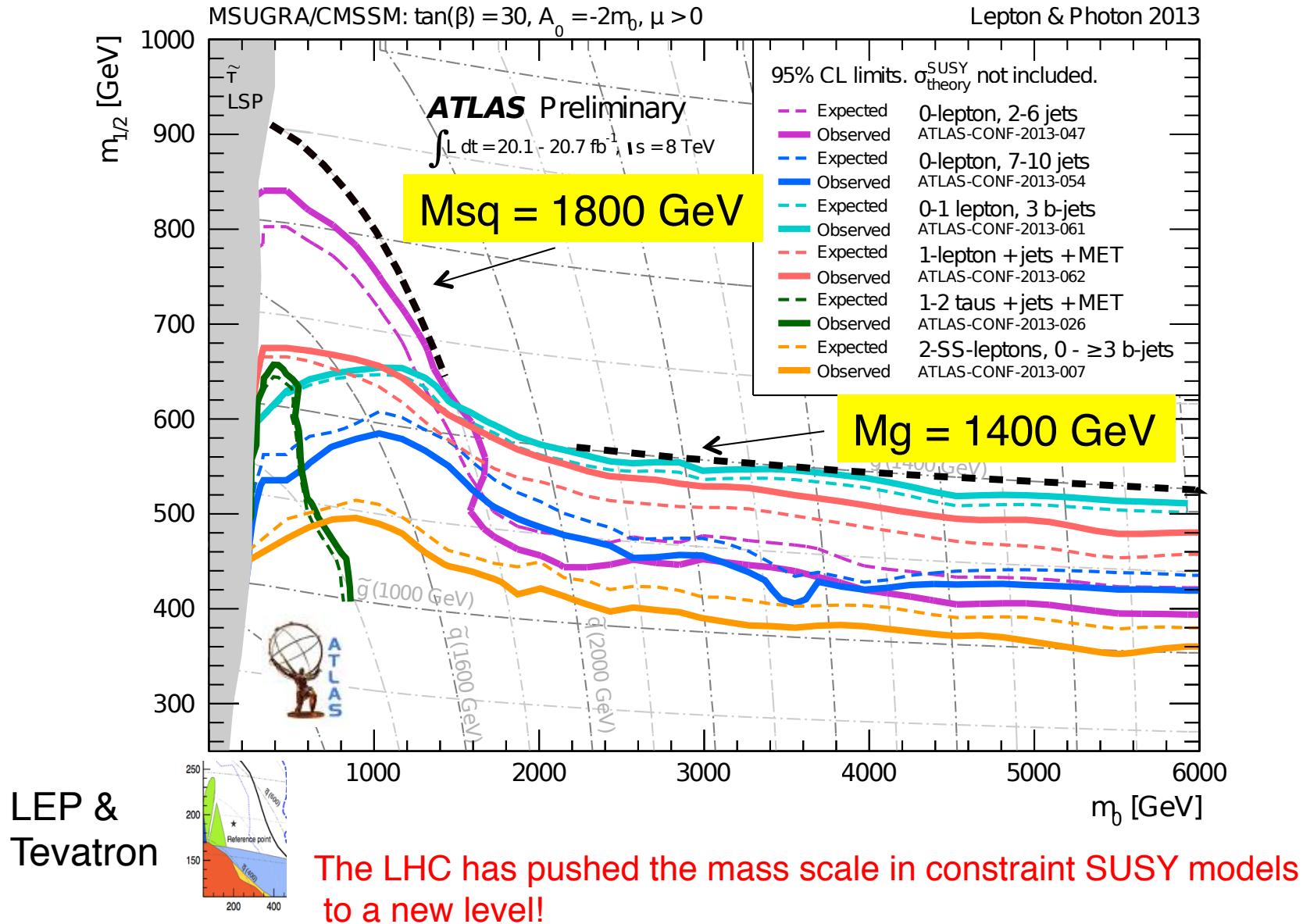
Model examples are Extra dimensions, Little Higgs, Technicolour, etc  
but a more generic definition for this signature is as follows.

# Inclusive SUSY Searches in 2013

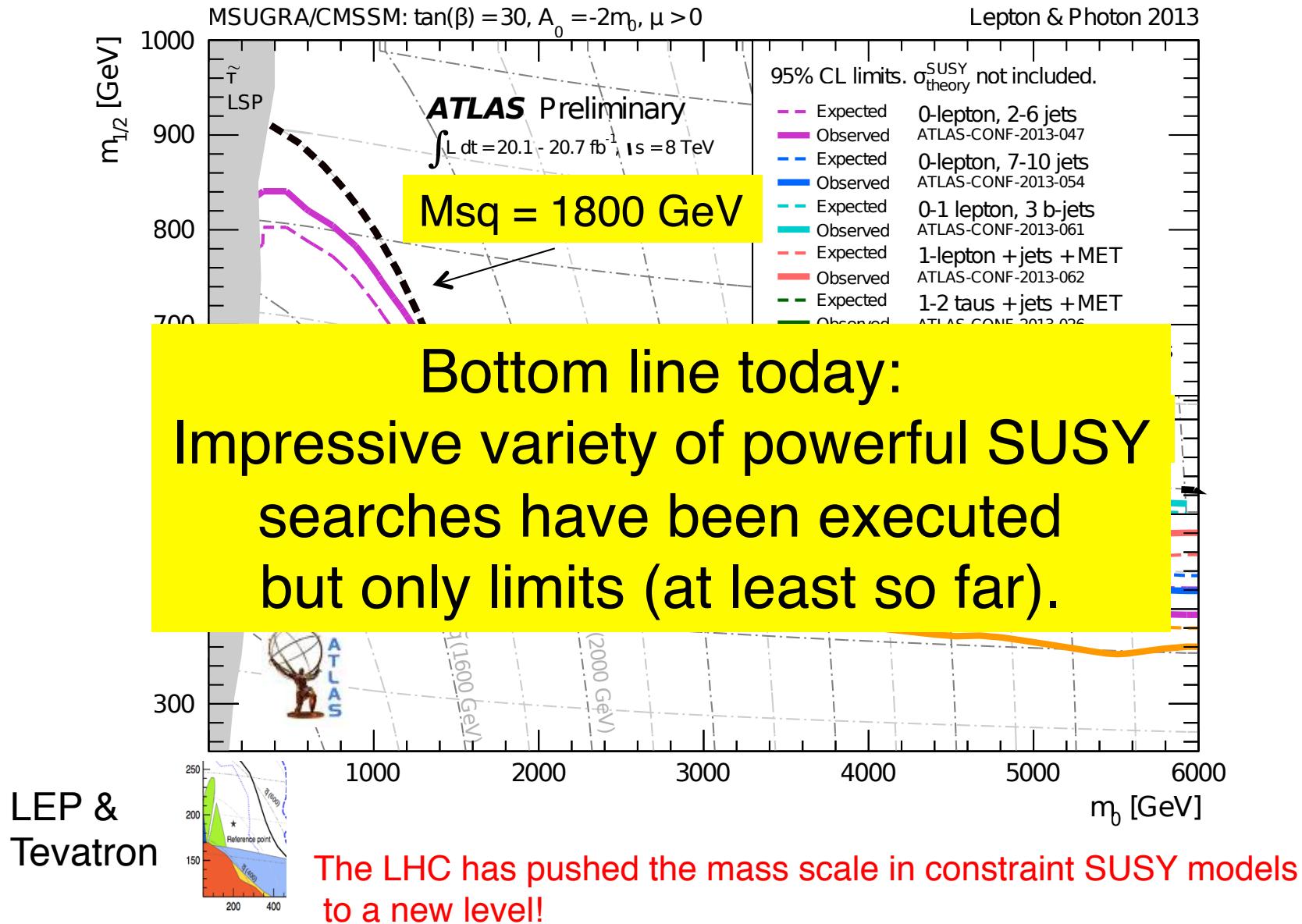


# Inclusive SUSY Searches in 2013

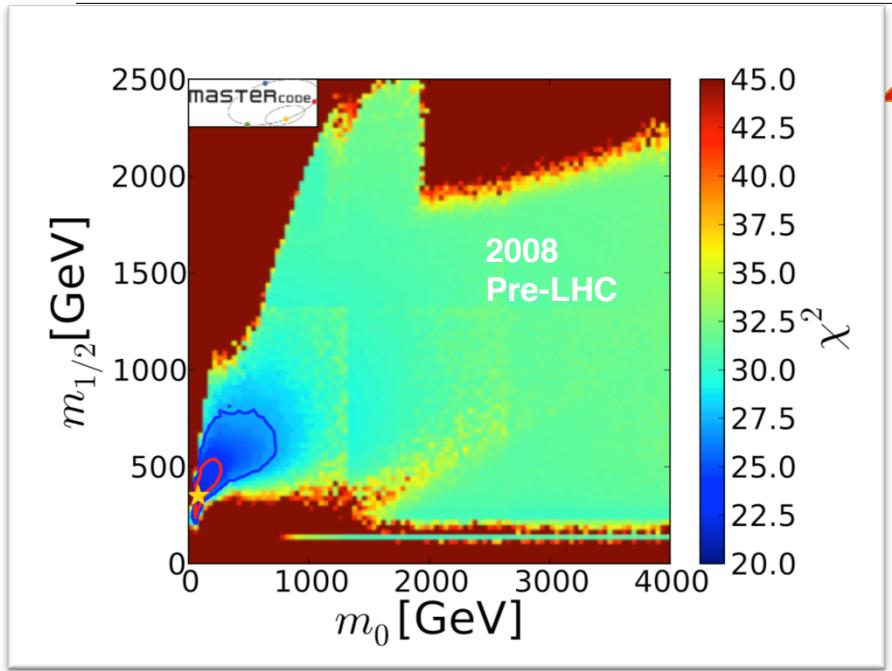
DM Searches @ collider & Direct Detection O. Buchmüller



# Inclusive SUSY Searches in 2013



# CMSSM: Evolution with time



$\chi^2$  increase from bluish to reddish



Source:  
<http://mastercode.web.cern.ch/mastercode/>

Observable	Source Th./Ex.	Constraint	$\Delta\chi^2$ (CMSSM)	$\Delta\chi^2$ (NUHM1)	$\Delta\chi^2$ ("SM")
$m_t$ [GeV]	[43]	$173.2 \pm 0.90$	0.05	0.06	-
$\Delta\alpha_{had}^{(b)}(M_Z)$	[42]	$0.02749 \pm 0.00010$	0.009	0.004	-
$M_Z$ [GeV]	[44]	$91.1875 \pm 0.0021$	$2.7 \times 10^{-6}$	0.26	-
$\Gamma_Z$ [GeV]	[26] / [44]	$2.4952 \pm 0.0023 \pm 0.001_{SUSY}$	0.078	0.047	0.14
$\sigma_{had}^0$ [nb]	[26] / [44]	$41.540 \pm 0.037$	2.50	2.57	2.54
$R_l$	[26] / [44]	$20.767 \pm 0.025$	1.05	1.08	1.08
$A_{fb}(\ell)$	[26] / [44]	$0.01714 \pm 0.00095$	0.72	0.69	0.81
$A_t(P_T)$	[26] / [44]	$0.1465 \pm 0.0032$	0.11	0.13	0.07
$R_b$	[26] / [44]	$0.21629 \pm 0.00066$	0.26	0.29	0.27
$R_c$	[26] / [44]	$0.1721 \pm 0.0030$	0.002	0.002	0.002
$A_{fb}(b)$	[26] / [44]	$0.0992 \pm 0.0016$	7.17	7.37	6.63
$A_{fb}(c)$	[26] / [44]	$0.0707 \pm 0.0035$	0.86	0.88	0.80
$A_b$	[26] / [44]	$0.923 \pm 0.020$	0.36	0.36	0.35
$A_c$	[26] / [44]	$0.670 \pm 0.027$	0.005	0.005	0.005
$A_t(SLD)$	[26] / [44]	$0.1513 \pm 0.0021$	3.16	3.03	3.51
$\sin\theta_W(Q_b)$	[26] / [44]	$0.2324 \pm 0.0012$	0.63	0.64	0.59
$M_W$ [GeV]	[26] / [44]	$80.399 \pm 0.023 \pm 0.010_{SUSY}$	1.77	1.39	2.08
$a_\mu^{EXP} - a_\mu^{SM}$	[53] / [42, 54]	$(30.2 \pm 8.8 \pm 2.0_{SUSY}) \times 10^{-10}$	4.35	1.82	11.19 (N/A)
$M_h$ [GeV]	[28] / [55, 56]	$> 114.4 \pm 1.5_{SUSY}$	0.0	0.0	0.0
$BR_{b \rightarrow s\gamma}^{EXPV/SM}$	[45] / [46]	$1.117 \pm 0.076_{EXP} \pm 0.082_{SM} \pm 0.050_{SUSY}$	1.83	1.09	0.94
$BR(B_s \rightarrow \mu^+ \mu^-)$	[29] / [41]	CMS & LHCb	0.04	0.44	0.01
$BR_{s \rightarrow \tau \nu}^{EXPV/SM}$	[29] / [46]	$1.43 \pm 0.43_{EXP+TH}$	1.43	1.59	1.00
$BR(B_d \rightarrow \mu^+ \mu^-)$	[29] / [46]	$< 4.6 \pm 0.01_{SUSY} \times 10^{-6}$	0.0	0.0	0.0
$BR_{d \rightarrow X_s \ell \bar{\nu}}^{EXPV/SM}$	[47] / [46]	$0.99 \pm 0.32$	0.02	$\ll 0.01$	$\ll 0.01$
$BR_{X_s \rightarrow \ell \bar{\nu}}^{EXPV/SM}$	[29] / [48]	$1.008 \pm 0.014_{EXP+TH}$	0.39	0.42	0.33
$BR_{X_s \rightarrow \tau \nu}^{EXPV/SM}$	[49] / [50]	$< 4.5$	0.0	0.0	0.0
$\Delta M_{B_s}^{EXPV/SM}$	[49] / [51, 52]	$0.97 \pm 0.01_{EXP} \pm 0.27_{SM}$	0.02	0.02	0.01
$\Delta M_{B_d}^{EXPV/SM}$	[29] / [46, 51, 52]	$1.00 \pm 0.01_{EXP} \pm 0.13_{SM}$	$\ll 0.01$	0.33	$\ll 0.01$
$\Delta t_K^{EXPV/SM}$	[49] / [51, 52]	$1.08 \pm 0.14_{EXP+TH}$	0.27	0.37	0.33
$\Omega_{CDM} h^2$	[31] / [13]	$0.1120 \pm 0.0056 \pm 0.012_{SUSY}$	$8.4 \times 10^{-4}$	0.1	N/A
$\sigma_p^{in}$	[25]	$(m_{1/2}, \sigma_p^{in})$ plane	0.13	0.13	N/A
jets + $E_T$	[18, 20]	$(m_0, m_{1/2})$ plane	1.55	2.20	N/A
$H/A, H^\pm$	[21]	$(M_A, \tan \beta)$ plane	0.0	0.0	N/A
Total $\chi^2/\text{d.o.f.}$	All	All	28.8/22	27.3/21	32.7/23 (21.5/22)
p-values			15%	16%	9% (49%)

**Global Fit to indirect and direct constraints on SUSY!**

Other “fitter” groups find very similar results: e.g.

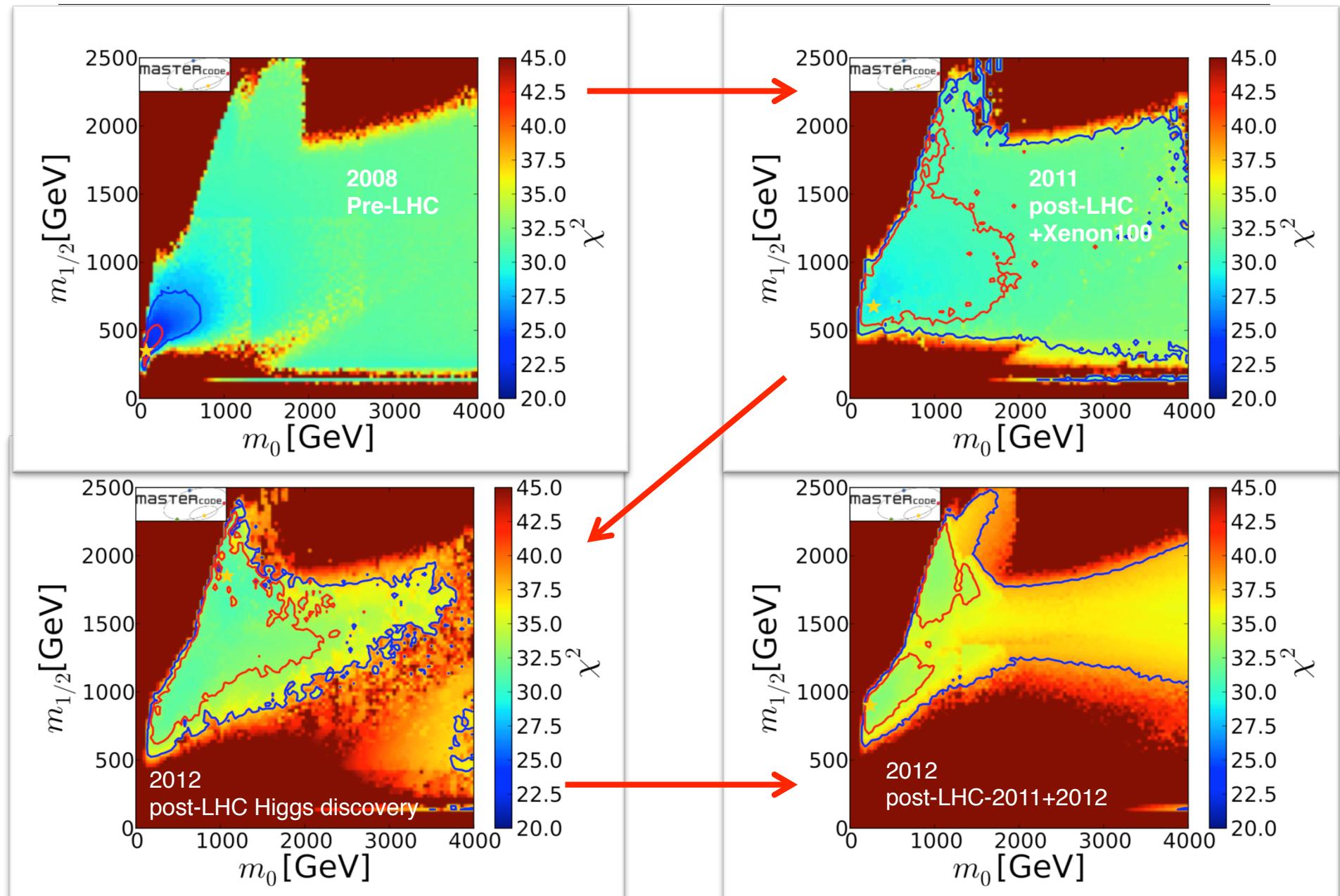
SuperBayeS: [arXiv:1212.2636](https://arxiv.org/abs/1212.2636)

Fittino group: [arXiv:1204.4199](https://arxiv.org/abs/1204.4199)

# CMSSM: Evolution with time



DM Searches @ collider & Direct Detection O. Buchmüller

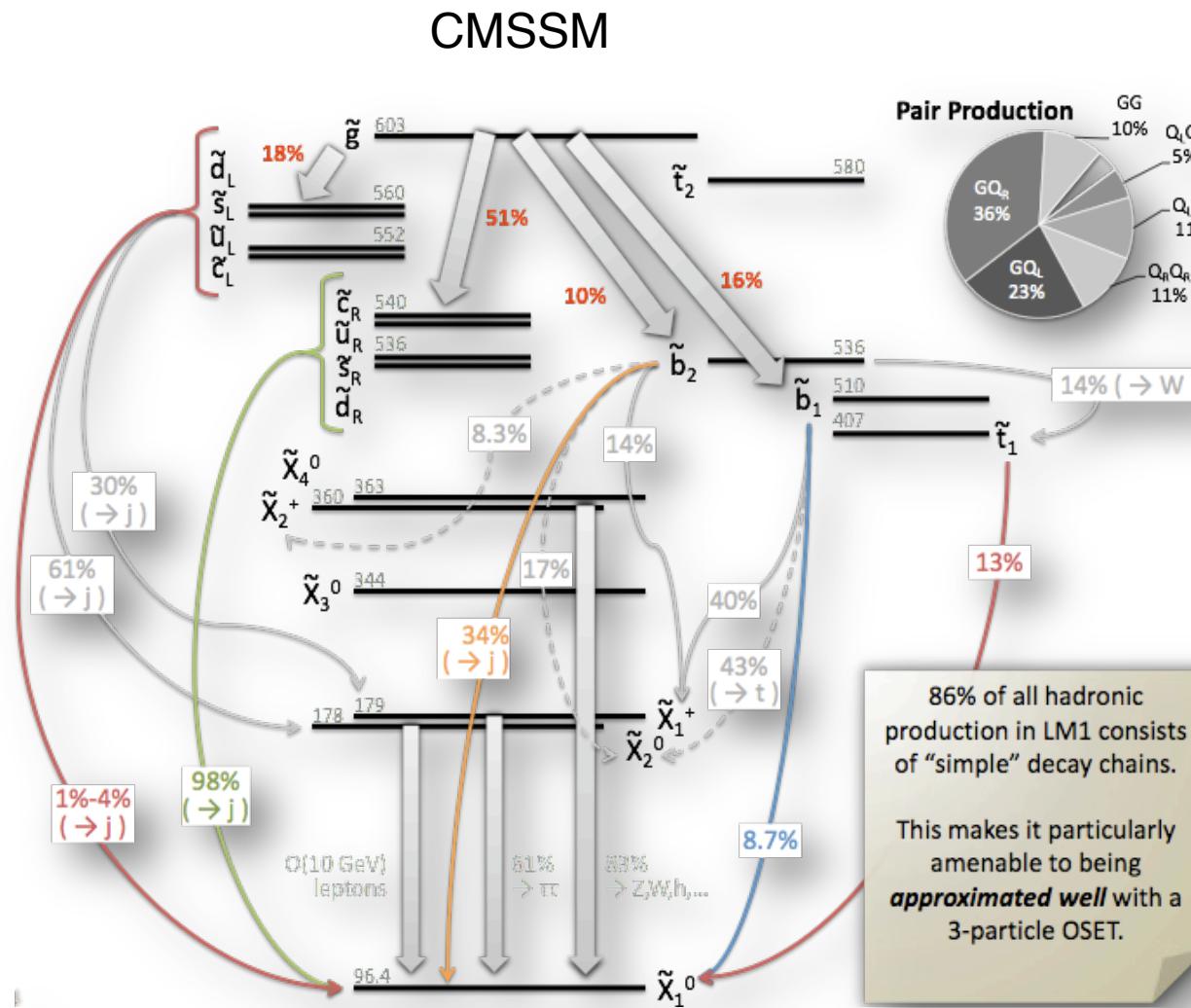


## SUSY Status – post 7 TeV LHC data

- Constrained SUSY models like the CMSSM are severely put under pressure by the LHC limits!
- Experiments define new benchmarks and less complex SUSY models in order to present the interpretation of their searches.
- Aided by the discovery of a Higgs boson, the focus of the experimental search strategy and corresponding interpretation shifts towards other scenarios like “Natural SUSY” (i.e. 3<sup>rd</sup> generation squark searches).

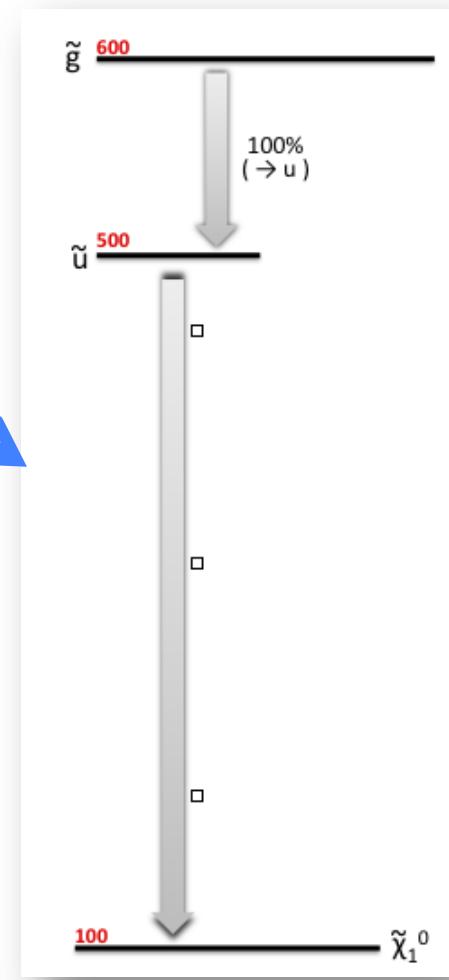
# Interpretation in Simplified Models

DM Searches @ collider & Direct Detection O. Buchmüller



**Simplified model spectrum (SMS)**  
with 3 particles, 2 decay modes

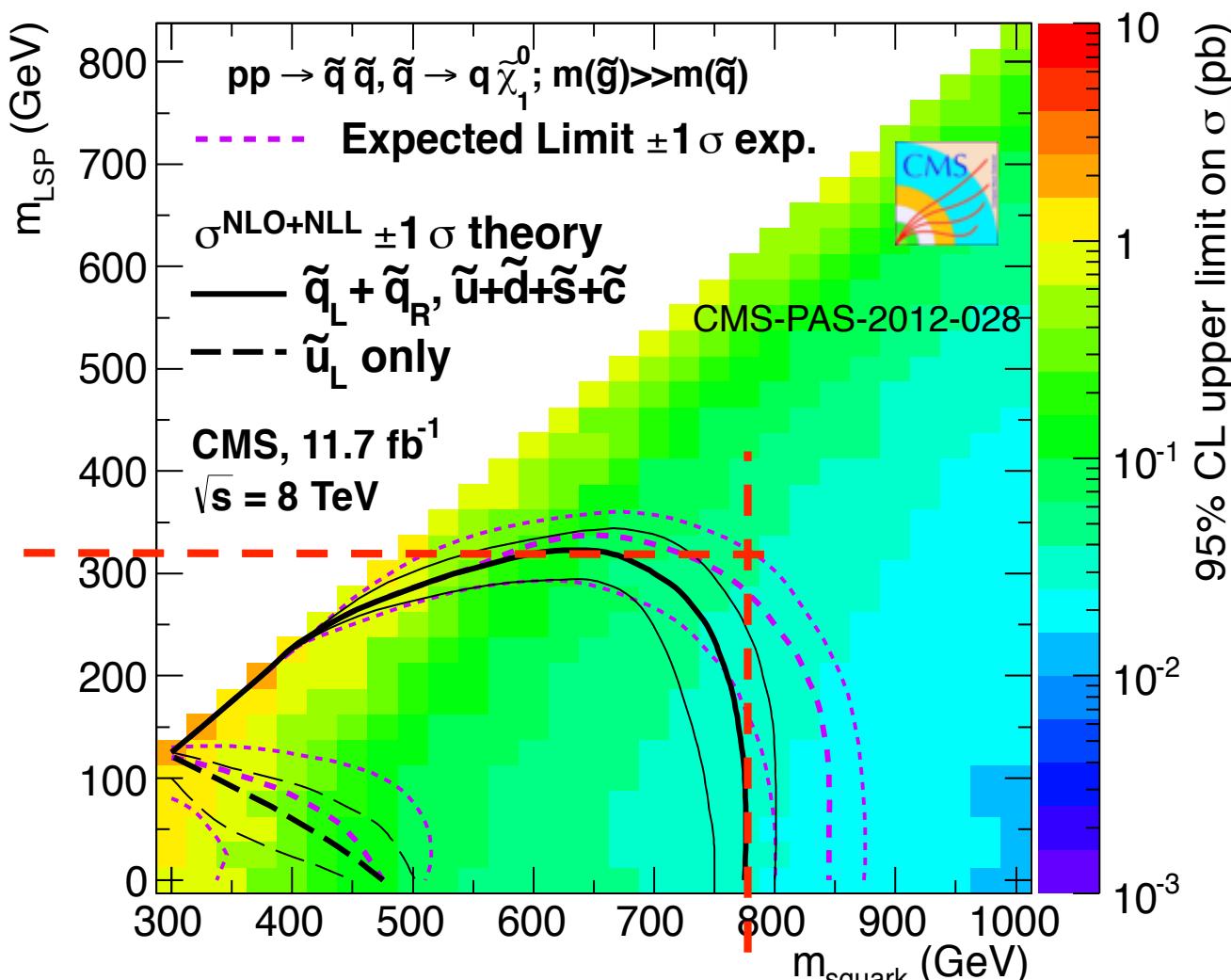
What the individual searches are sensitive to is much more simple...



## SMS: a few interesting features

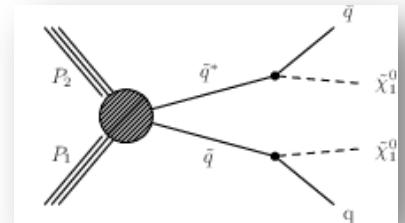
DM Searches @ collider & Direct Detection O. Buchmüller

$m_{\text{LSP}}^{\text{max}} \approx 0.3 \text{ TeV}$  : LSP mass above which there is NO limit anymore



$m_G^{\text{max}} \approx 0.8 \text{ TeV}$  : Best limit in plane

Assumes 100%  
BR for decay chain  
considered.



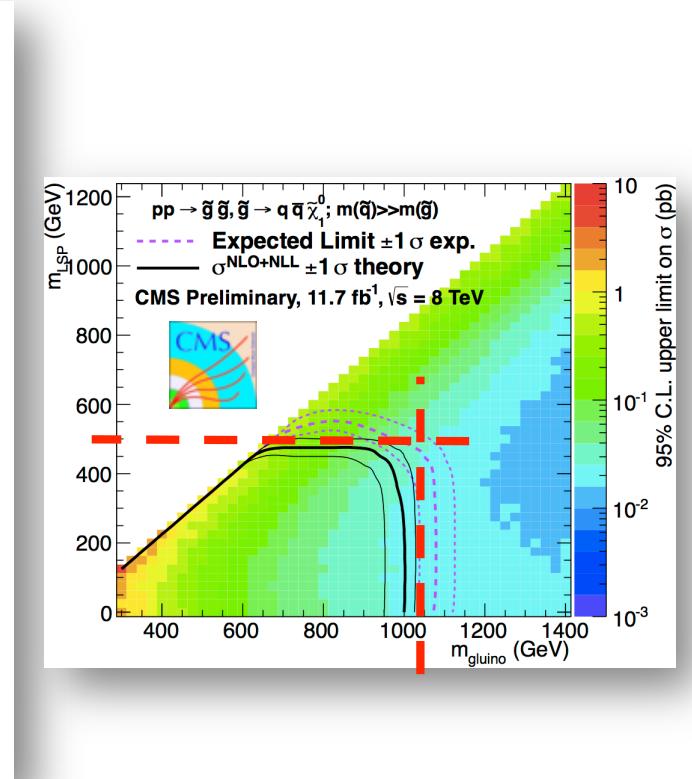
$$\tilde{q}\tilde{q} \rightarrow q\tilde{\chi}_1^0\bar{q}\tilde{\chi}_1^0$$

# How to summarize SMS limits?

*Approach taken in the 2012 and 2013 Experimental SUSY PDG reviews  
[OB & Paul De Jong]:*

<http://pdg.lbl.gov/2012/reviews/rpp2012-rev-susy-2-experiment.pdf>  
<http://pdg.lbl.gov/2013/reviews/rpp2013-rev-susy-2-experiment.pdf>

Model	Assumption	$m_{\tilde{q}}$	$m_{\tilde{g}}$
CMSSM	$m_{\tilde{q}} \approx m_{\tilde{g}}$	1400	1400
	all $m_{\tilde{q}}$	-	800
	all $m_{\tilde{g}}$	1300	-
Simplified model $\tilde{g}\tilde{g}$	$m_{\tilde{\chi}_1^0} = 0$	-	900
	$m_{\tilde{\chi}_1^0} > 300$	-	no limit
Simplified model $\tilde{q}\tilde{q}$	$m_{\tilde{\chi}_1^0} = 0$	750	-
	$m_{\tilde{\chi}_1^0} > 250$	no limit	-
Simplified model $\tilde{g}\tilde{q}, \tilde{g}\bar{\tilde{q}}$	$m_{\tilde{\chi}_1^0} = 0, m_{\tilde{q}} \approx m_{\tilde{g}}$	1500	1500
	$m_{\tilde{\chi}_1^0} = 0, \text{ all } m_{\tilde{g}}$	1400	-
	$m_{\tilde{\chi}_1^0} = 0, \text{ all } m_{\tilde{q}}$	-	900



This was an appropriate approach for the rather limited amount of inclusive searches and corresponding SMS interpretations available in 2011 (7 TeV).

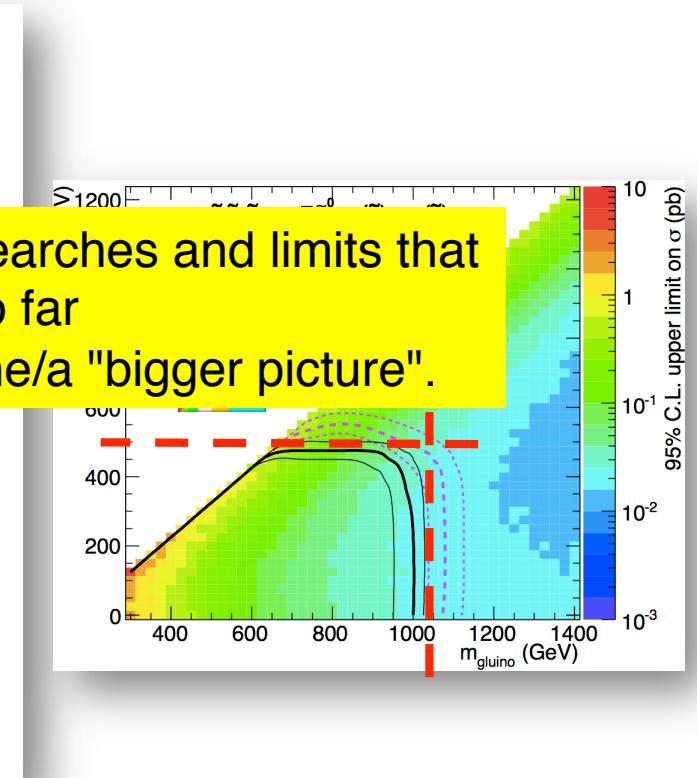
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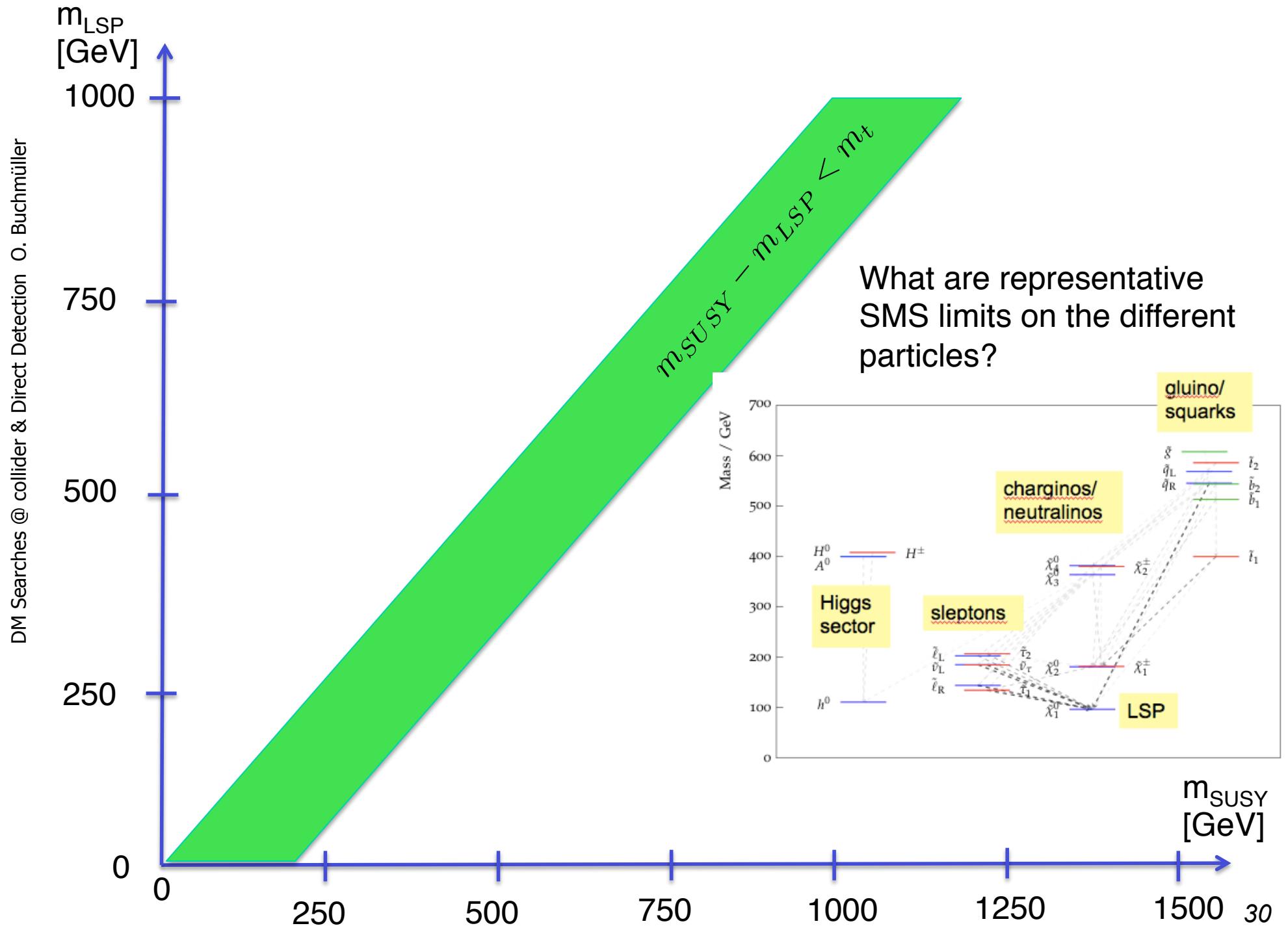
<http://pdg.lbl.gov/2012/reviews/rpp2012-rev-susy-2-experiment.pdf>  
<http://pdg.lbl.gov/2013/reviews/rpp2013-rev-susy-2-experiment.pdf>

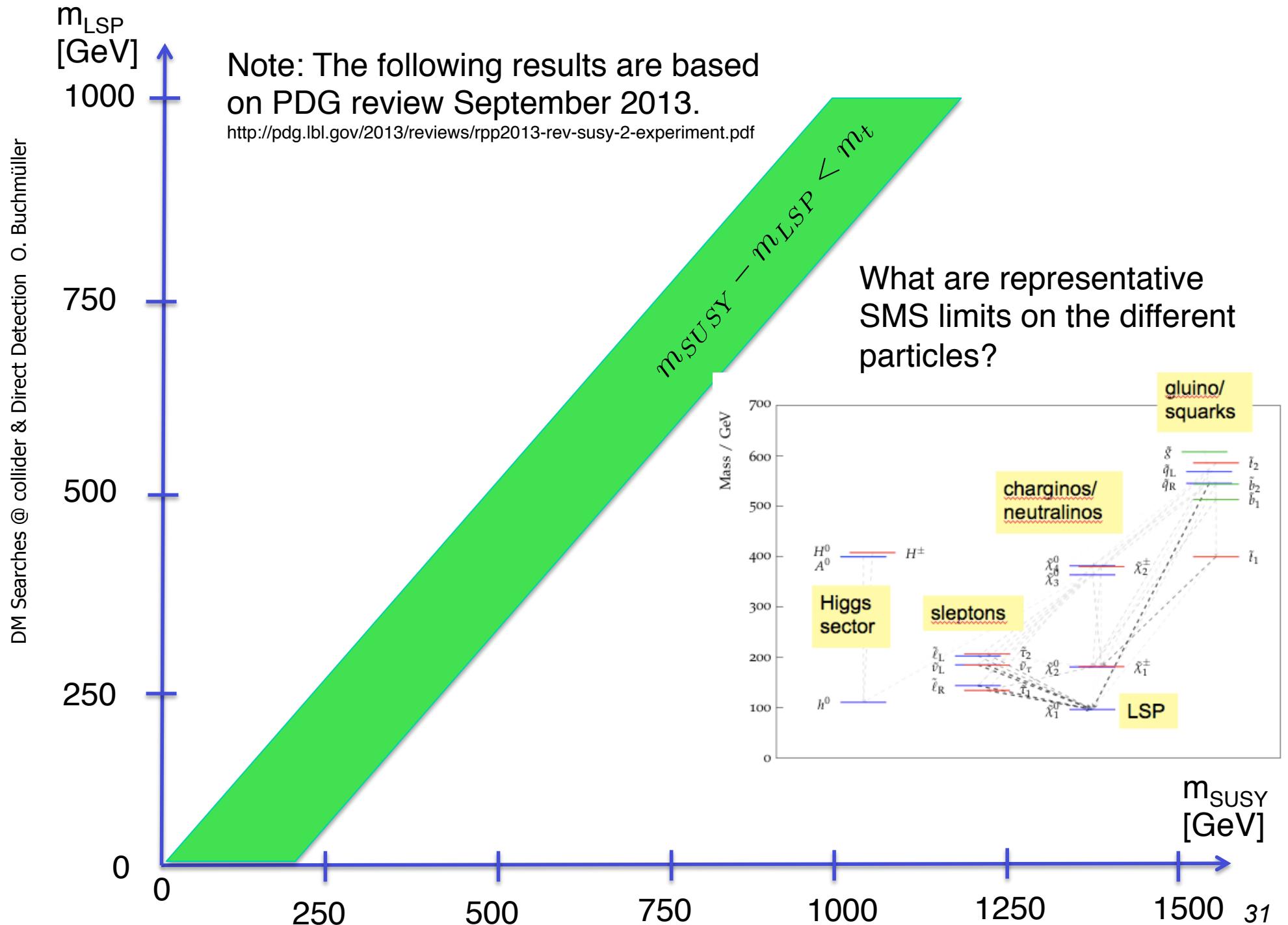
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CMSSM	$m_{\tilde{q}} \approx m_{\tilde{g}}$	1400	1400
	all $m_{\tilde{g}}$	-	800
Simplified	$m_{\tilde{\chi}_1^0} = 0$	750	-
	$m_{\tilde{\chi}_1^0} > 250$	no limit	-
Simplified model $\tilde{q}\tilde{q}$	$m_{\tilde{\chi}_1^0} = 0, m_{\tilde{q}} \approx m_{\tilde{g}}$	1500	1500
	$m_{\tilde{\chi}_1^0} = 0, \text{ all } m_{\tilde{g}}$	1400	-
	$m_{\tilde{\chi}_1^0} = 0, \text{ all } m_{\tilde{q}}$	-	900

It is a challenge to do justice to the many searches and limits that have been established so far  
- even more so to put it all together into the/a "bigger picture".

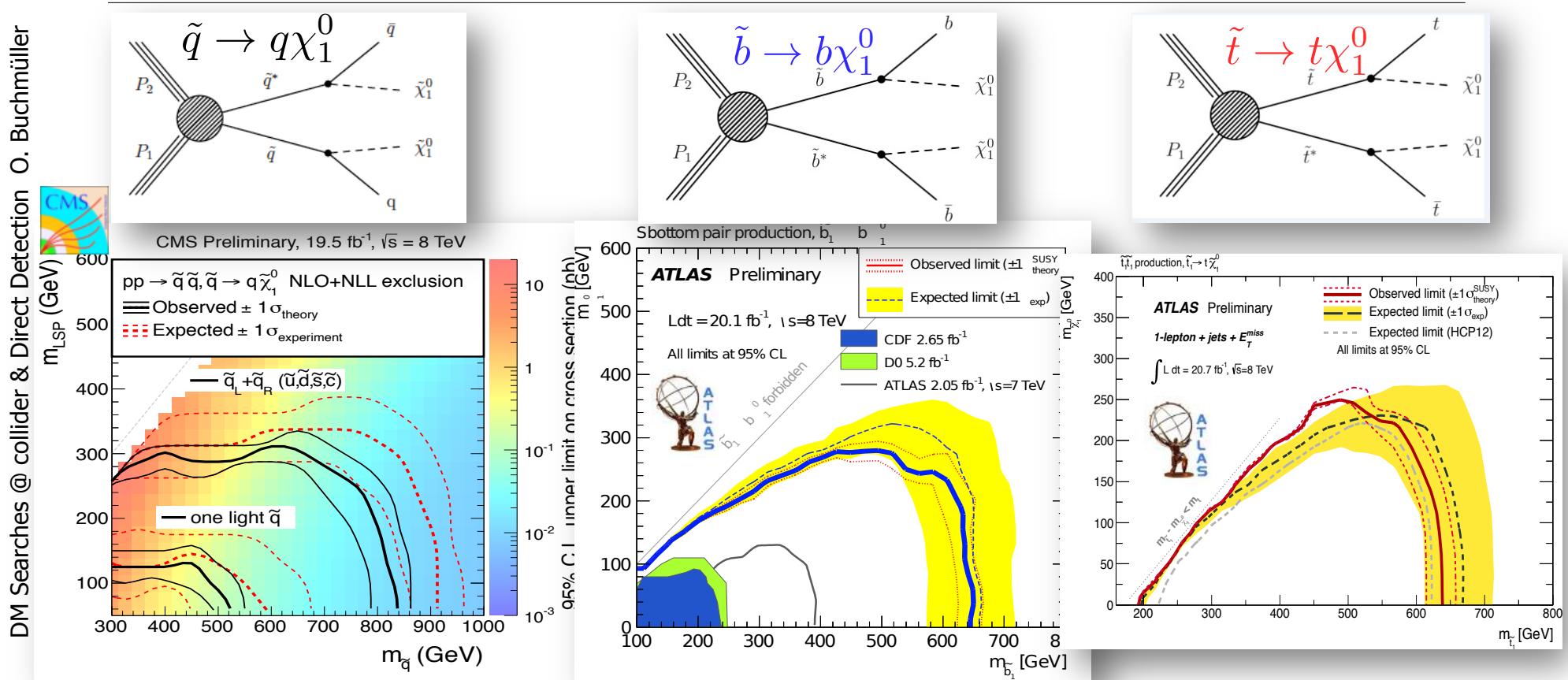


This was an appropriate approach for the rather limited amount of inclusive searches and corresponding SMS interpretations available in 2011 (7 TeV).





# Direct squark production – chosen limits



**CMS-SUS-PAS-13-012**

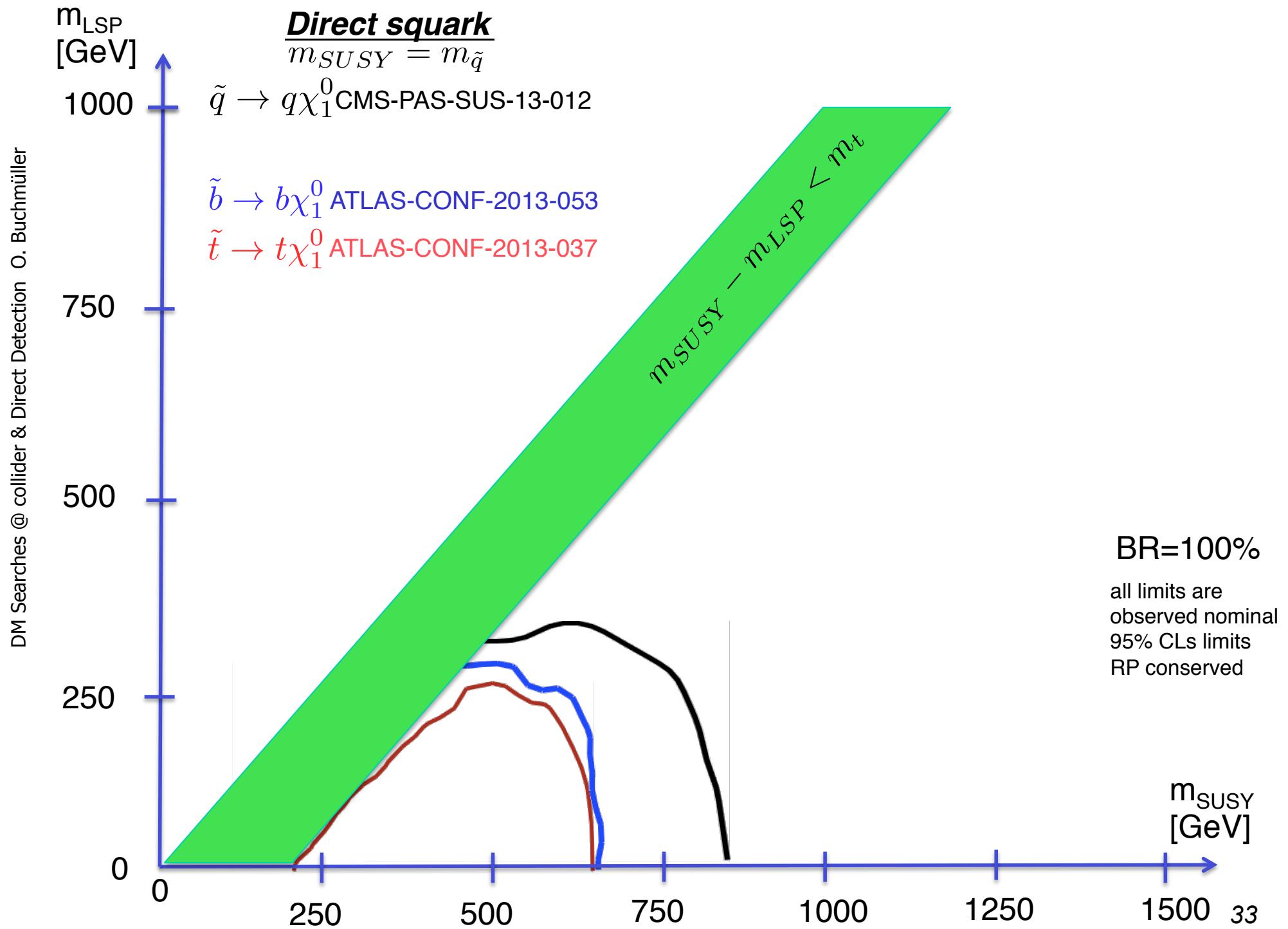
Signature: Jets +  $E_T^{\text{miss}} + H_T$   
Limit assumes all 1<sup>st</sup> & 2<sup>nd</sup> gen  
squarks to be mass degenerate  
or only one light squark!

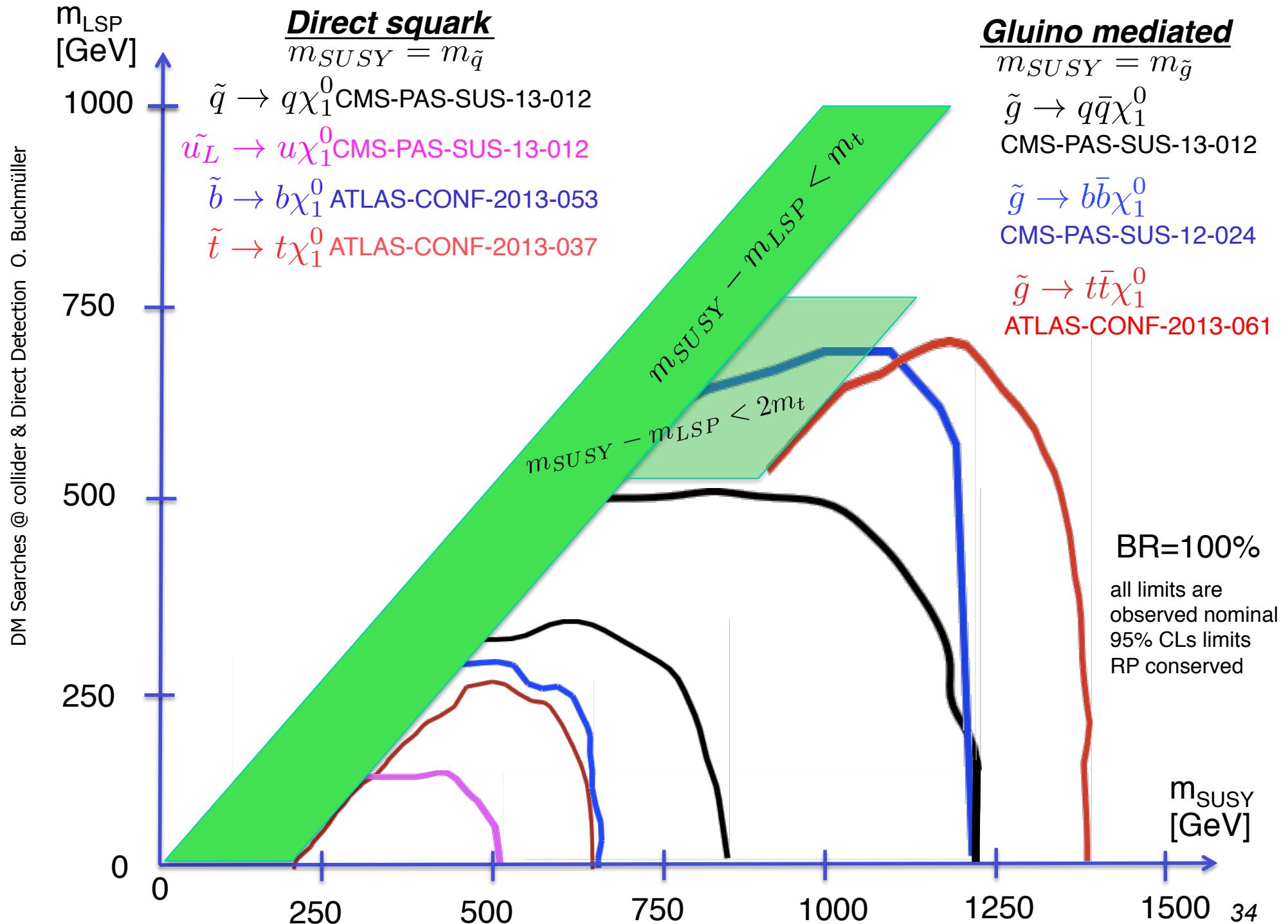
**ATLAS-CONF-2013-053**

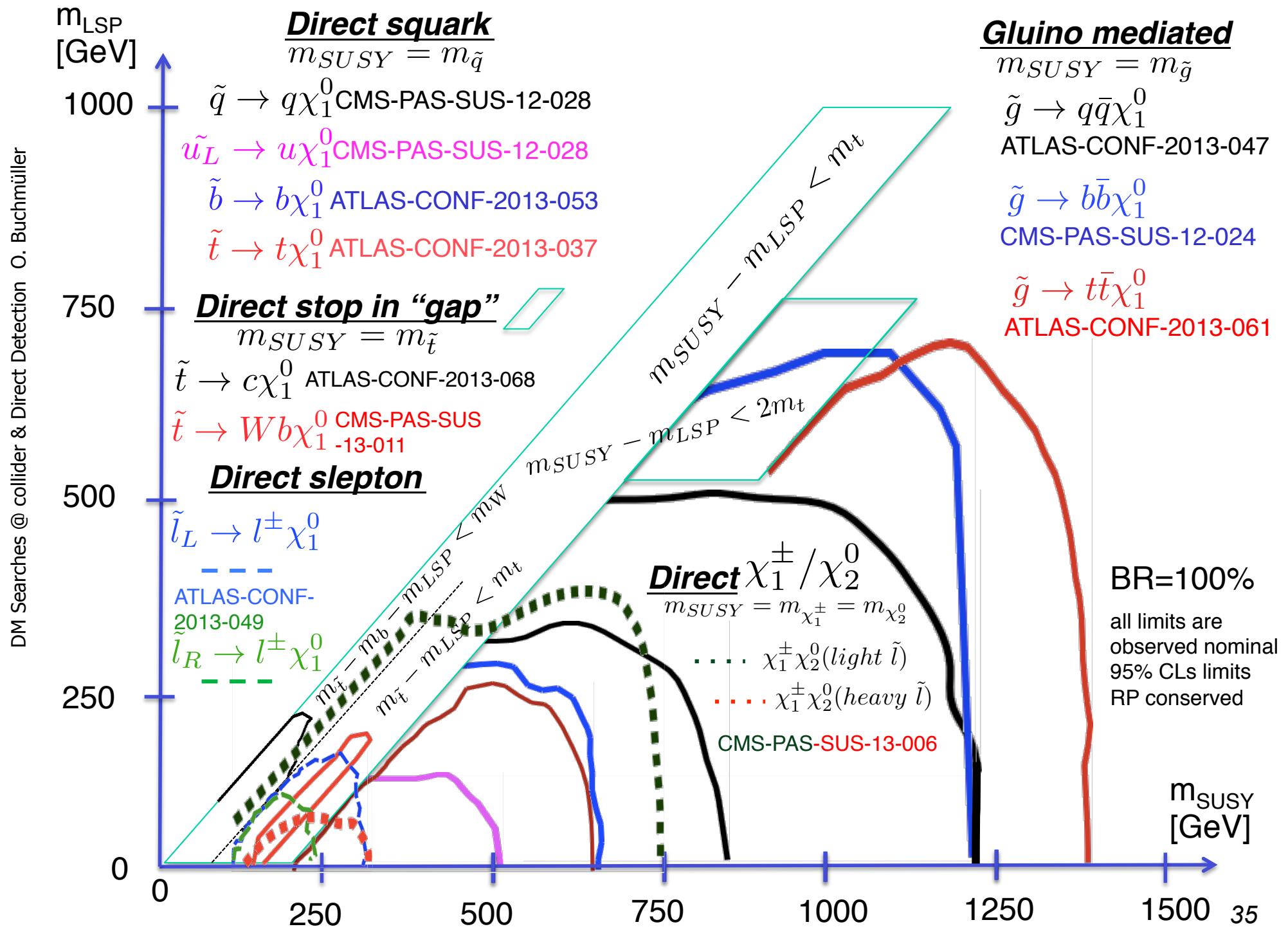
Signature: 2 b-jets +  $E_T^{\text{miss}}$

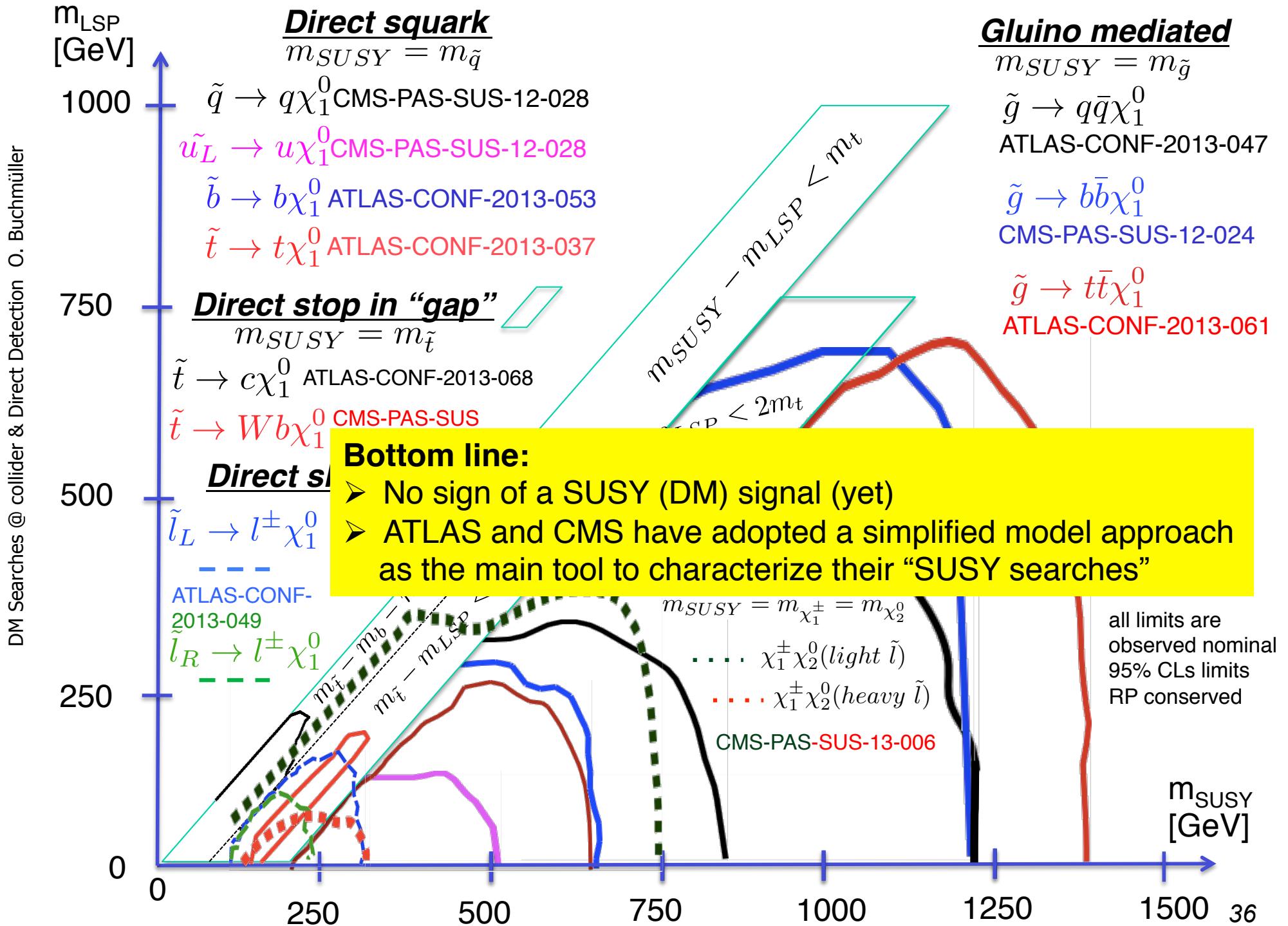
**ATLAS-CONF-2013-037**

Signature: 1Lepton + jets +  
 $E_T^{\text{miss}}$







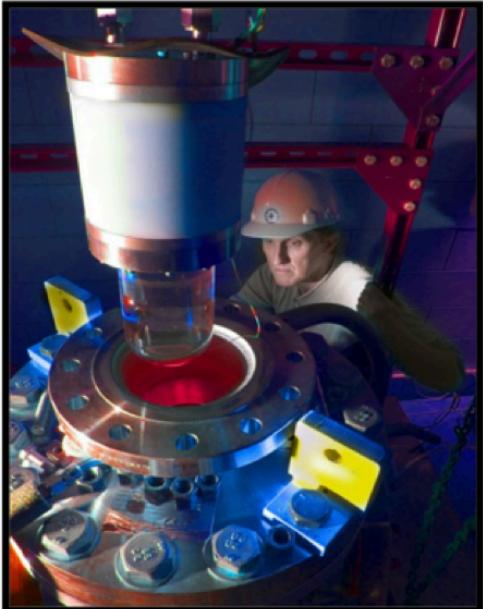


# Dark Matter Searches at Direct Detection Experiments

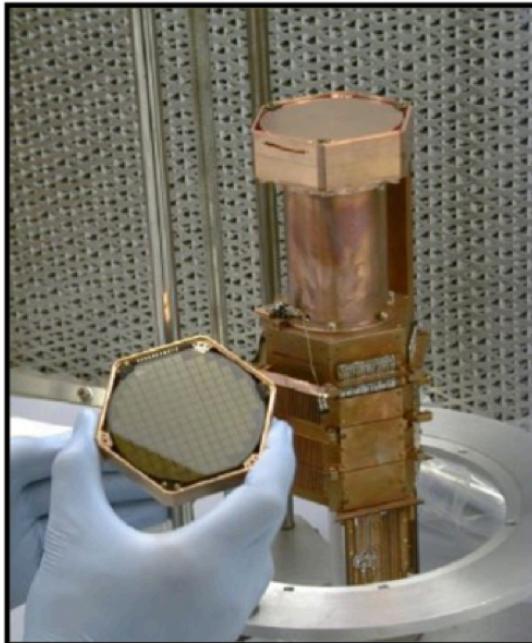


## Direct Detection Experiments: Examples

COUPP



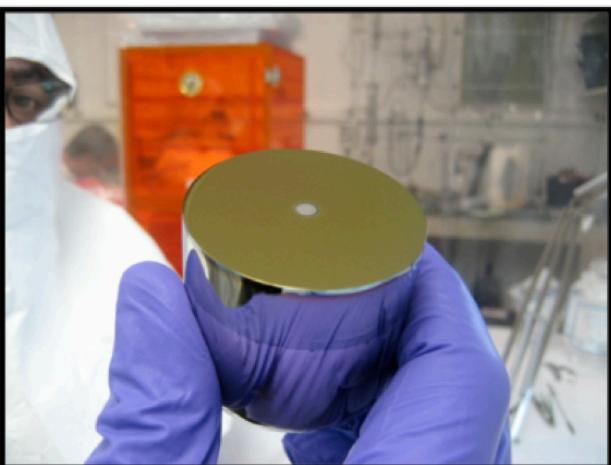
CDMS



CRESST



CoGeNT



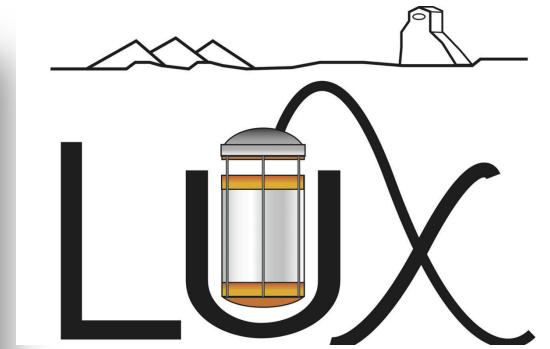
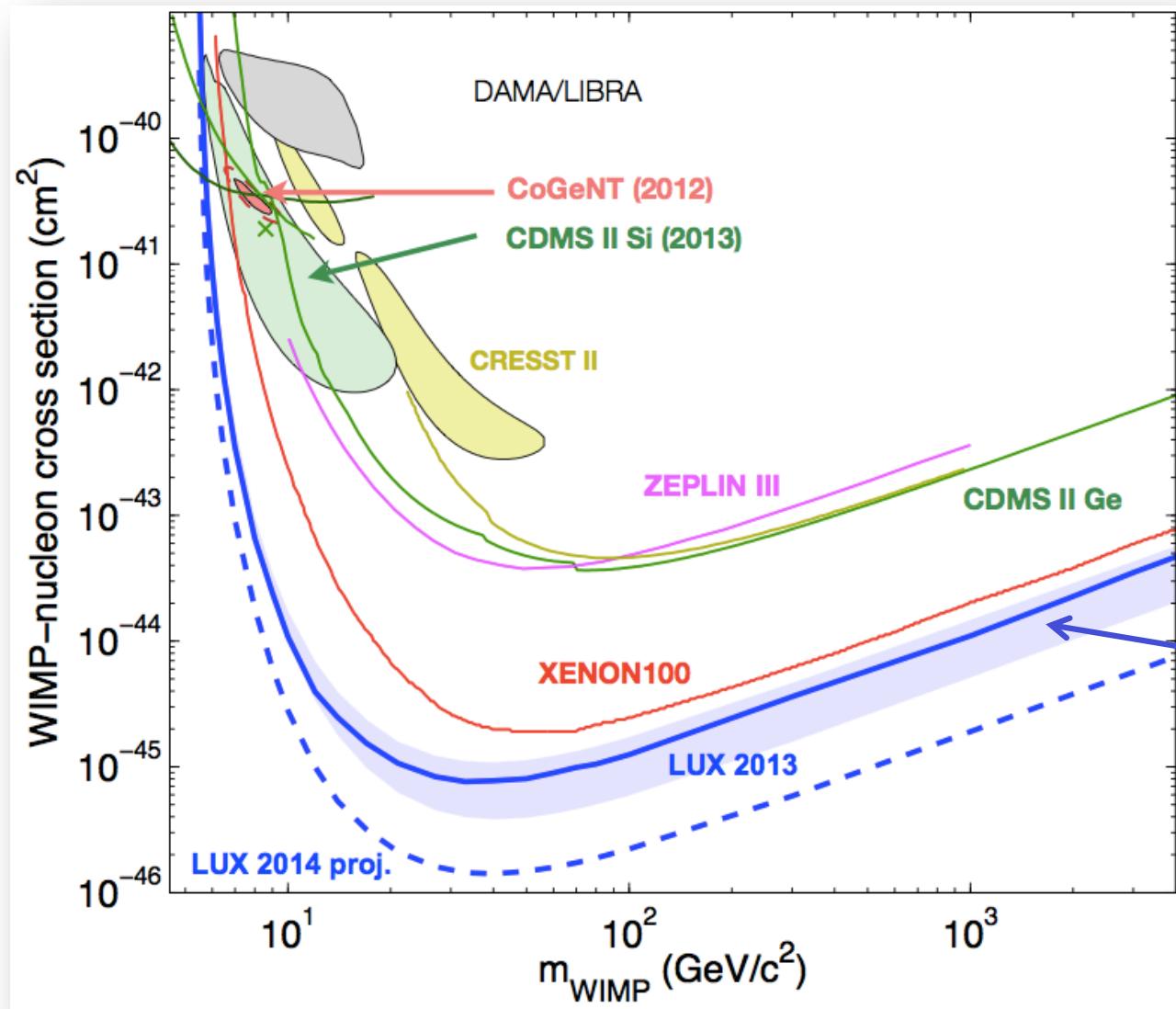
( + EDELWEISS,  
DAMA, EURECA,  
ZEPLIN, DEAP, ArDM,  
WARP, LUX, SIMPLE,  
PICASSO, DMTPC,  
DRIFT, KIMS, ...)

Xenon



# Dark Matter Searches: Direct Detection Experiments

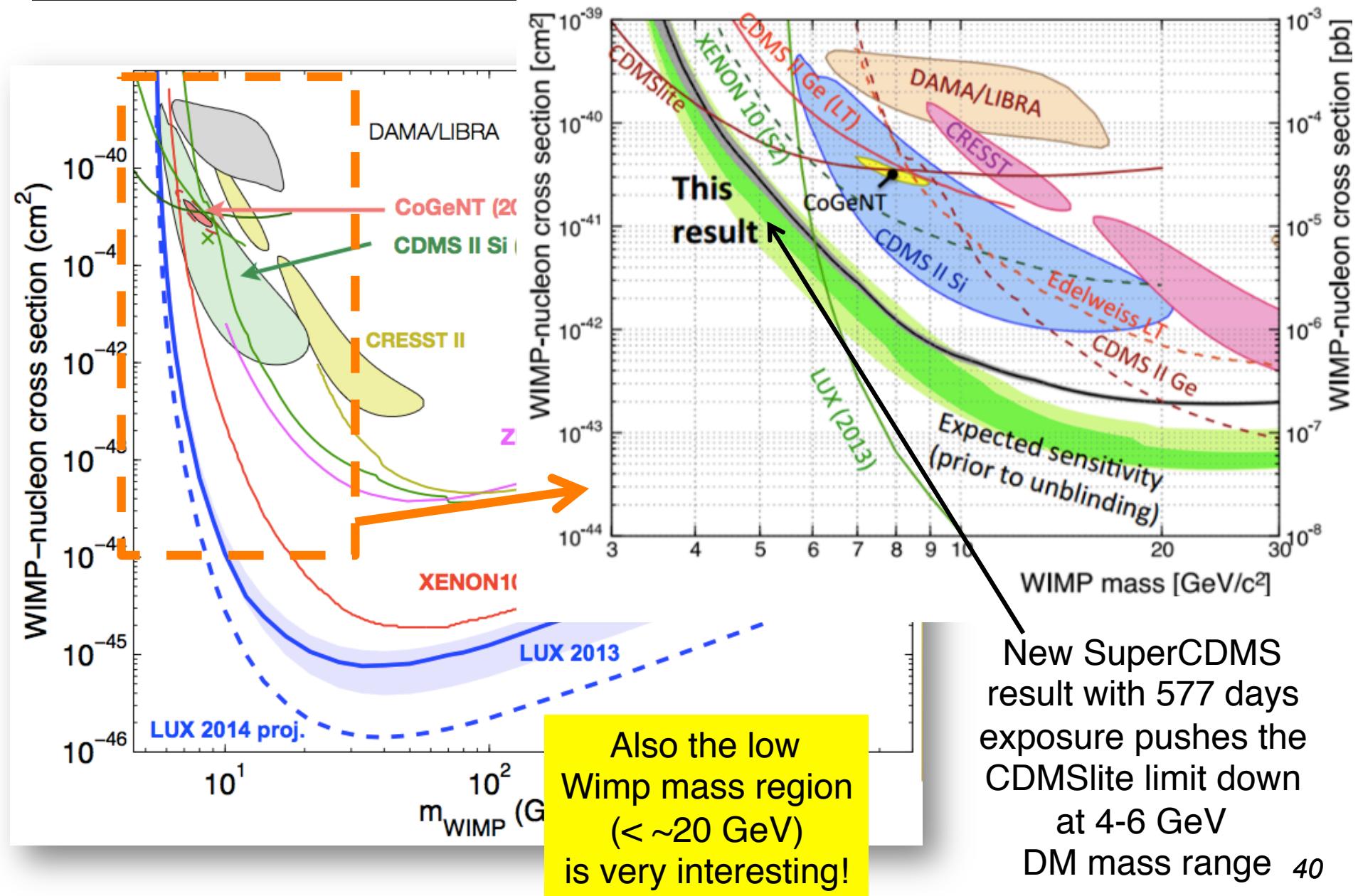
DM Searches @ collider & Direct Detection O. Buchmüller



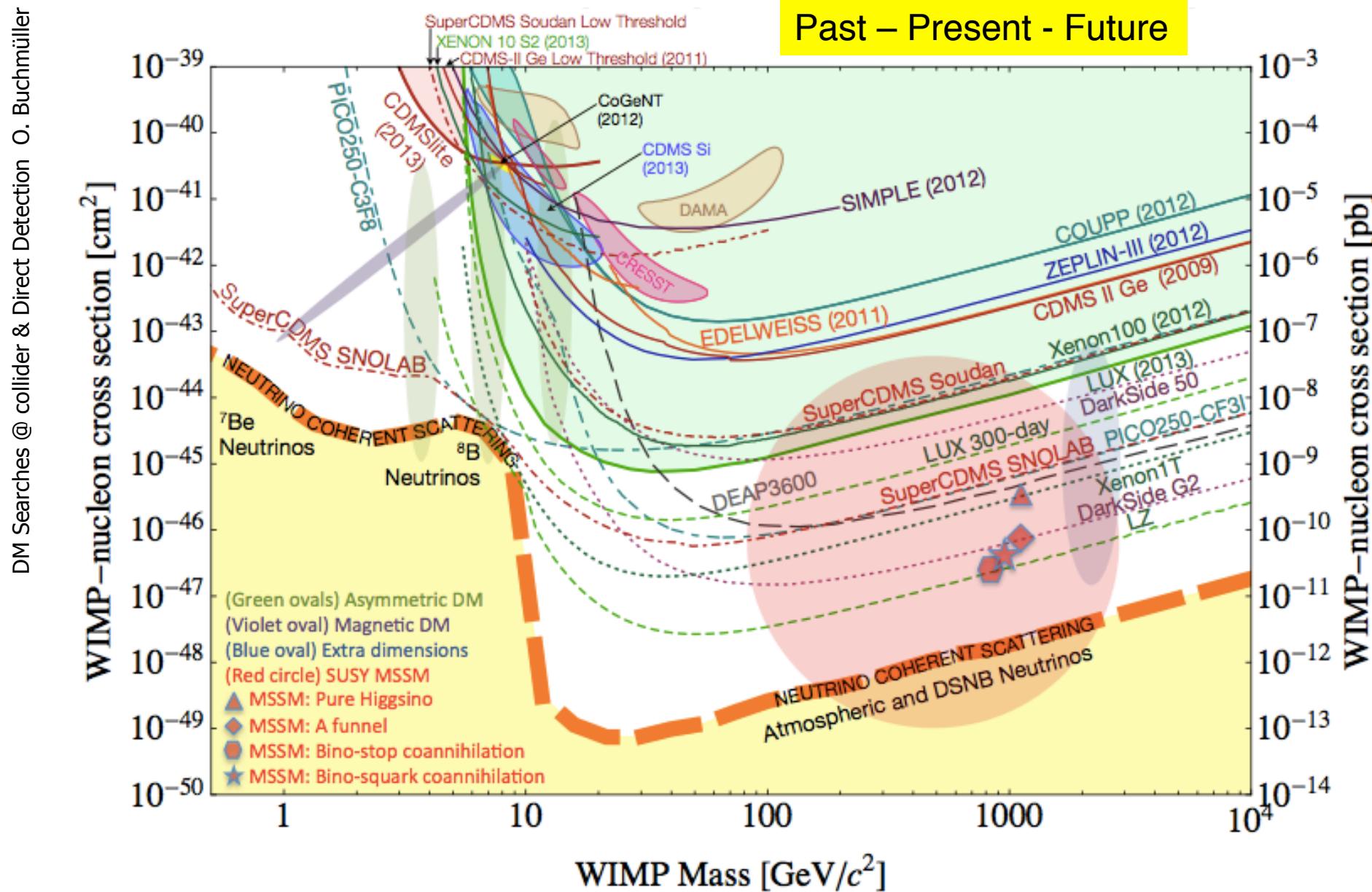
New 2013 Lux  
85.3 live-days of data  
with a fiducial volume  
of 118 kg. Conservative  
assumptions!  
Improved results  
expected soon! 39

# Dark Matter Searches: Direct Detection Experiments

DM Searches @ collider & Direct Detection O. Buchmüller

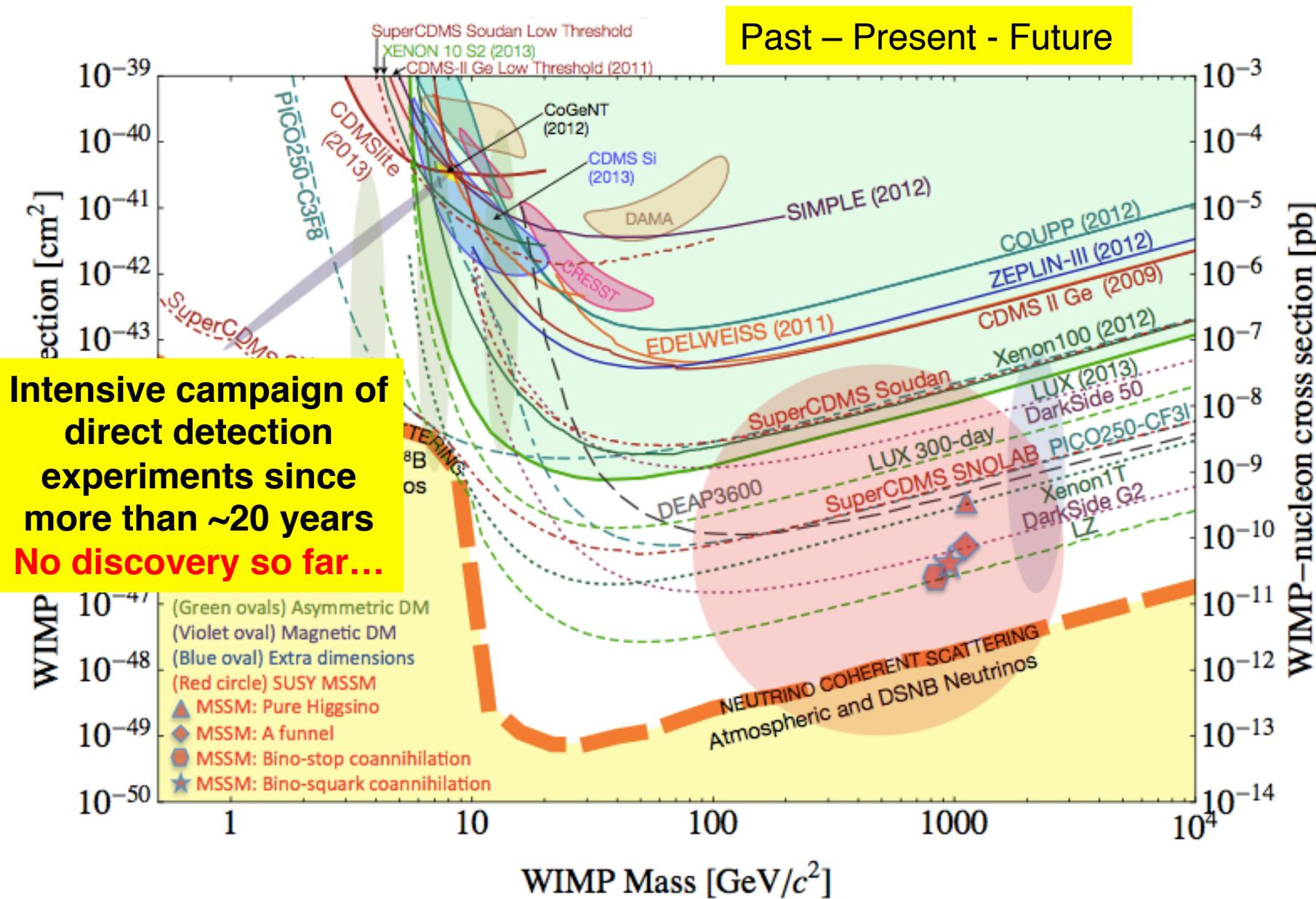


# Direct Detection Landscape in a nutshell!



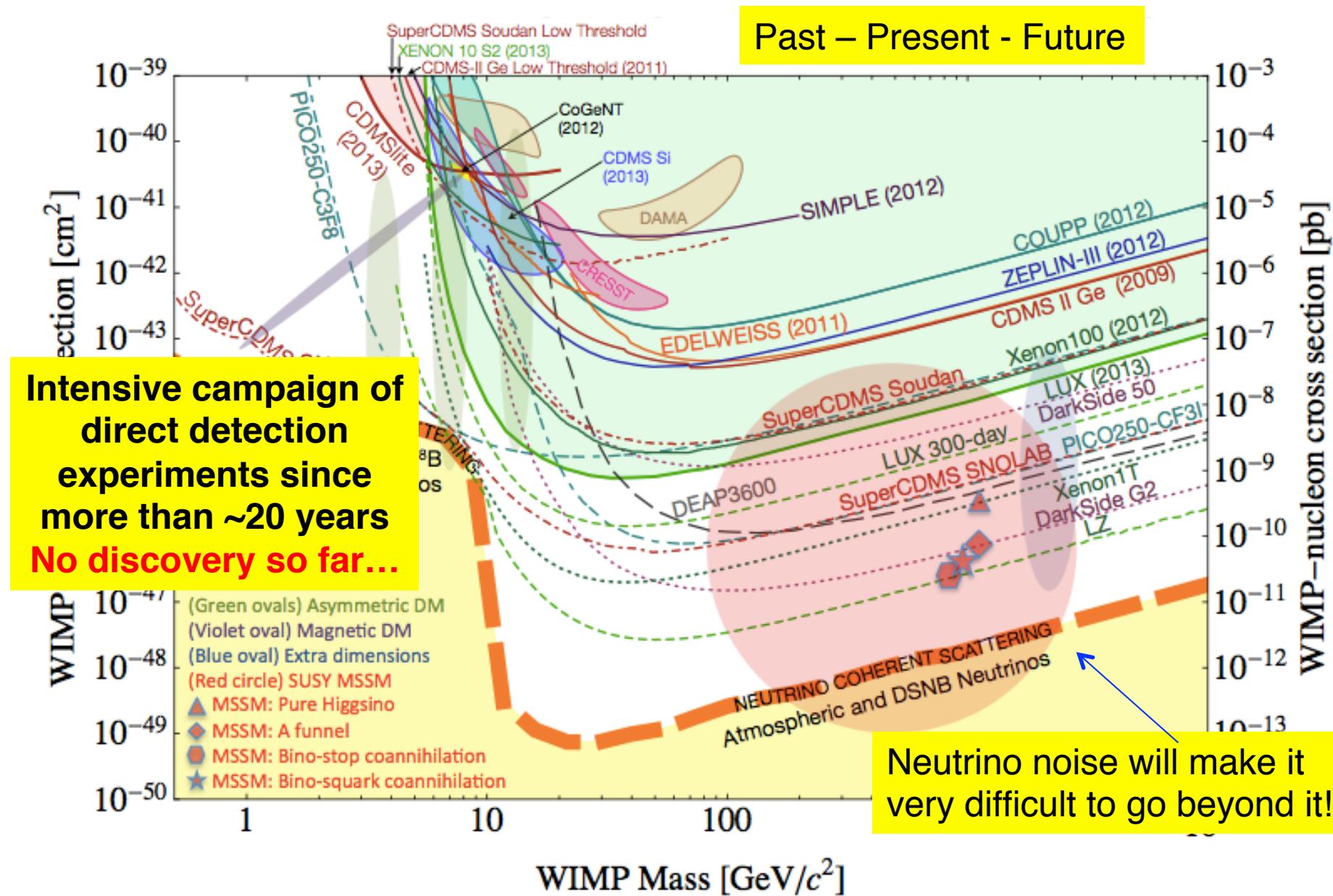
# Direct Detection Landscape in a nutshell!

DM Searches @ collider & Direct Detection O. Buchmüller



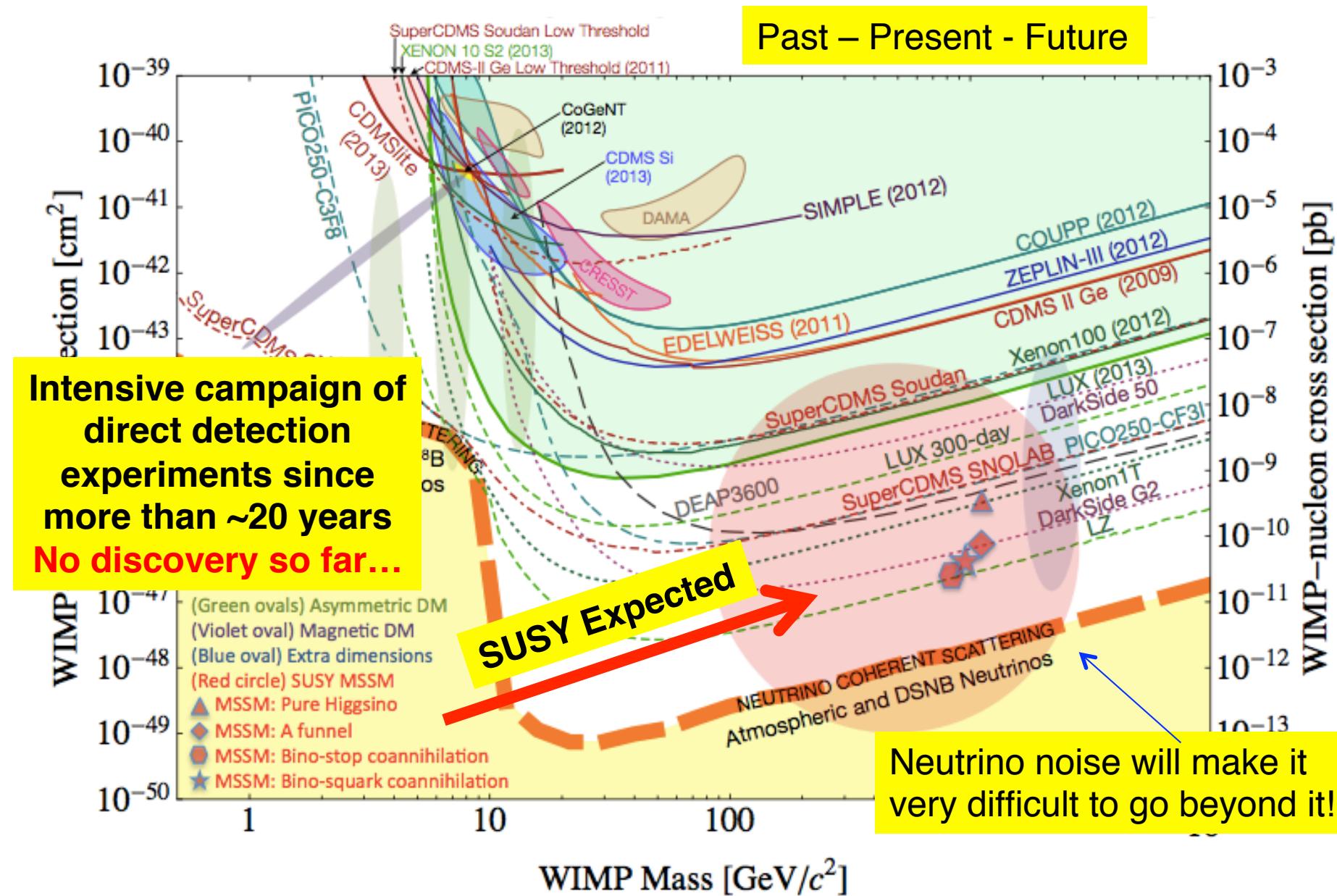
# Direct Detection Landscape in a nutshell!

DM Searches @ collider & Direct Detection O. Buchmüller



# Direct Detection Landscape in a nutshell!

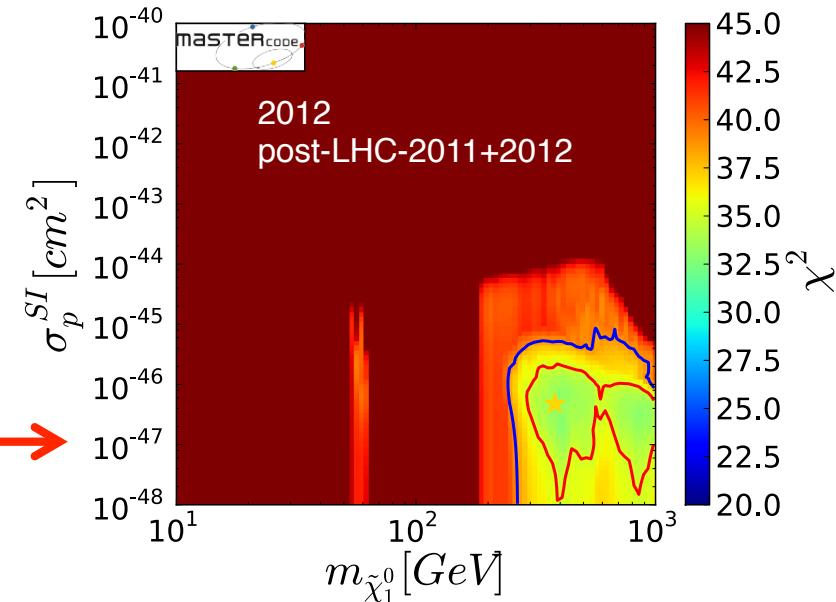
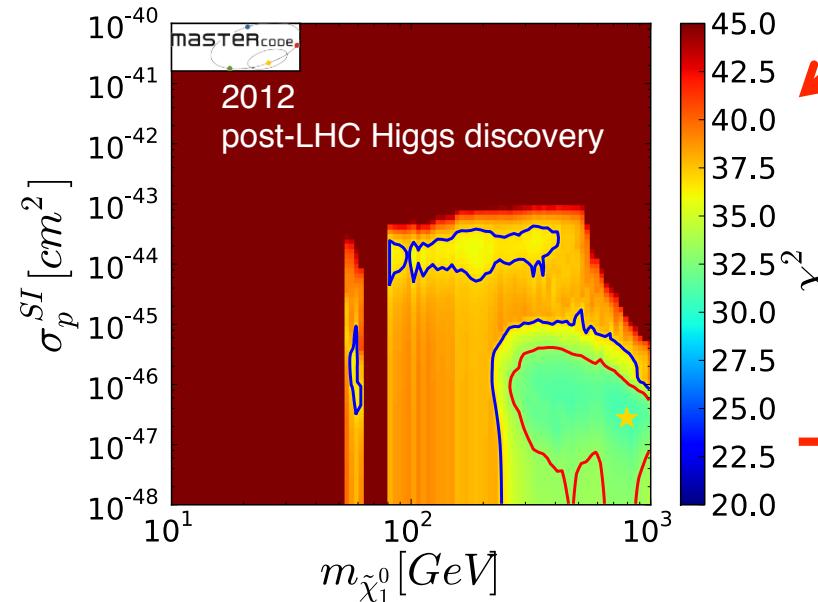
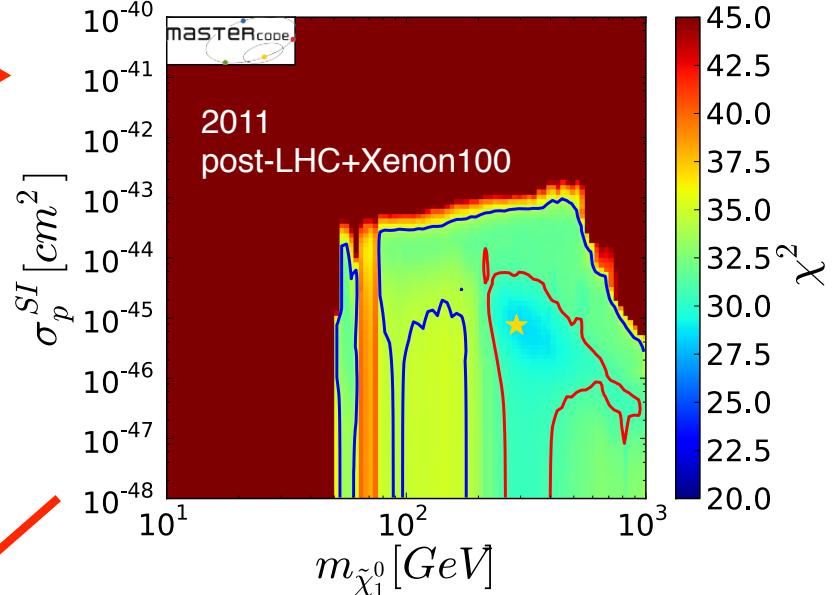
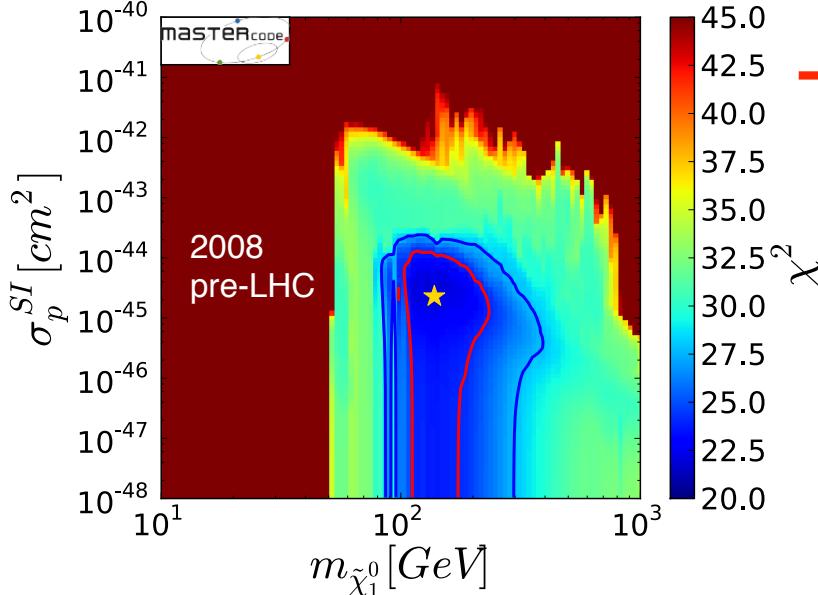
DM Searches @ collider & Direct Detection O. Buchmüller



# In Supersymmetry at Colliders and Direct Detection (i.e. a famous model ...but only one model)

# SUSY & Dark Matter: Evolution with time

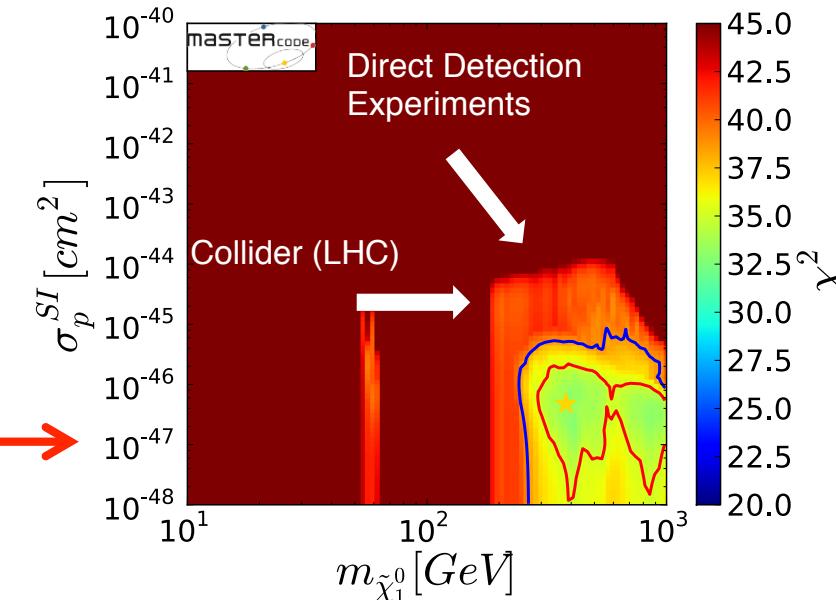
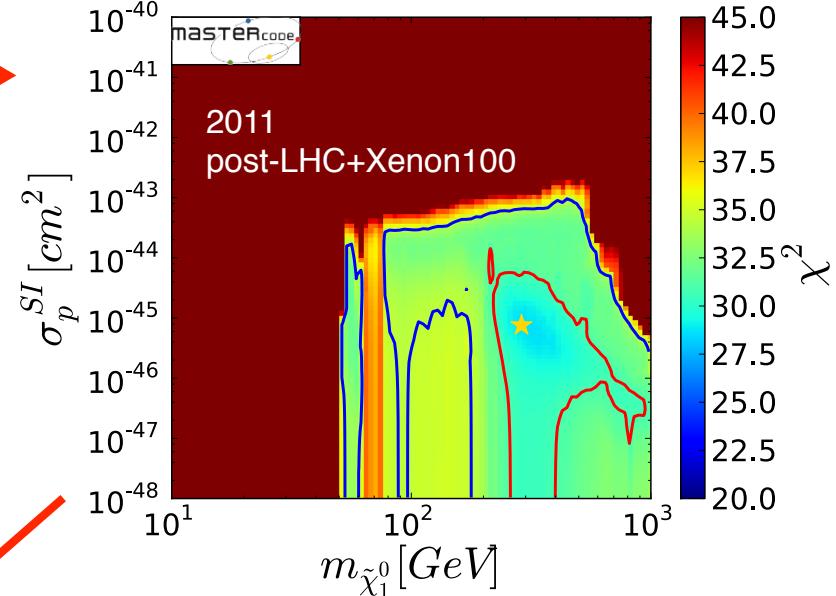
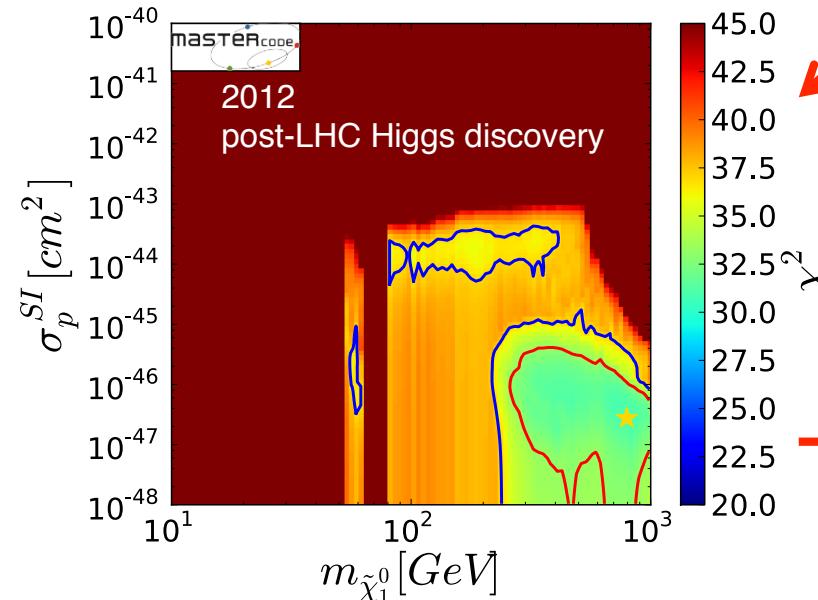
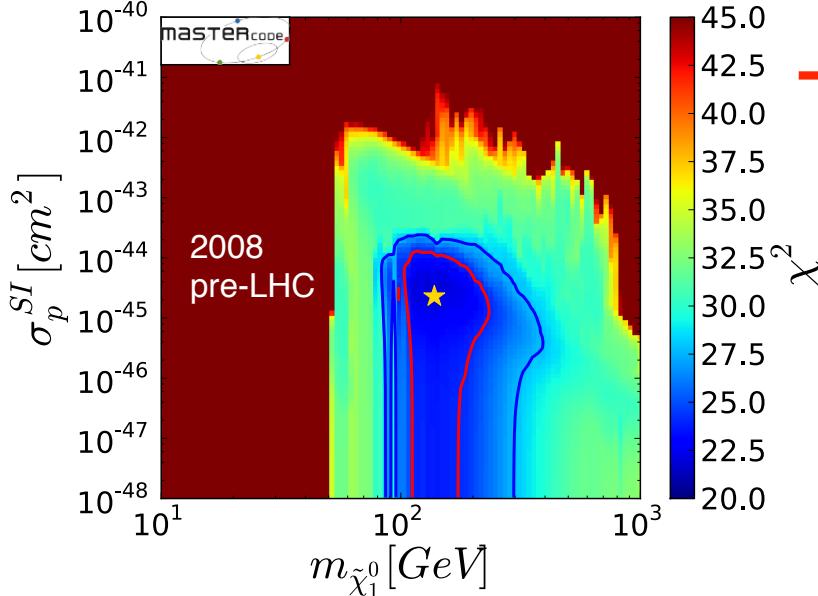
DM Searches @ collider & Direct Detection O. Buchmüller



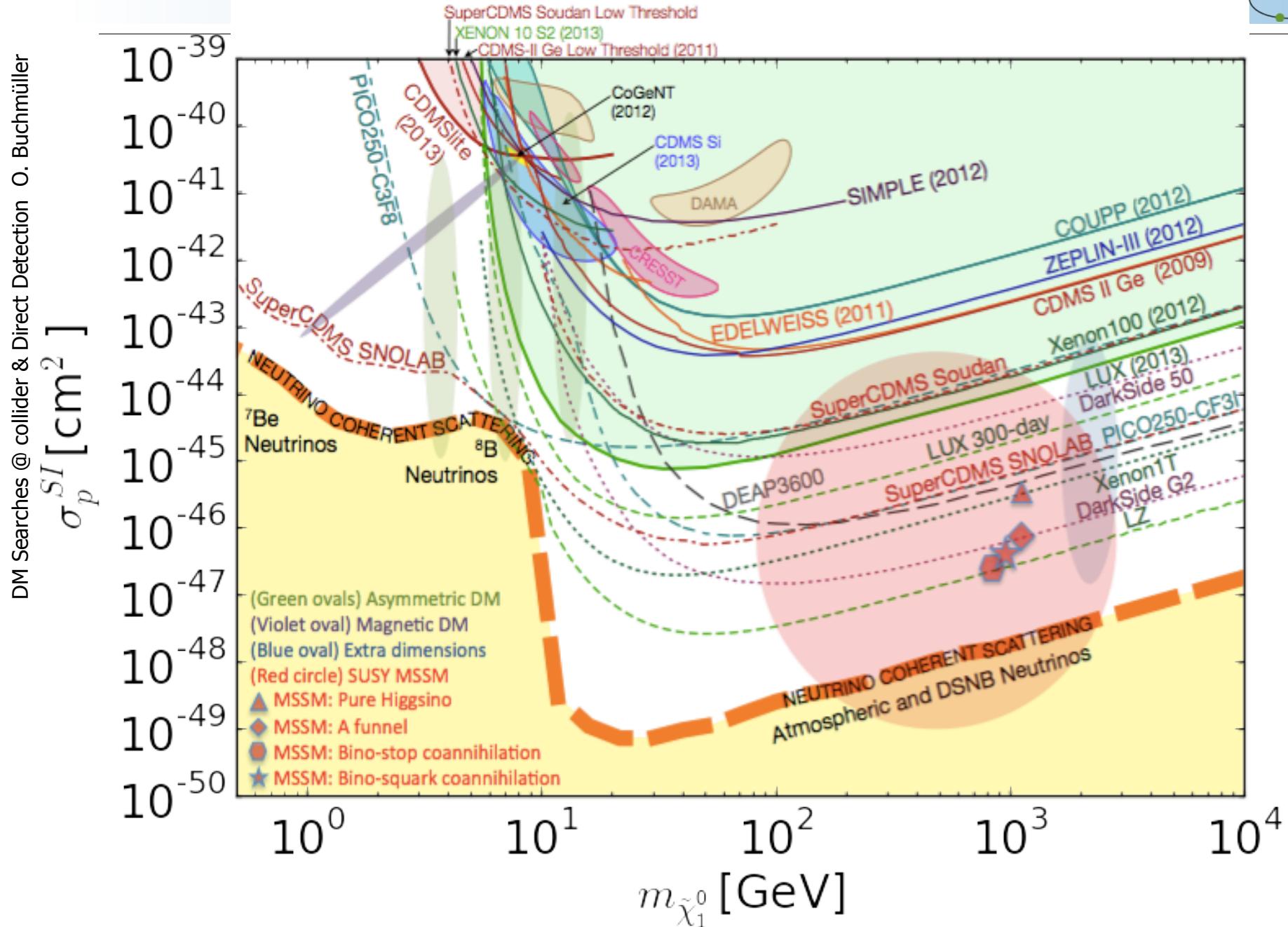
# SUSY & Dark Matter: Evolution with time



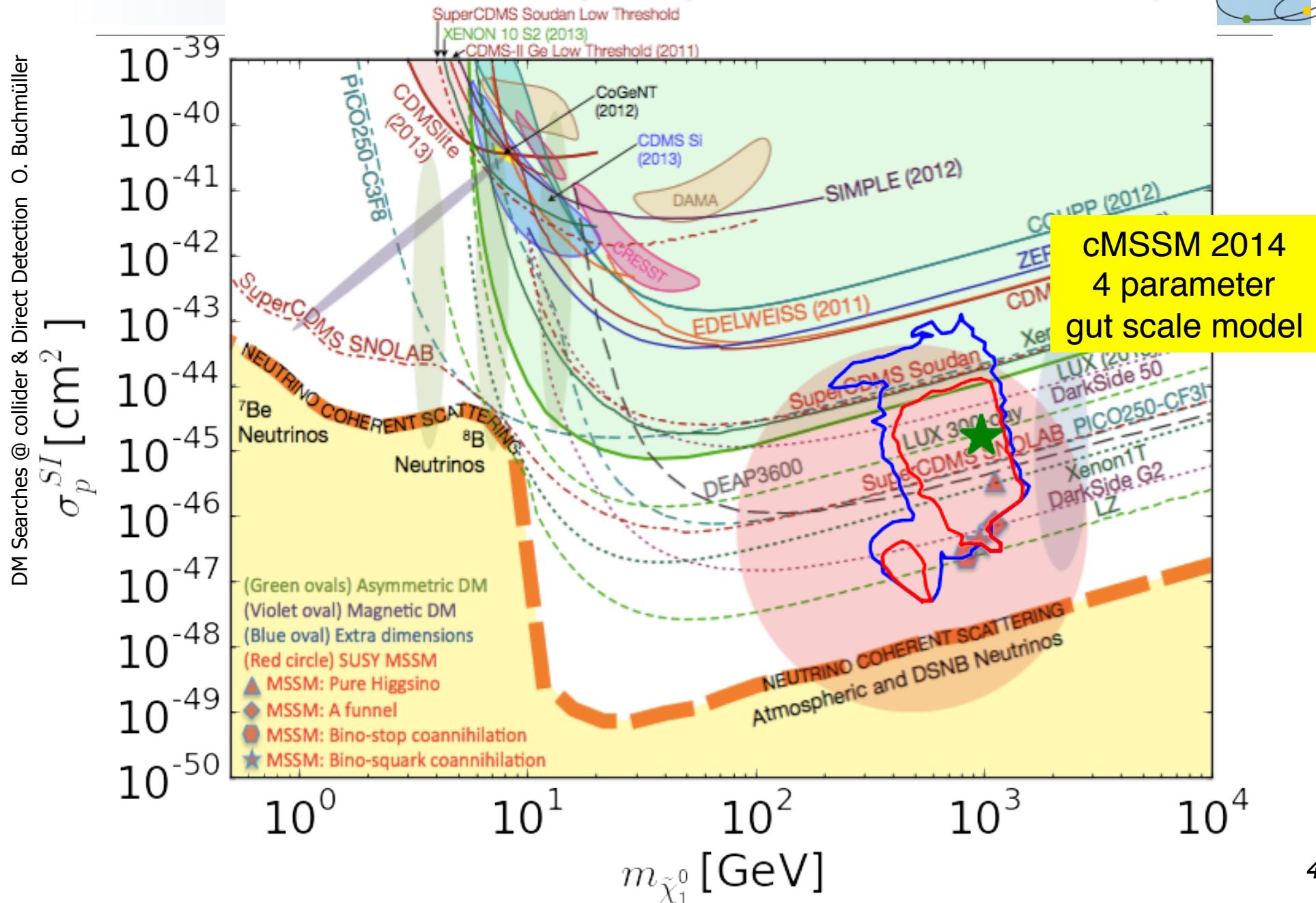
DM Searches @ collider & Direct Detection O. Buchmüller



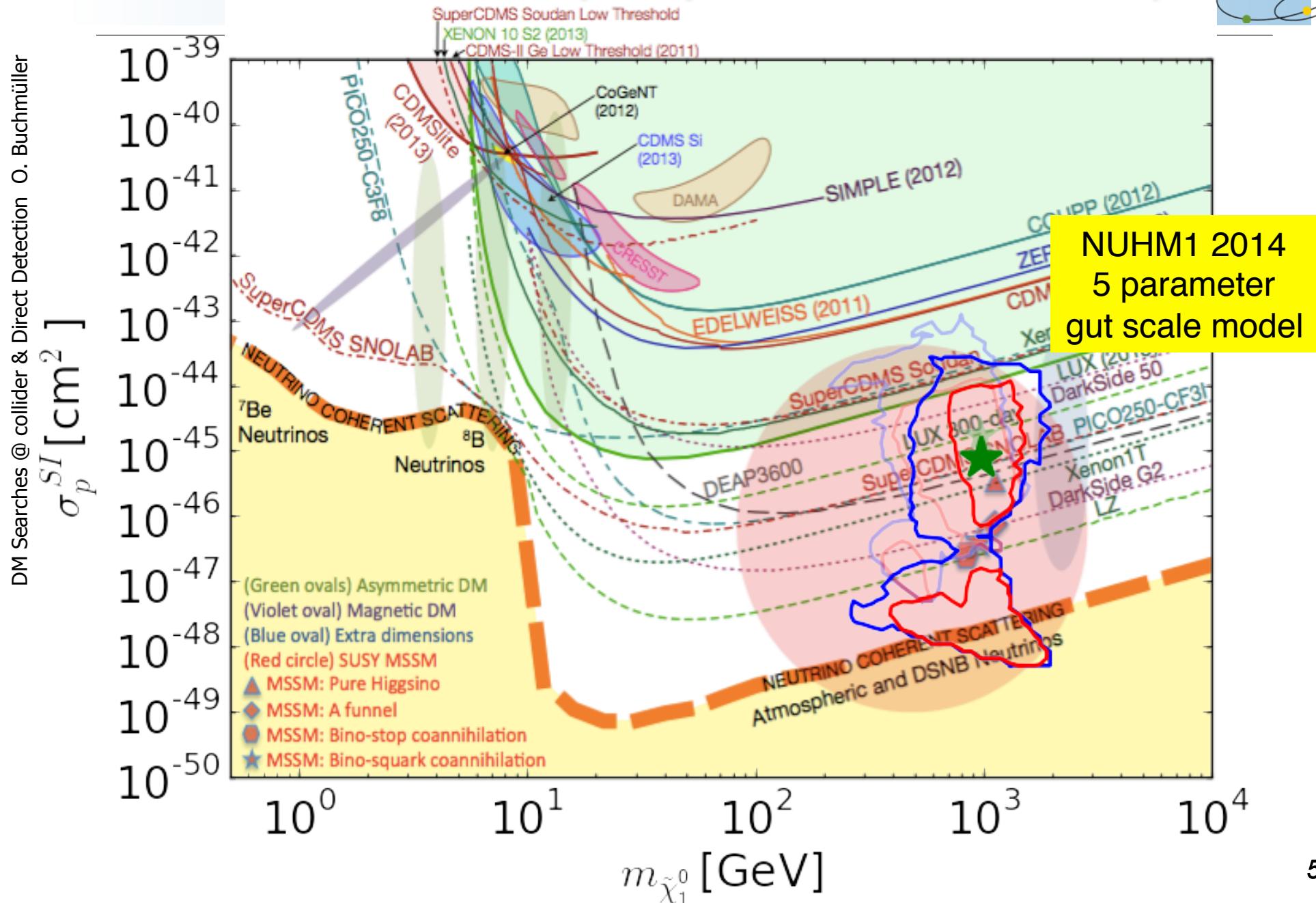
# What to expect for SUSY?



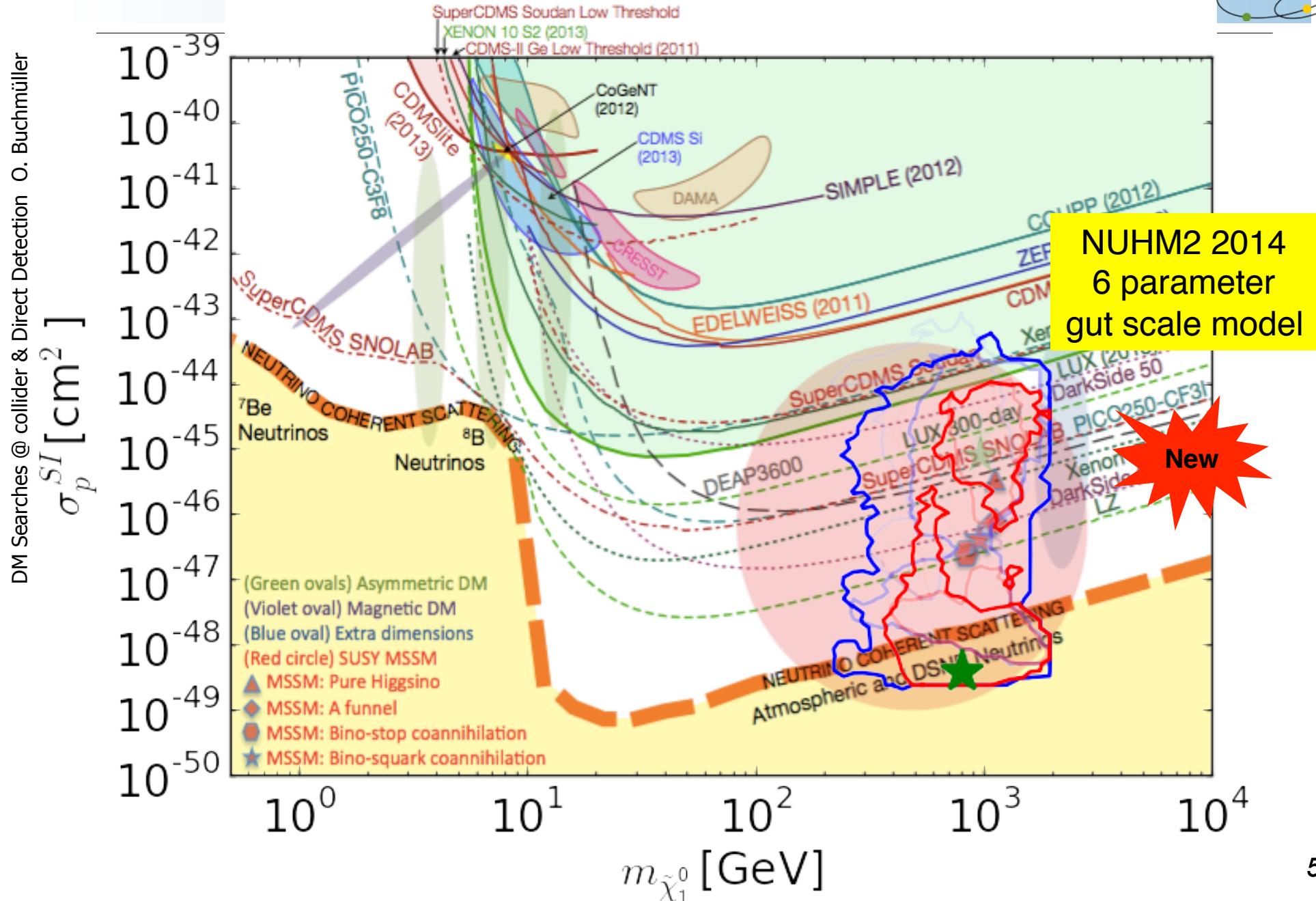
# What to expected for SUSY?



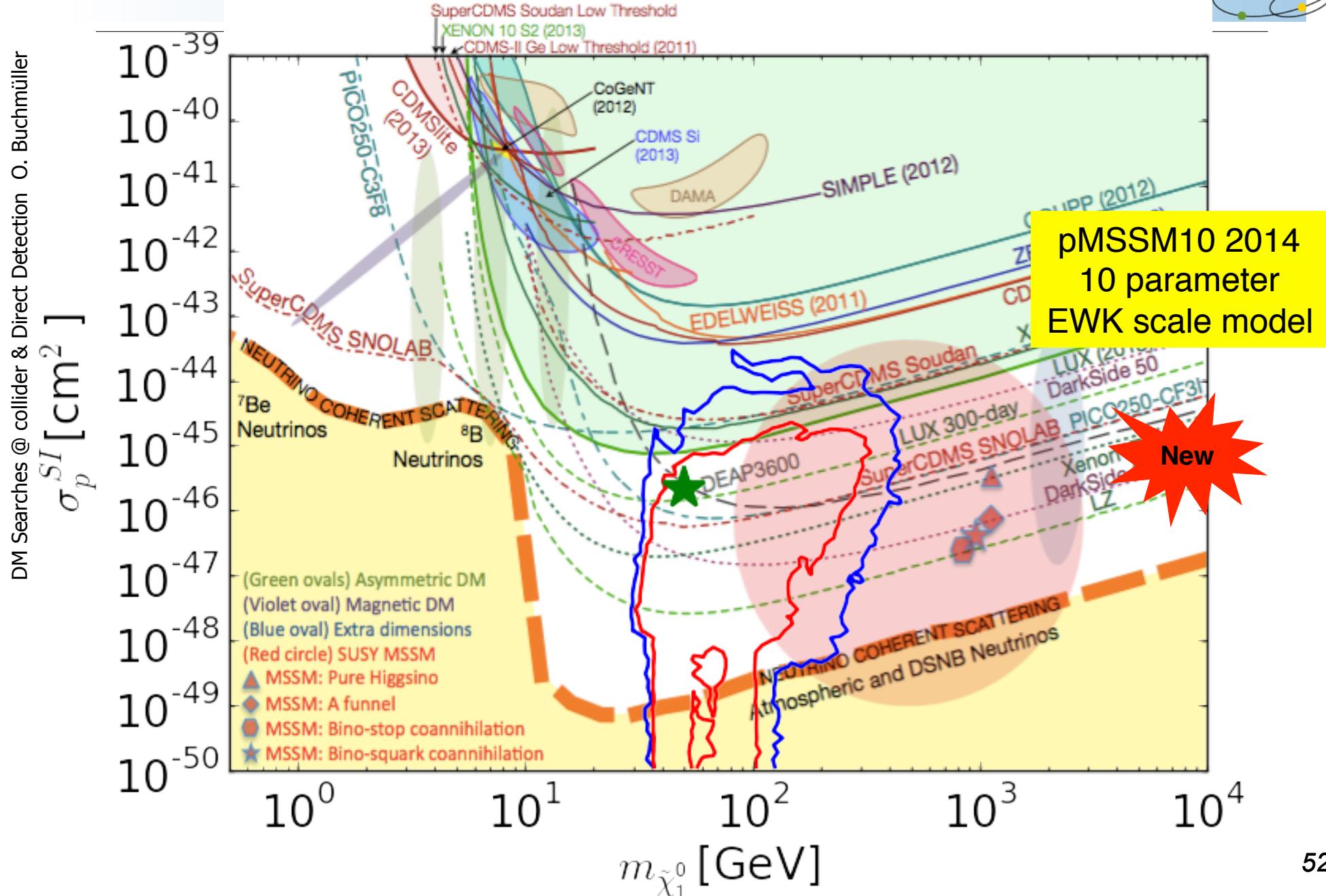
# What to expected for SUSY?



# What to expected for SUSY?



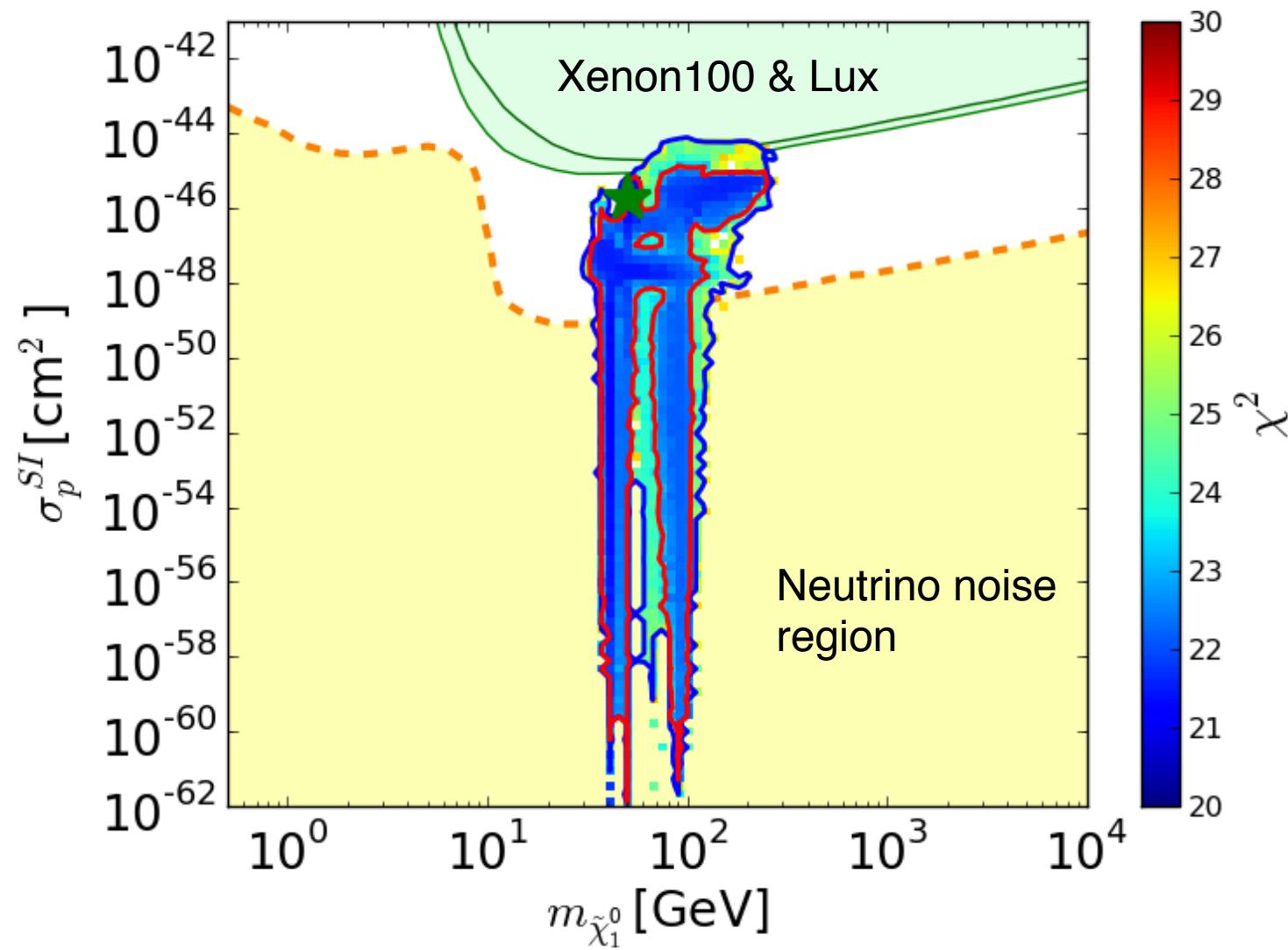
# What to expect for SUSY?



# What to expect for SUSY (pMSSM)



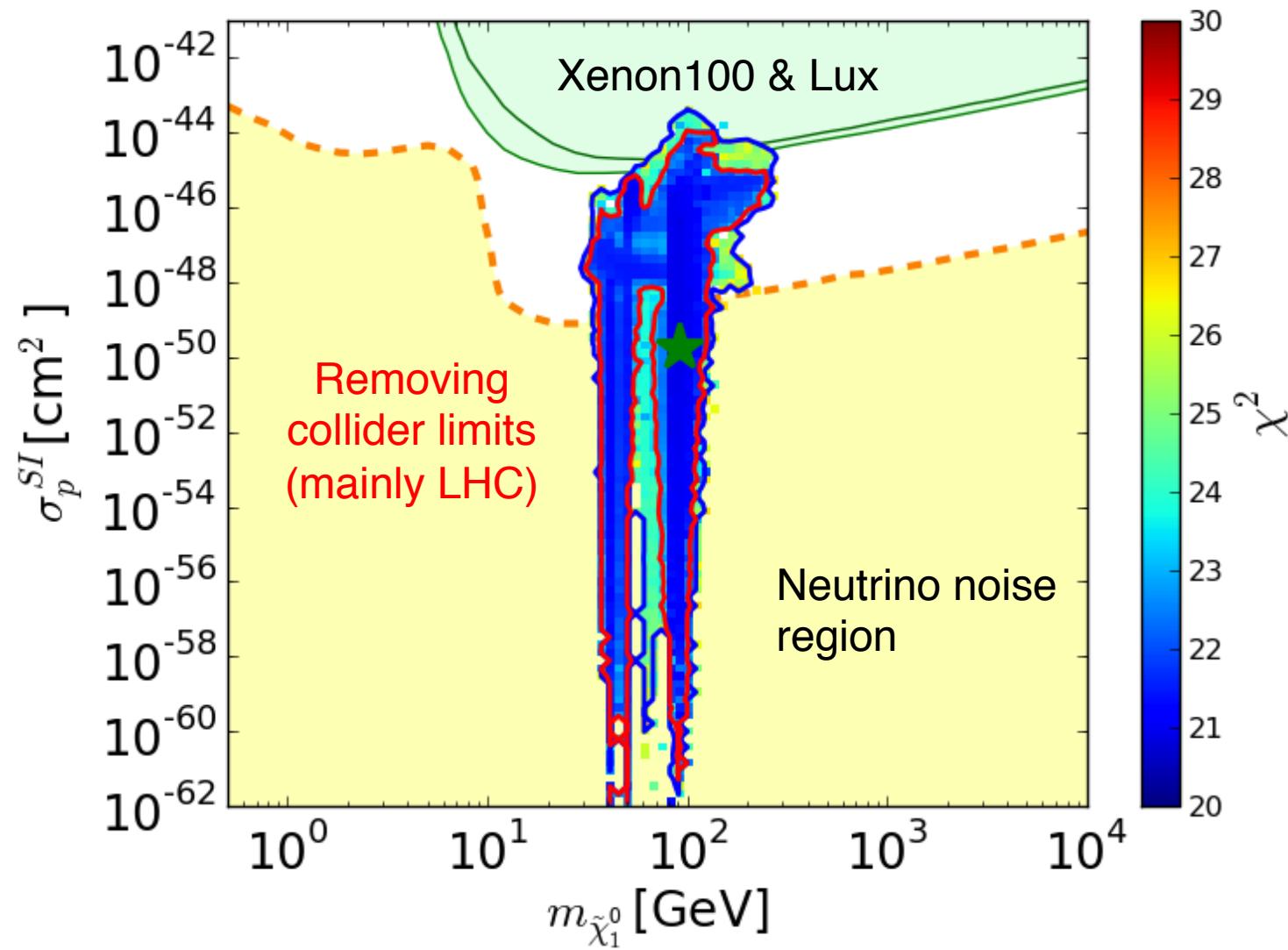
DM Searches @ collider & Direct Detection O. Buchmüller



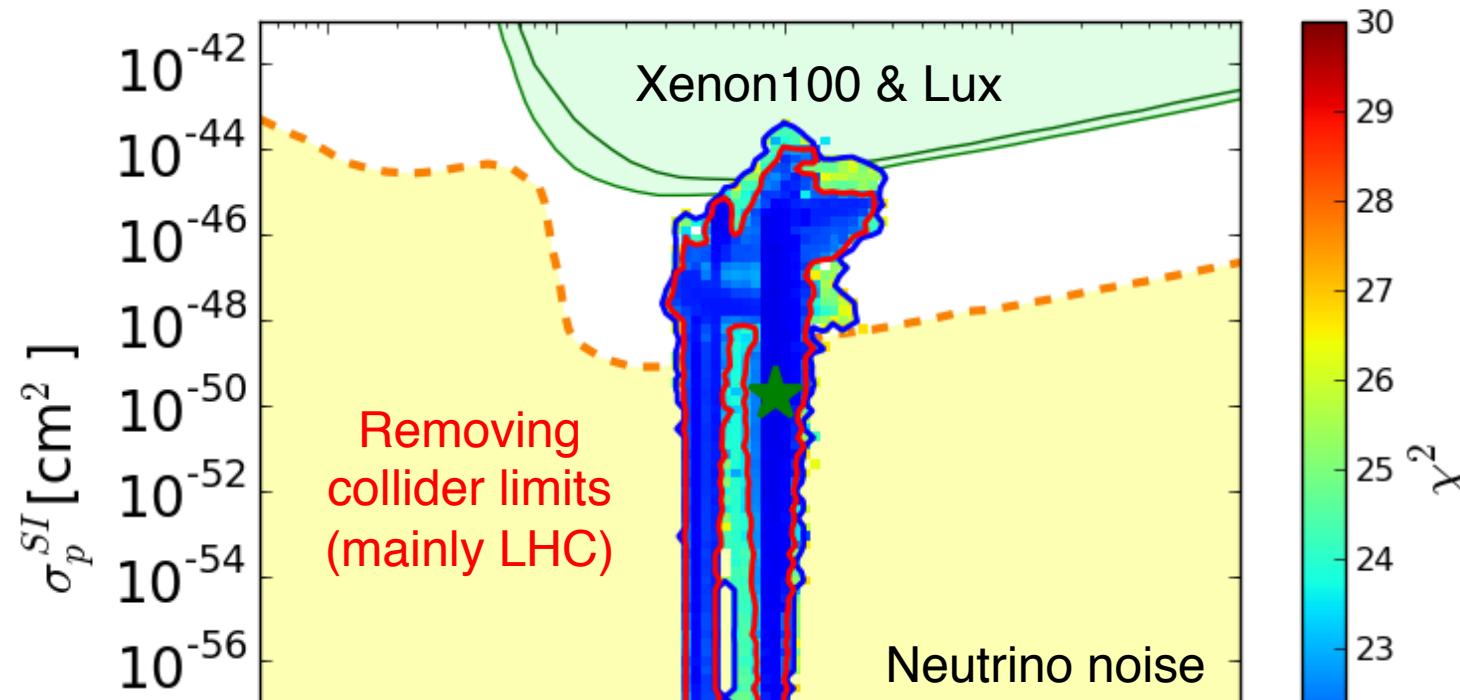
# What to expect for SUSY (pMSSM)



DM Searches @ collider & Direct Detection O. Buchmüller



# What to expect for SUSY (pMSSM)

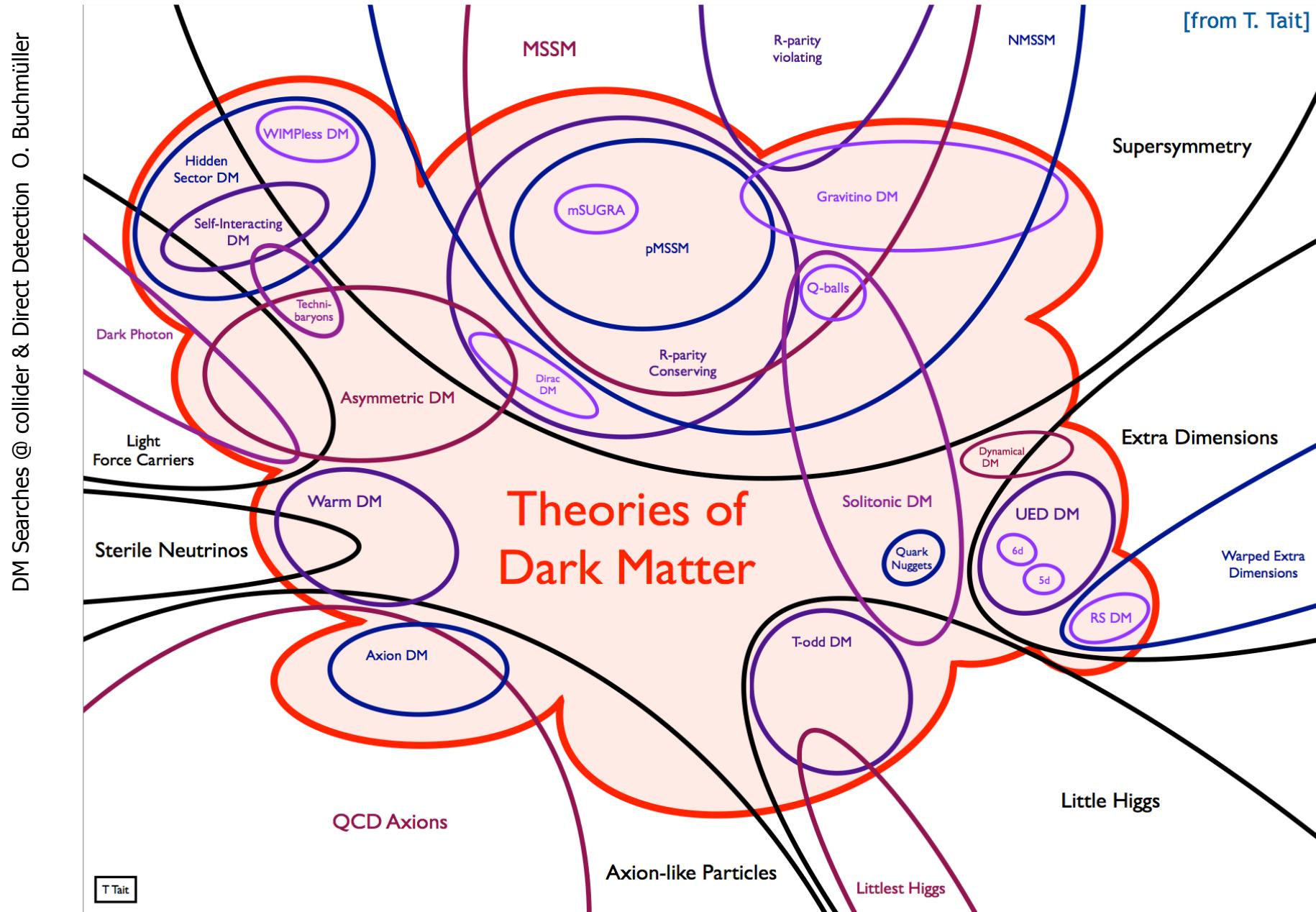


## SUSY summary:

- ◆ There remains still a lot of unexplored parameter space in SUSY
  - ◆ Large regions are within the neutrino noise region
  - ◆ LHC searches are can probe regions not accesible to DM Direct detection experiments
- **Need to better work out complementarity between collider and DD**

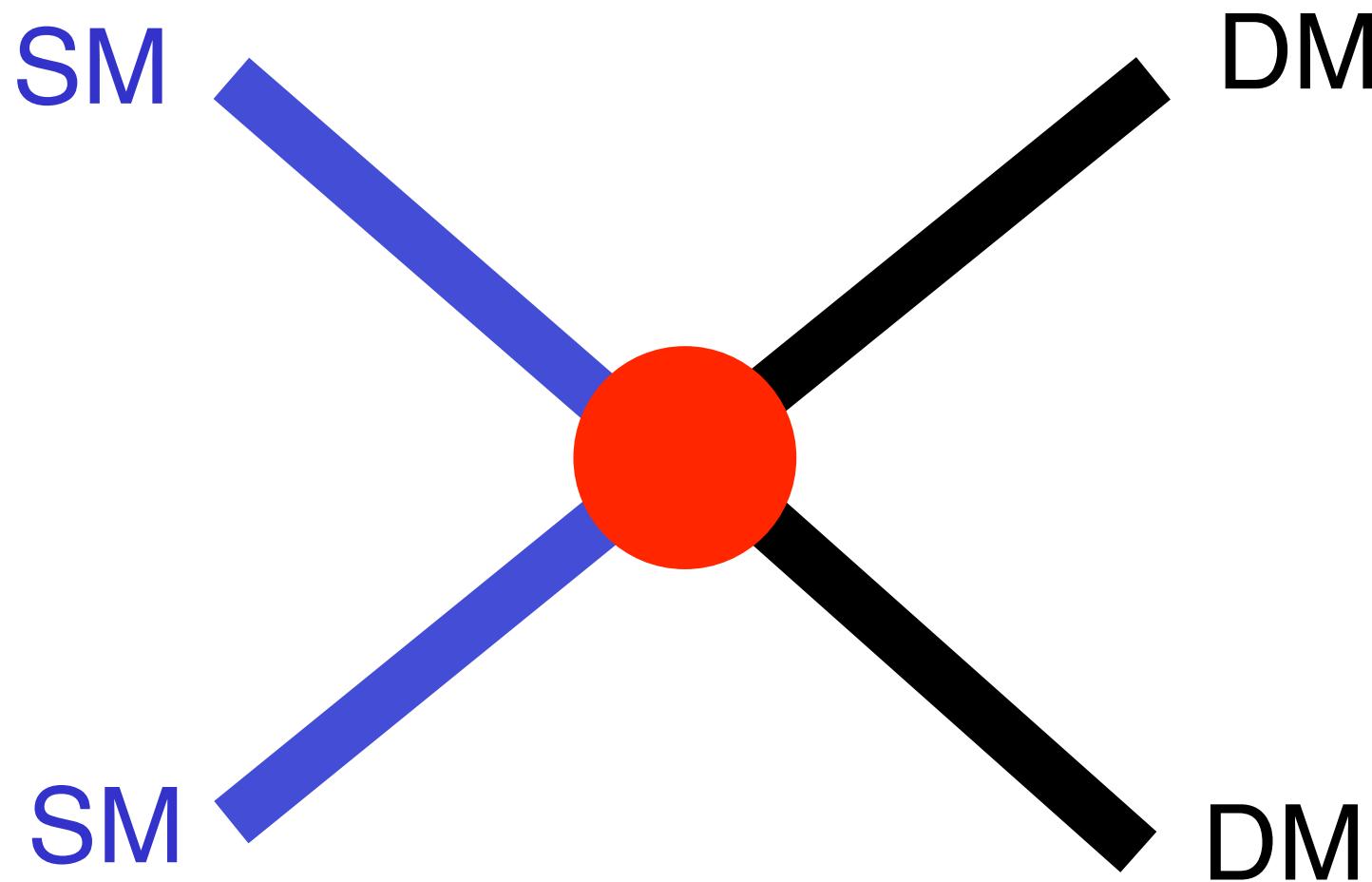
# Beyond Models! Generic Direct Searches for Dark Matter

# The Theoretical Landscape of DM Theories



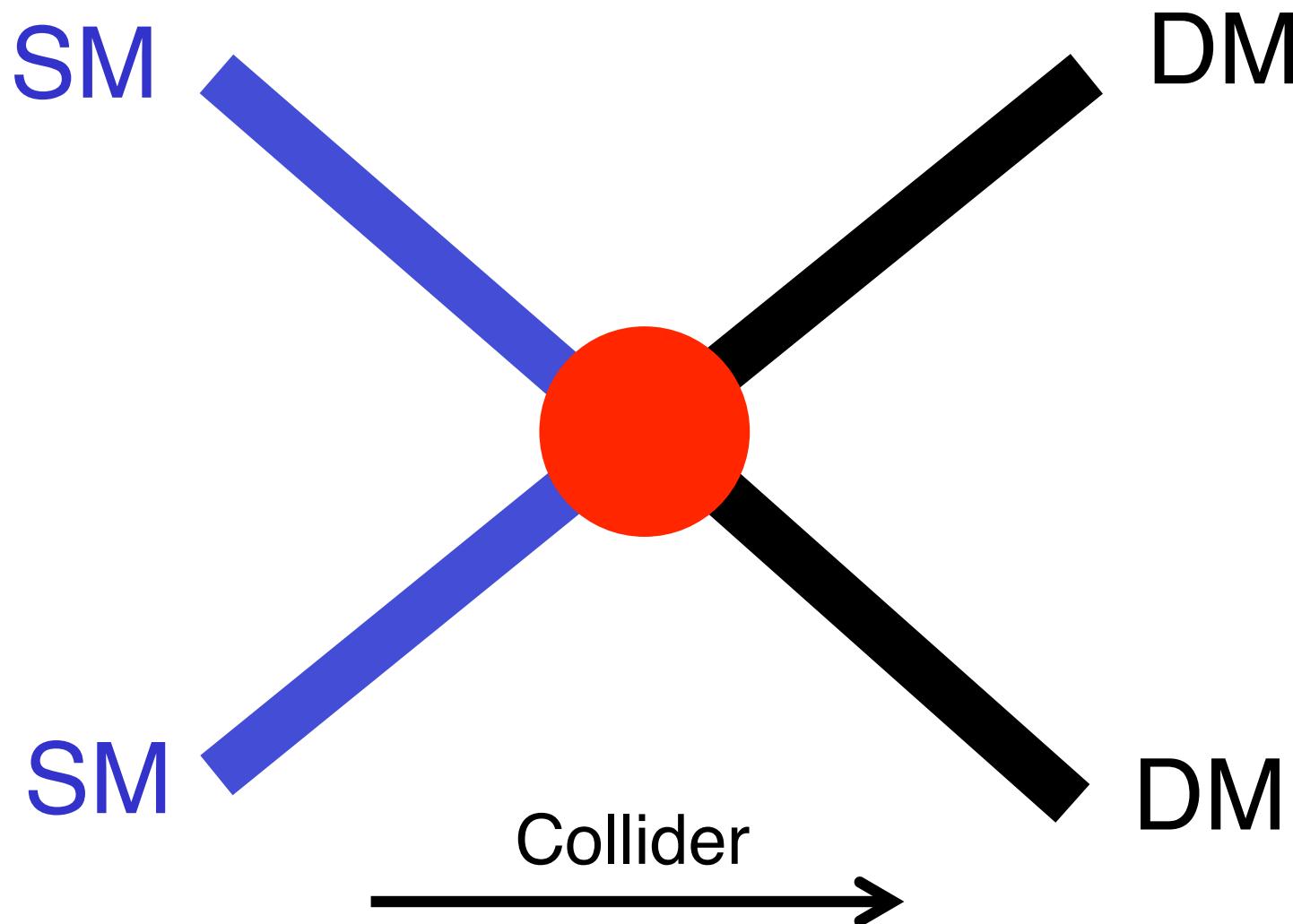
## Dark Matter Searches

DM Searches @ collider & Direct Detection O. Buchmüller



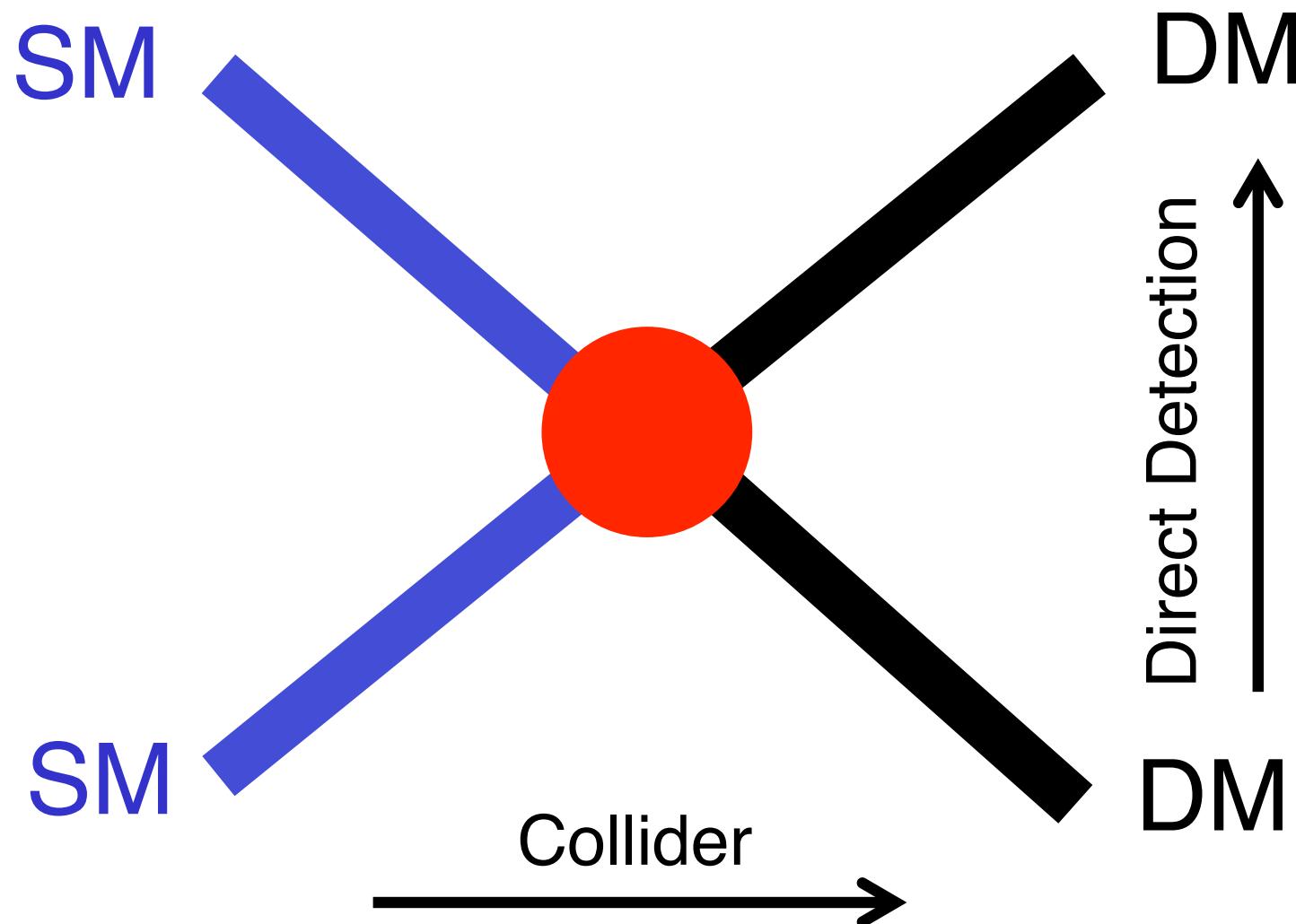
## Dark Matter Searches

DM Searches @ collider & Direct Detection O. Buchmüller



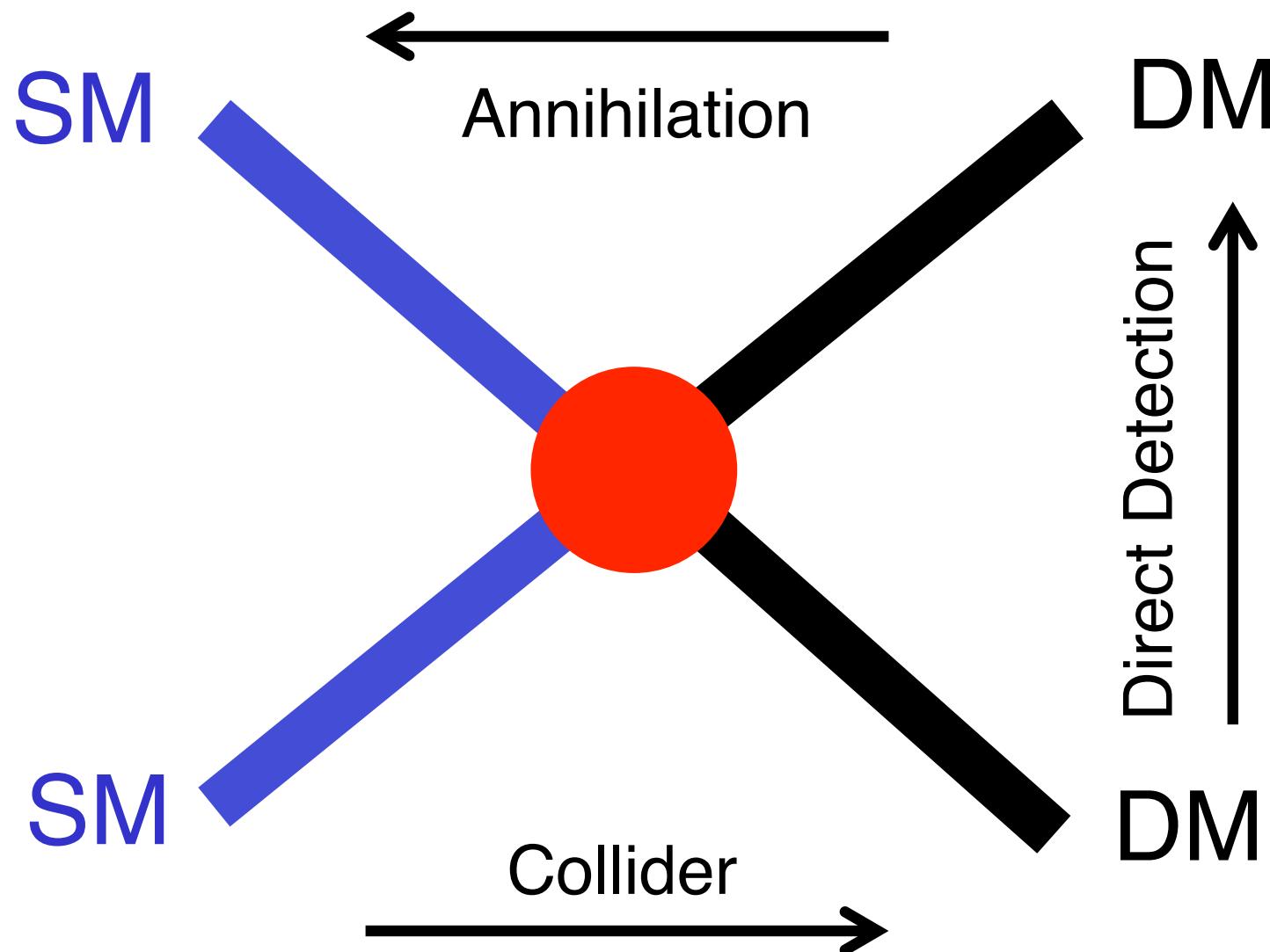
## Dark Matter Searches

DM Searches @ collider & Direct Detection O. Buchmüller



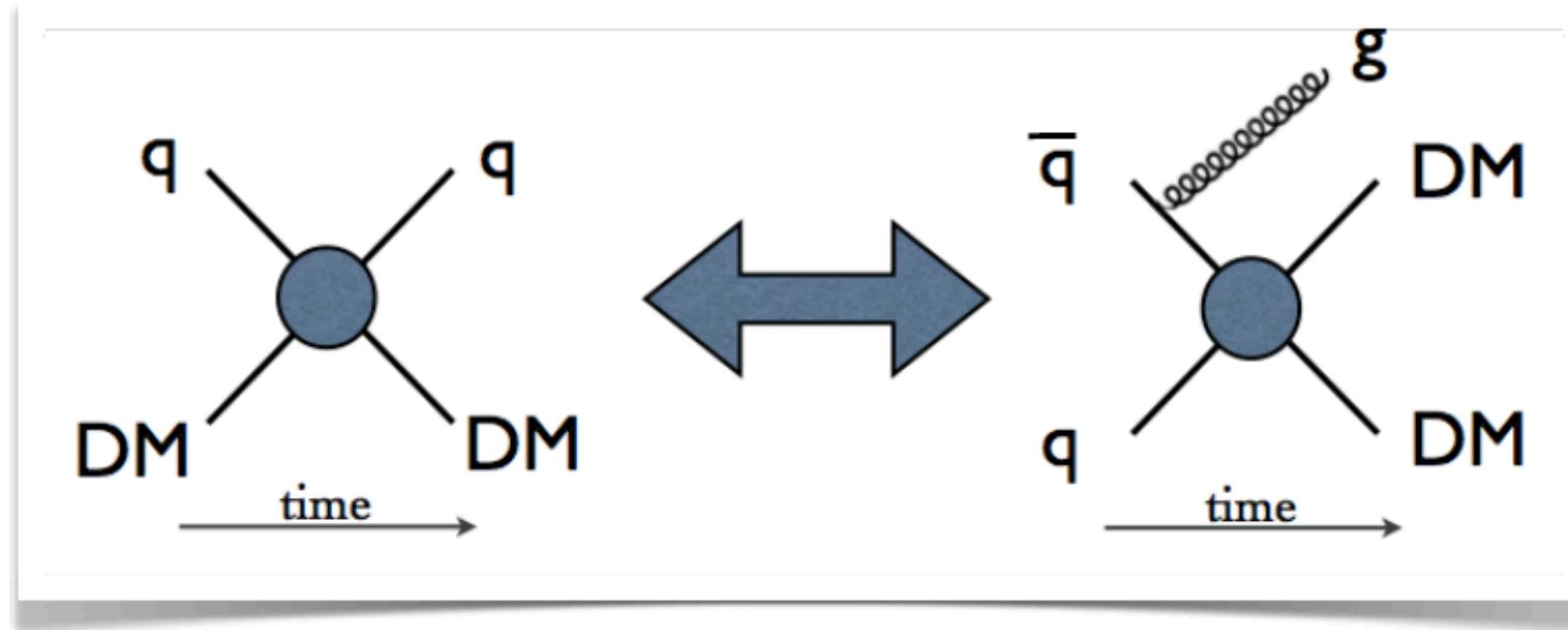
## Dark Matter Searches

DM Searches @ collider & Direct Detection O. Buchmüller



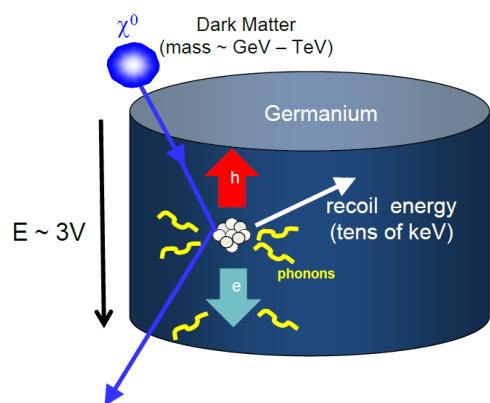
# Dark Matter Searches: Direct Detection vs Colliders

DM Searches @ collider & Direct Detection O. Buchmüller



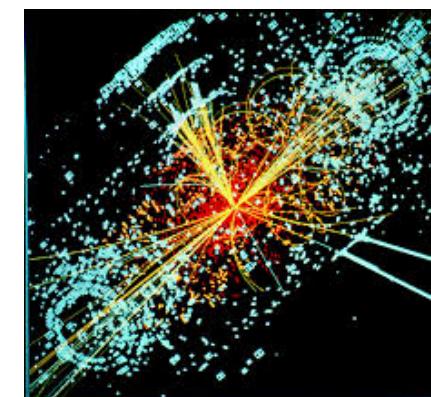
## Direct Detection Experiments

- DM-nucleus scattering



## Collider Experiments

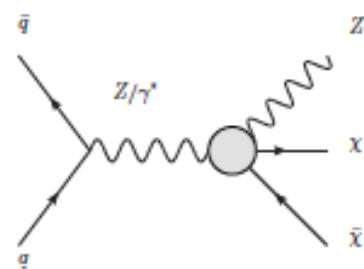
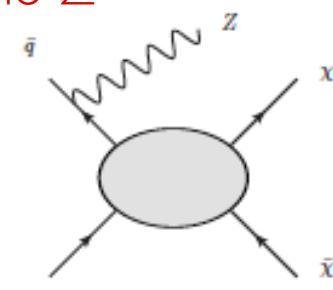
- Pair-production of DM
- missing energy signature



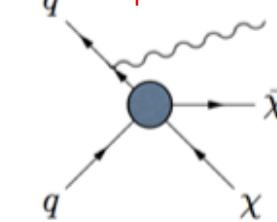
# Mono-Mania (at the LHC)

DM Searches @ collider & Direct Detection O. Buchmüller

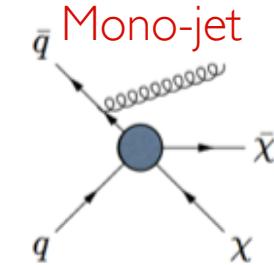
Mono-Z



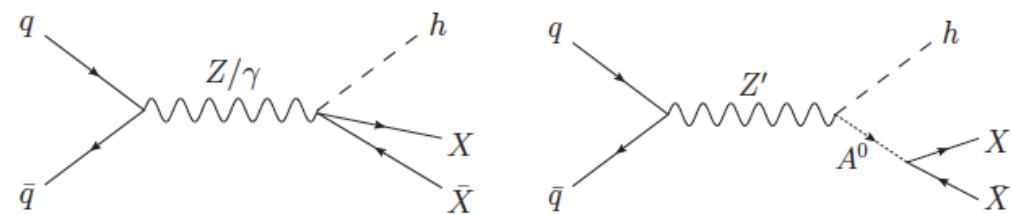
Mono-photon



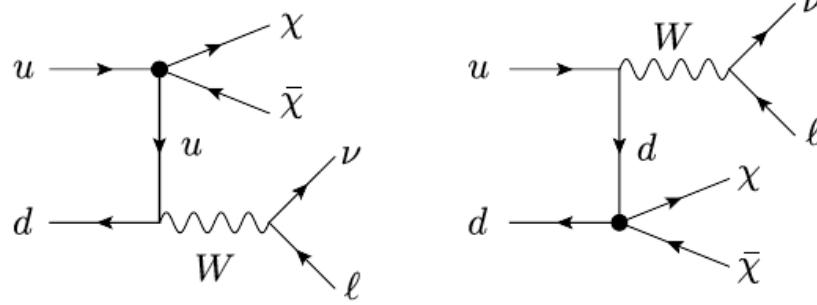
Mono-jet



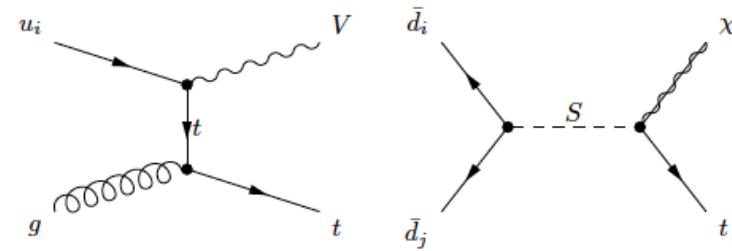
Mono-Higgs



Mono-W



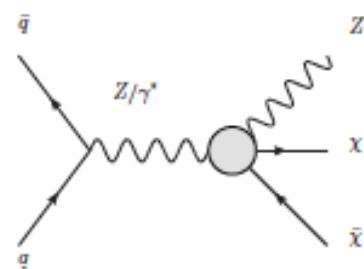
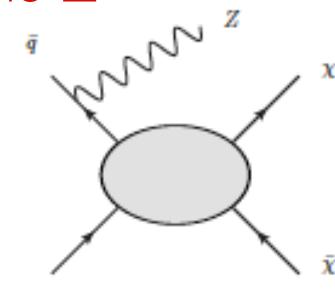
Mono-top



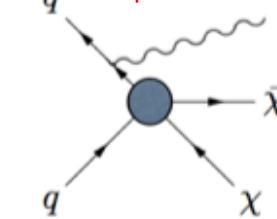
# Mono-Mania (at the LHC)

DM Searches @ collider & Direct Detection O. Buchmüller

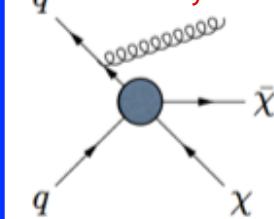
Mono-Z



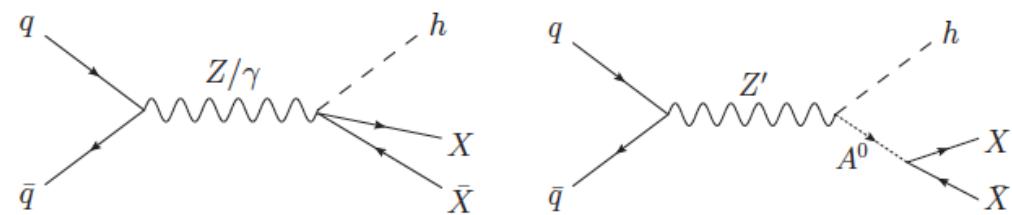
Mono-photon



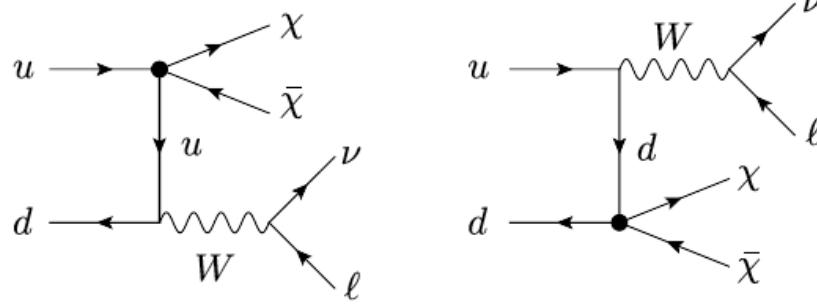
Mono-jet



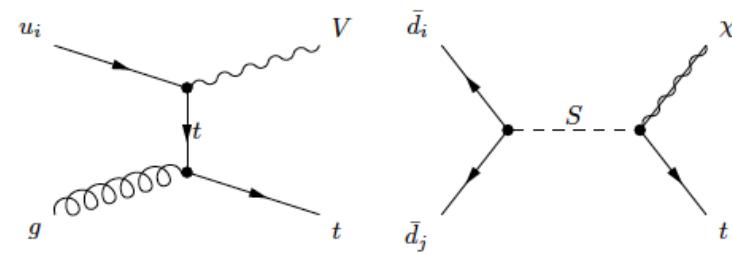
Mono-Higgs



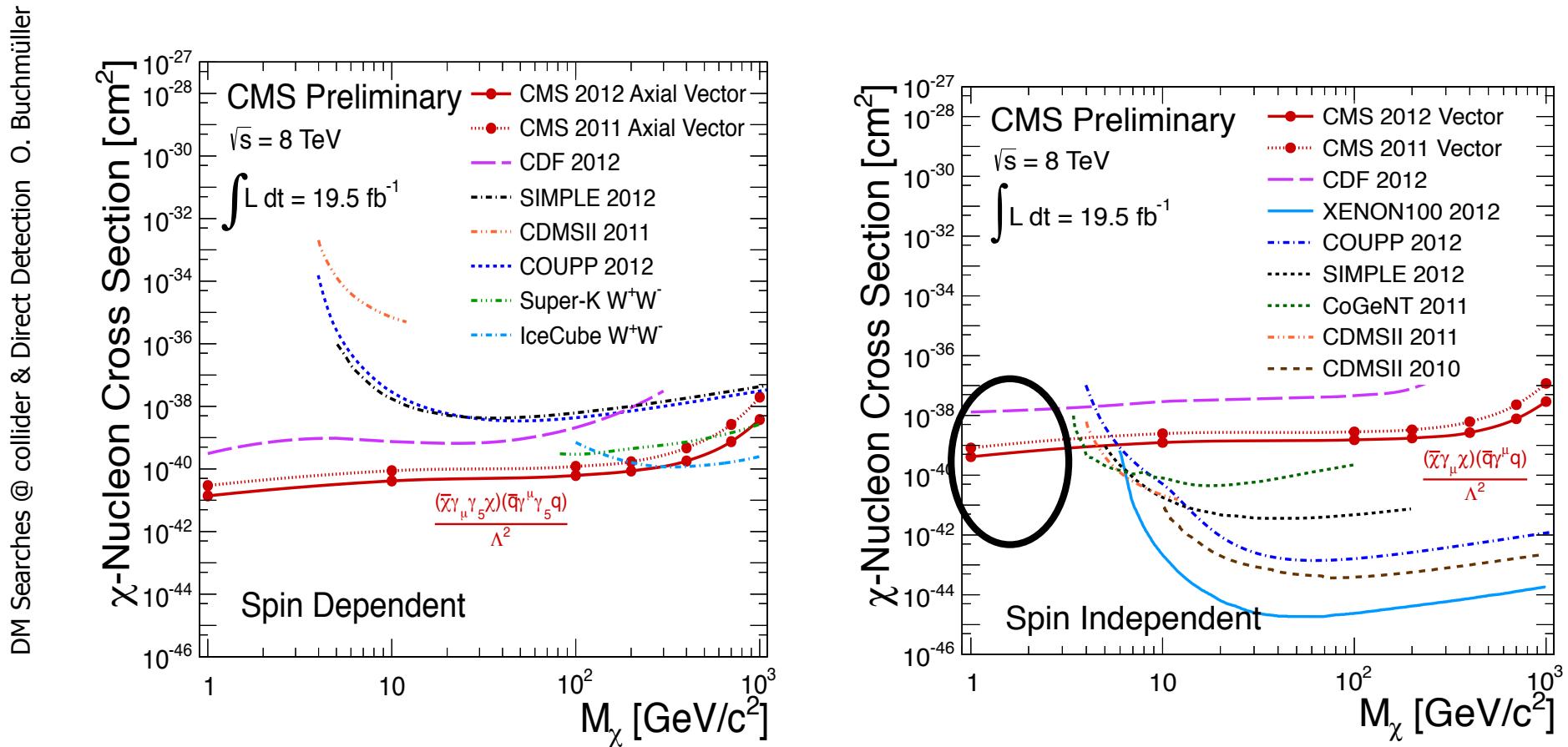
Mono-W



Mono-top



# Monojet analyses better than direct detection?!



Claim [often made]:

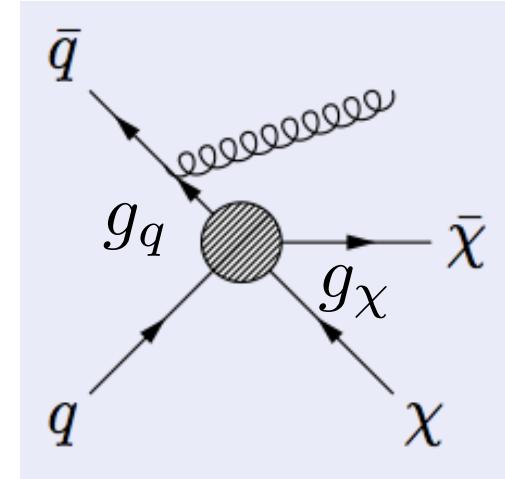
For **low mass** and the entire **spin-dependent** case monojet limits are stronger than direct detection limits!

# Effective Field Theory (EFT) Interpretation

Example of considered operators:

$$O_V = \frac{(\bar{\chi} \gamma_\mu \chi)(\bar{q} \gamma_\mu q)}{\Lambda^2} \quad \text{Vector operator, s-channel}$$

$$O_{AV} = \frac{(\bar{\chi} \gamma_\mu \gamma_5 \chi)(\bar{q} \gamma_\mu \gamma_5 q)}{\Lambda^2} \quad \text{Axial vector operator, s-channel}$$



## Assumption of EFT

If the operator (e.g. V or AV) mediator is **suitably(!!)** heavy it can be integrated out to obtain the effective V or AV contact operator. **In this case (and only this case)**, the contact interaction scale  $\Lambda$  is related to the parameters entering the Lagrangian:

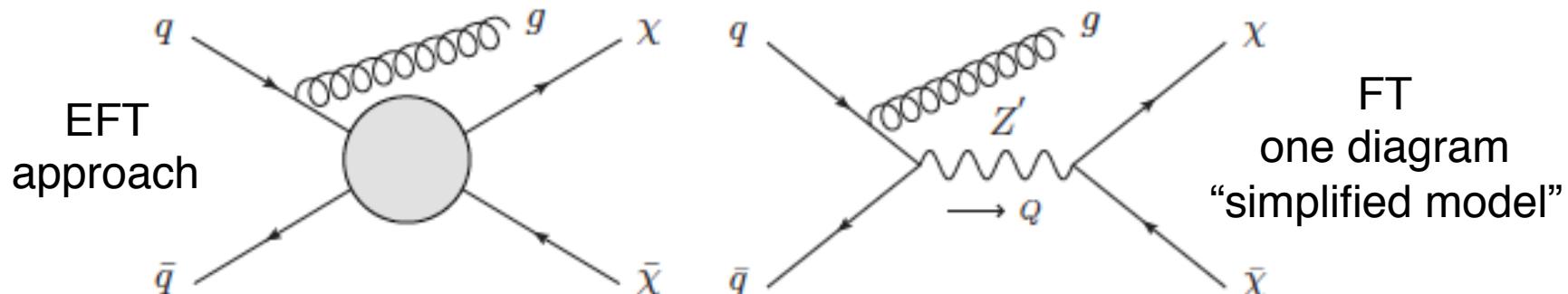
$$\Lambda = \frac{M_{mediator}}{\sqrt{g_q g_\chi}} \quad (\text{relation in the full theory})$$

# Validity of Effective Field Theory Limits

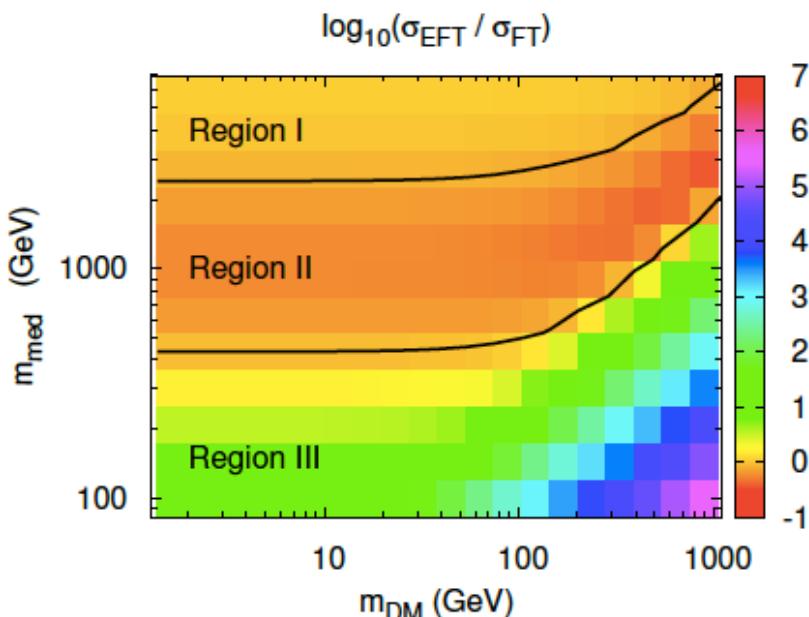
↳ collider & Direct Detection O. Buchmüller

Recent work from OB, M.Dolan,C.McCabe: arXiv:1308.6799

- Compare Effective Field Theory (EFT) with Full Theory (FT)



Use vector and axial-vector mediators (e.g.  $Z'$ ) as example - scalar are similar in conclusion!



Compare prediction of FT with EFT in  $m_{\text{med}} - m_{\text{DM}}$  plane.  
Three regions become visible:

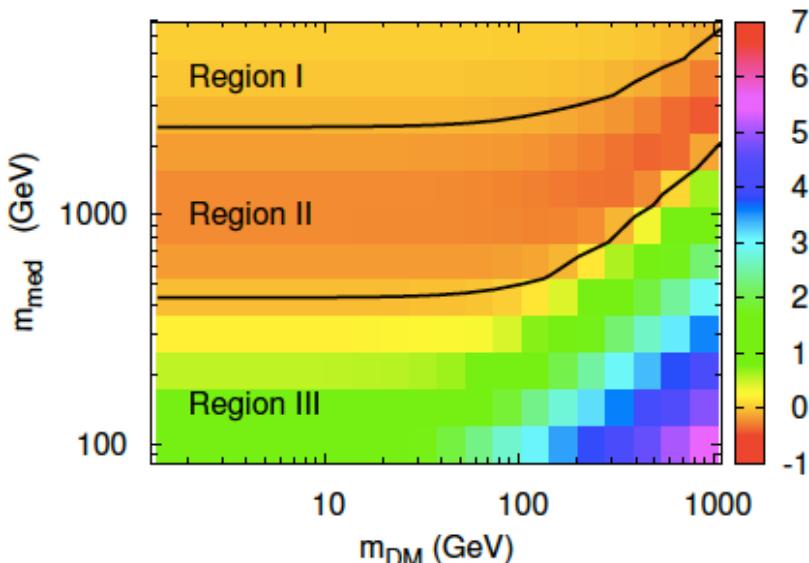
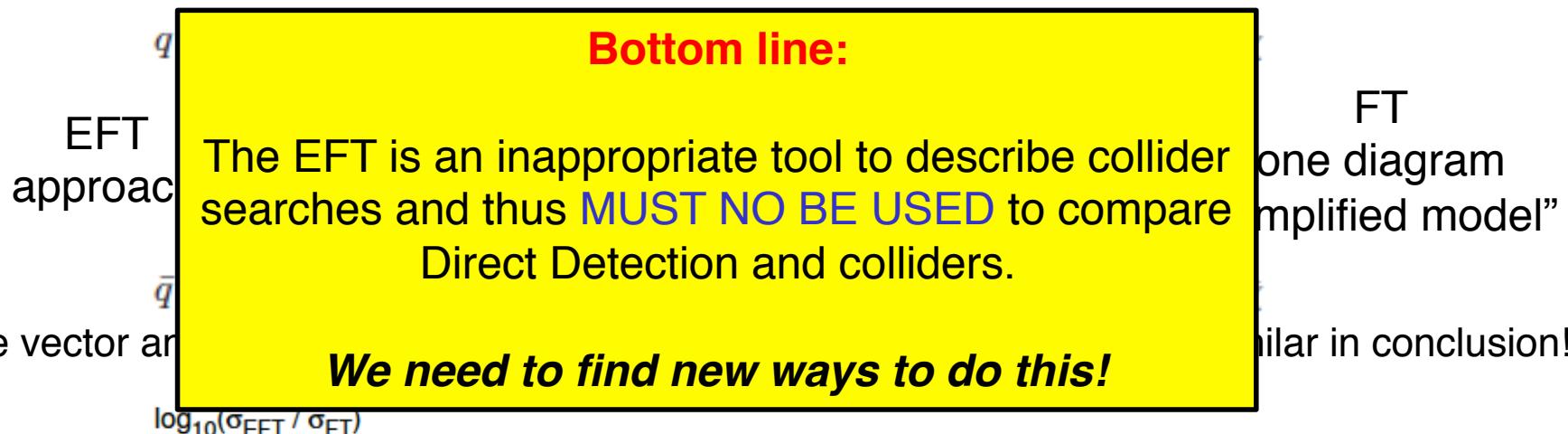
- Region I:** EFT and FT agree better than 20%
  - EFT is valid!
- Region II:** EFT yields significant weaker limits than FT
  - EFT limits are too conservative!
- Region III:** EFT yields significant stronger limits than FT
  - EFT limits are too aggressive!

# Validity of Effective Field Theory Limits

↳ collider & Direct Detection O. Buchmüller

Recent work from OB, M.Dolan,C.McCabe: arXiv:1308.6799

- Compare Effective Field Theory (EFT) with Full Theory (FT)

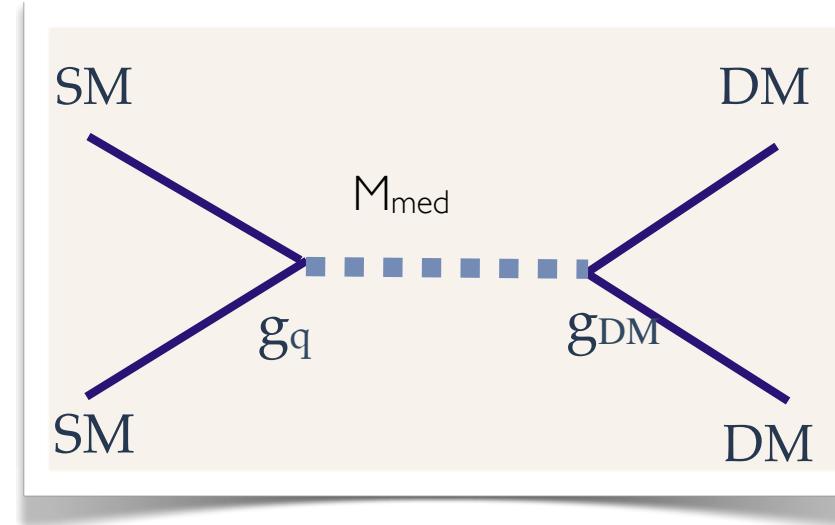


Compare prediction of FT with EFT in  $m_{\text{med}} - m_{\text{DM}}$  plane.  
Three regions become visible:

- **Region I:** EFT and FT agree better than 20%  
➤ EFT is valid!
- **Region II:** EFT yields significant weaker limits than FT  
➤ EFT limits are too conservative!
- **Region III:** EFT yields significant stronger limits than FT  
➤ EFT limits are too aggressive!

# Minimal Simplified Dark Matter Model

Based on work from :  
 OB, S. Malik,  
 M.Dolan,C.McCabe



s-channel

Define simplified model with  
(minimum) 4 parameters

Mediator mass ( $M_{\text{med}}$ )	DM mass ( $M_{\text{DM}}$ )
$g_q$	$g_{\text{DM}}$

DM

Dirac fermion	Scalar - real
Majorana fermion	Scalar - complex

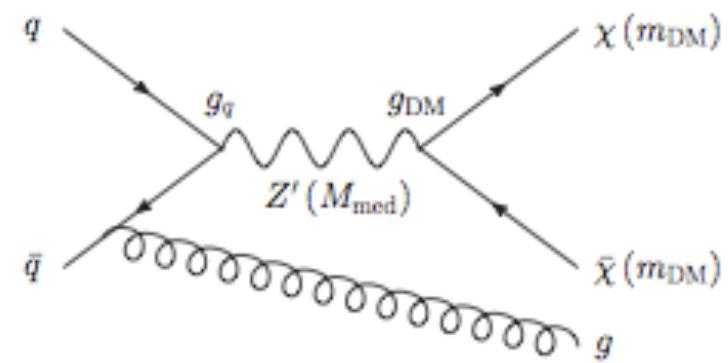
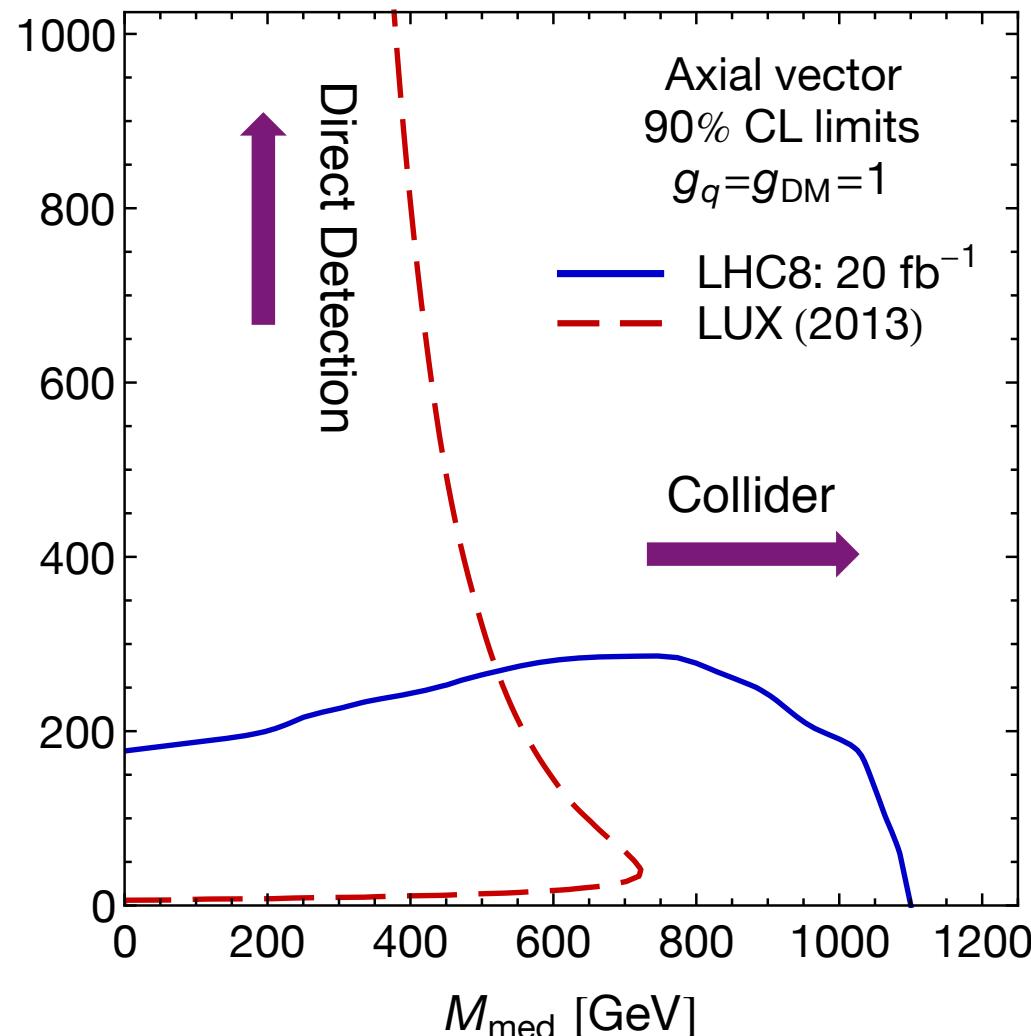
Consider comprehensive set  
of diagrams for mediator

Vector	Axial-vector
Scalar	Pseudoscalar

# Dark Matter Limits from Direct Searches: Today

DM Searches @ collider & Direct Detection O. Buchmüller

**Examples:** CMS monojet search  
and recent LUX result:  
interpretation in simplified models

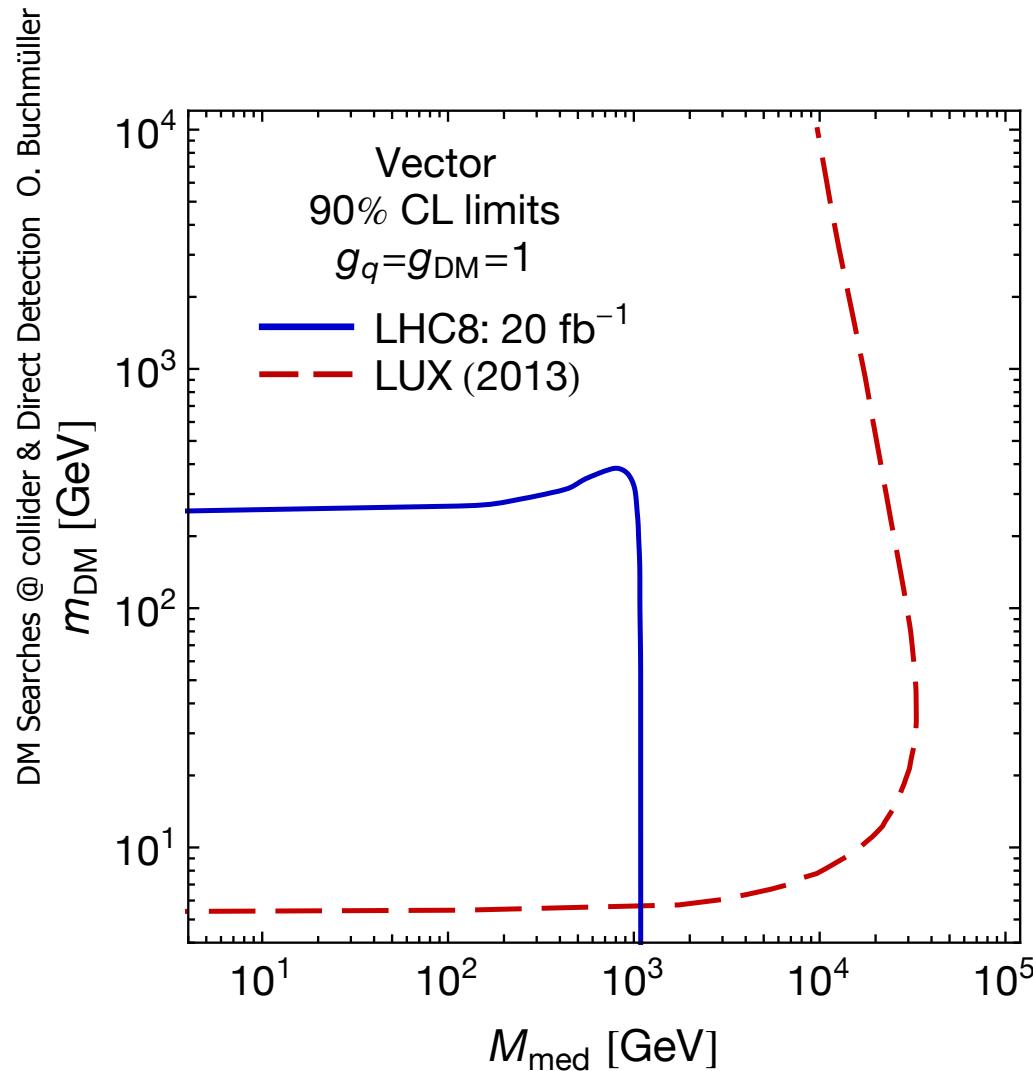


## Assumes:

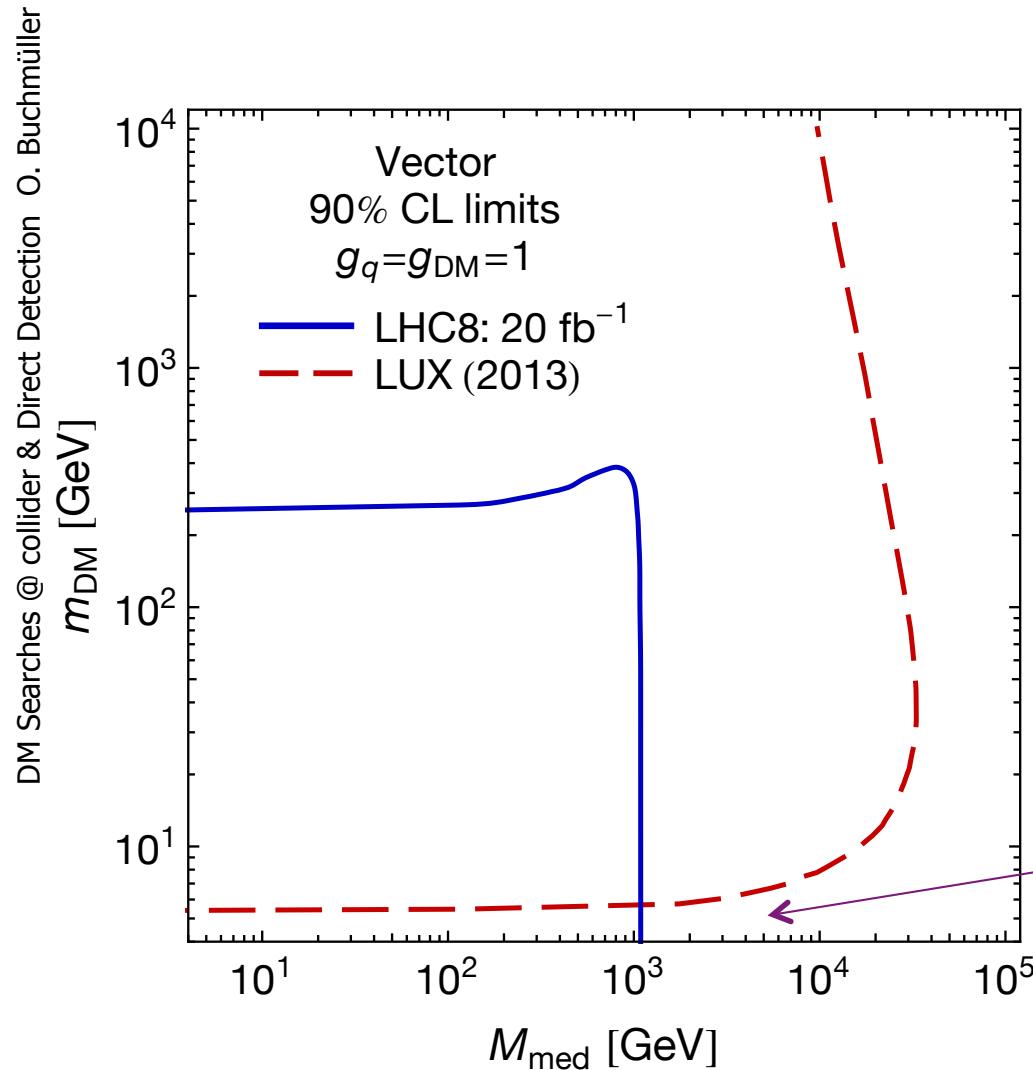
- couplings to all quarks
- $g_q=g_{\text{DM}}=1$
- width calculated from  $g$

**Direct Detection experiments  
and collider are complementary  
probing different regions of  
the relevant parameter space!**

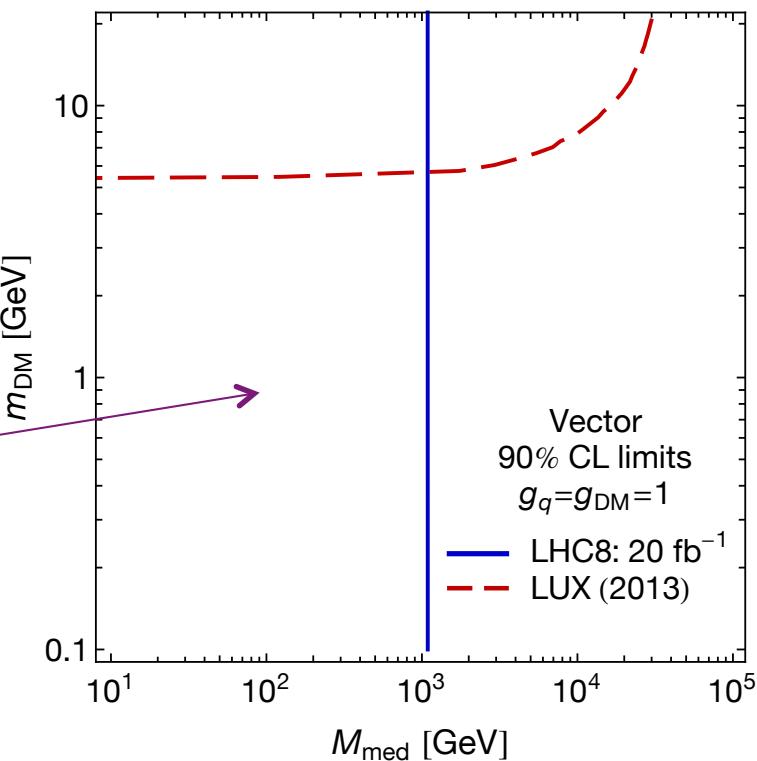
## The Vector Case



## The Vector Case

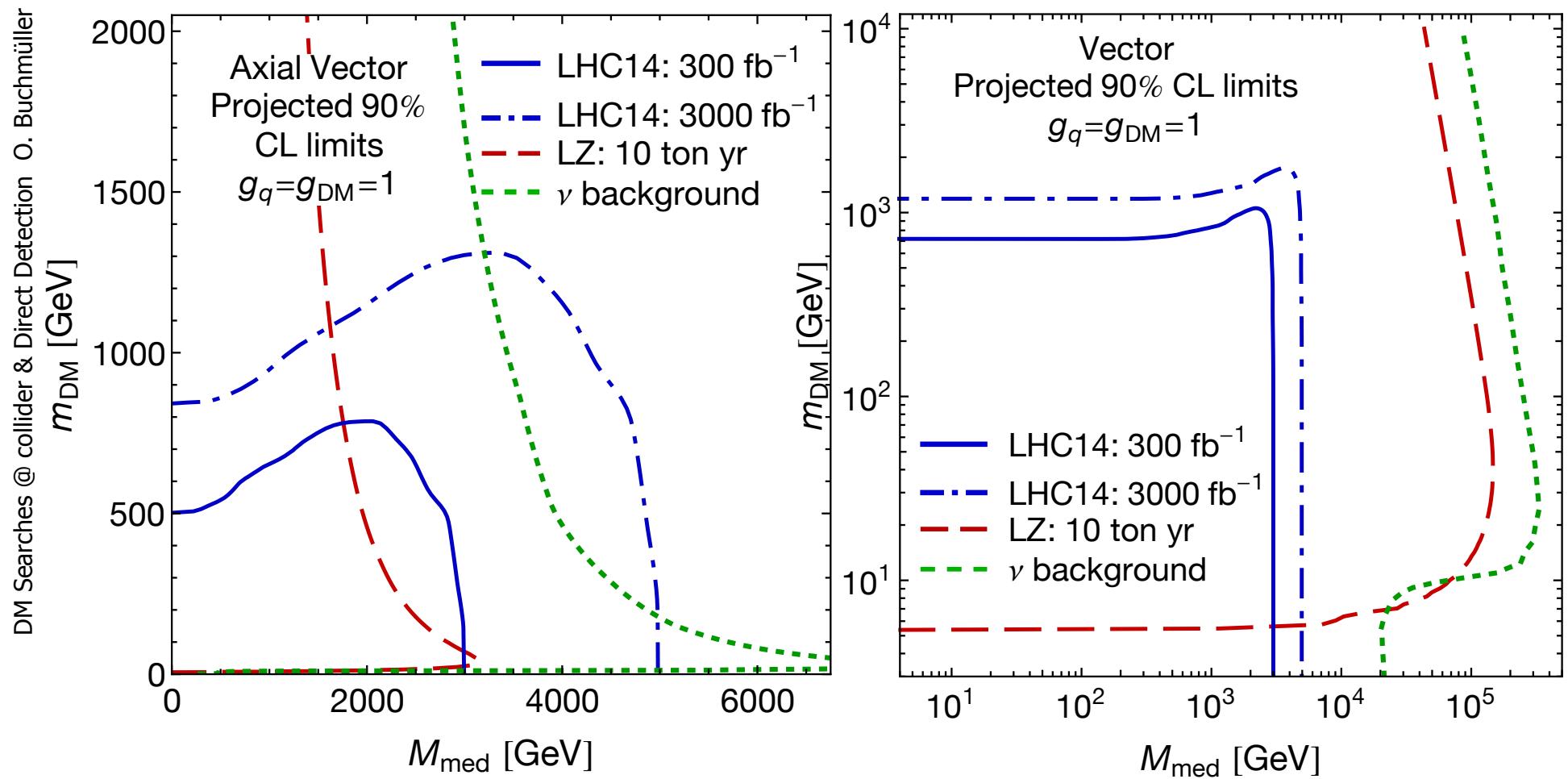


While for Axial-Vector collider and DD are complementary in the full parameter space, for the pure Vector case, DD limits are strong EXCEPT for low DM masses!



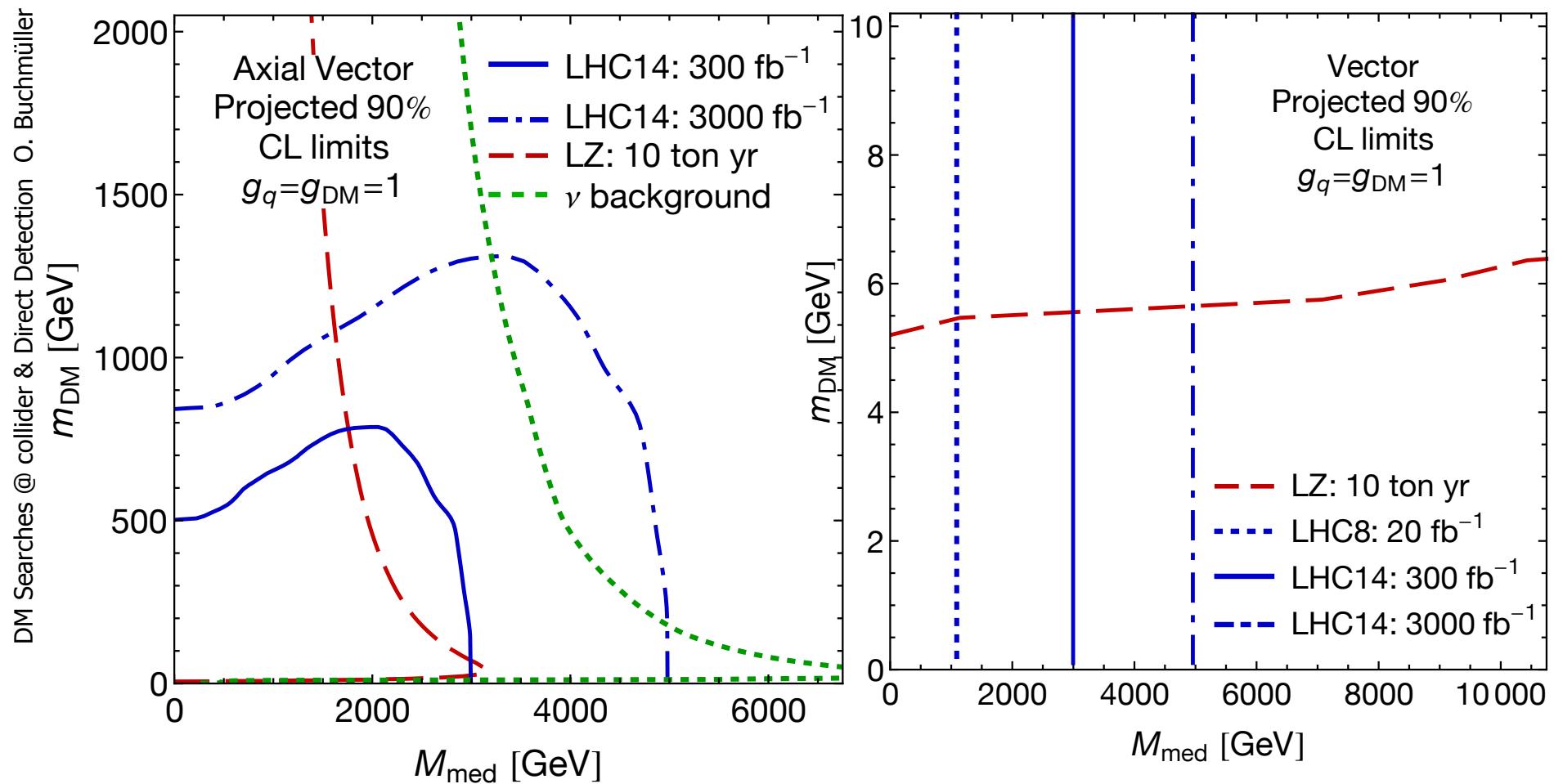
Note: Vector scales like Atomic number squared ( $\sim 130^{12}$ ) while AV scales like the spin (~1).

## Future Projections



**Compare:**  
**LHC @ 300/fb and HL-LHC@/3000/fb**  
**with Direct Detection:**  
**Lux-Zepline (~2025) and Neutrino noise border**

# Future Projections



**Important complementarity of the two experimental approaches will allow good coverage of the relevant parameter space!**

**Big discovery potential!**

## Summary

- So far the origin Dark Matter has not revealed itself!
- Both the LHC as well as Direct Detection experiments are probing very interesting regions in the parameter space!
- WE need to better work out the complementarity between collider and DD
  - Effective Field Theory is not the answer!
  - Simplified models might be helpful!
- The forthcoming data taking campaign of DD experiments as well as the high energy running of the LHC starting 2015 will be our next very (as in very) real chance for discovery!

The story continues ... stay tuned!

## Outlook: 8 TeV vs 14 TeV

Use parton luminosities to illustrate the gain of 14 vs 8 TeV

Higgs:

$pp \rightarrow H, H \rightarrow WW, ZZ$  and  $\gamma\gamma$

mainly gg: factor  $\sim 2$

SUSY – 3<sup>rd</sup> Generation:

Mass scale  $\sim 500$  GeV

qq and gg: factor  $\sim 3$  to 6

SUSY – Squarks/Gluino:

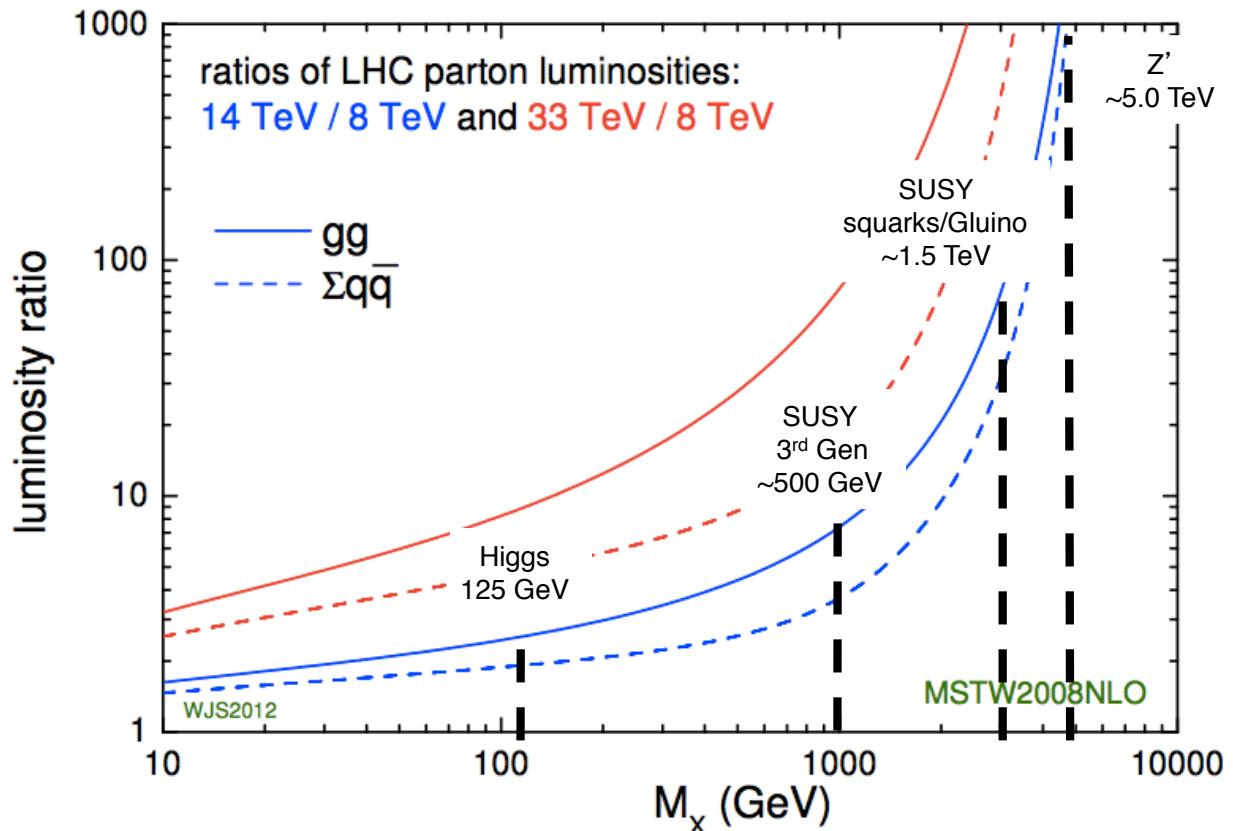
Mass scale  $\sim 1.5$  TeV

qq,gg,qg: factor  $\sim 40$  to 80

$Z'$ :

Mass scale  $\sim 5$  TeV

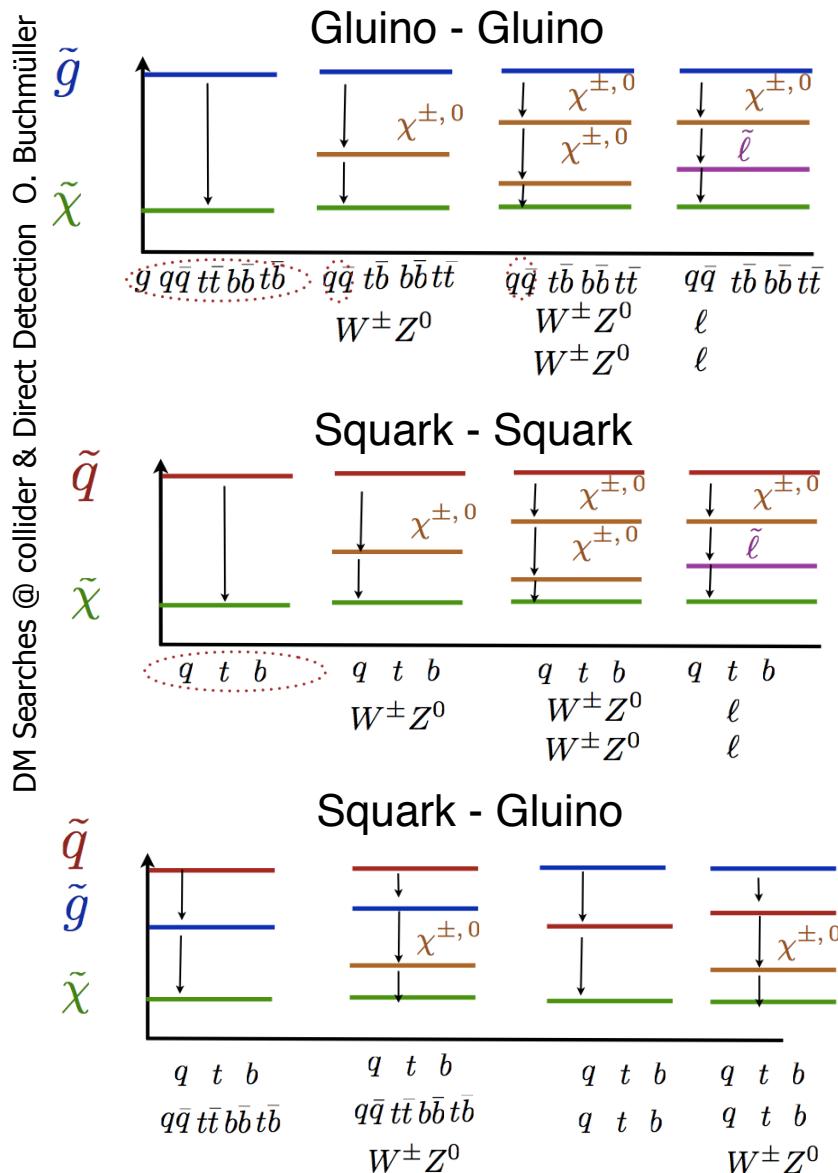
qq: factor  $\sim 1000$



Increase in energy will help a lot!  
Not just for SUSY...

# BACKUP

# Early SUSY Search Strategy at the LHC



## Search Signatures

- SUSY-like decay chains range from short to long and simple to very complicated.
- All physics objects, MET, jets, leptons, photons, b's taus, tops, W, Z, etc are involved
- Comprehensive coverage of all possible signature requires a topology oriented search strategy:

## References Analyses

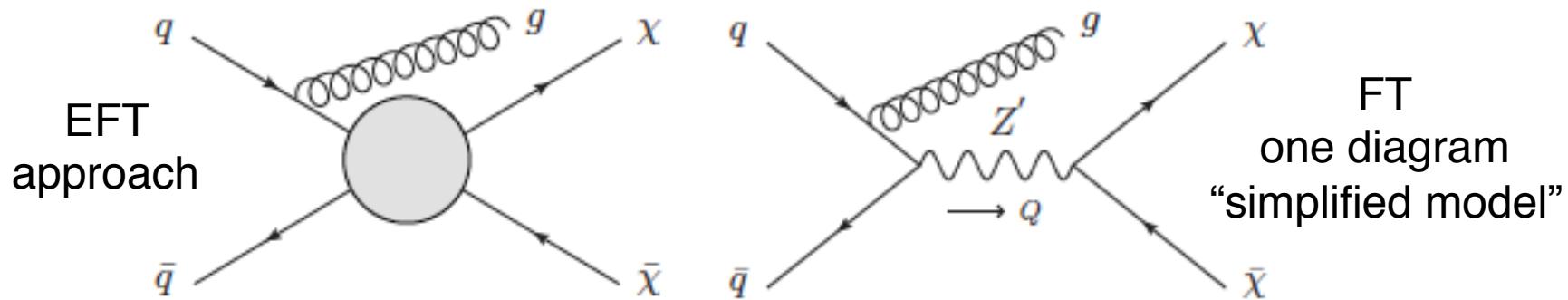
0-leptons	1-lepton	OSDL	SSDL	$\geq 3$ leptons	2-photons	$\gamma +$ lepton
Jets + MET	Single lepton + Jets + MET	Opposite-sign di-lepton + jets + MET	Same-sign di-lepton + jets + MET	Multi-lepton	Di-photon + jet + MET	Photon + lepton + MET

Already in less than two years of operation  
ATLAS & CMS managed to carry out  
the full list of these core  
“SUSY References Analyses”!

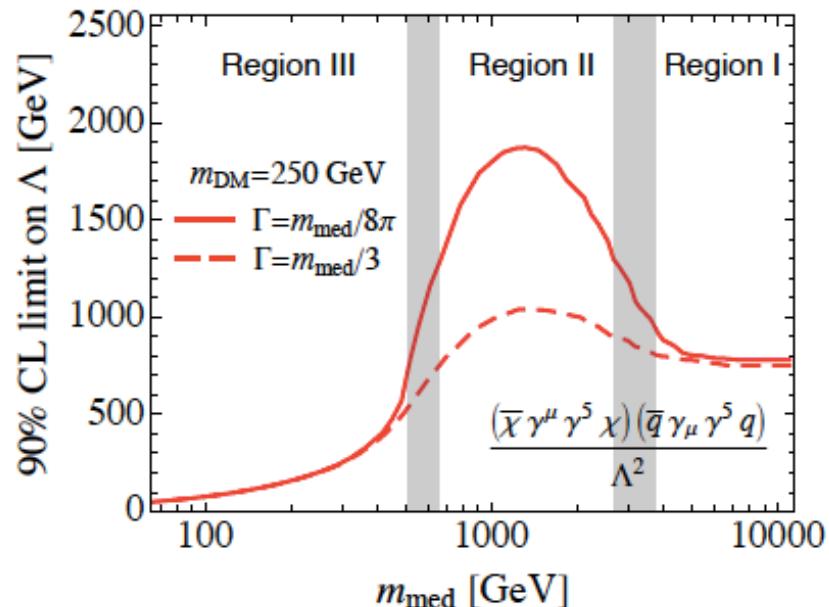
# Validity of Effective Field Theory Limits

Recent work from OB, M.Dolan,C.McCabe: arXiv:1308.6799

➤ Compare Effective Field Theory (EFT) with Full Theory (FT)



Use vector and axial-vector mediators (e.g.  $Z'$ ) as example - scalar are similar in conclusion!



Three Regions as function of mediator mass:

**Region I:** Heavy  $m_{\text{med}}$

➤ EFT is valid!

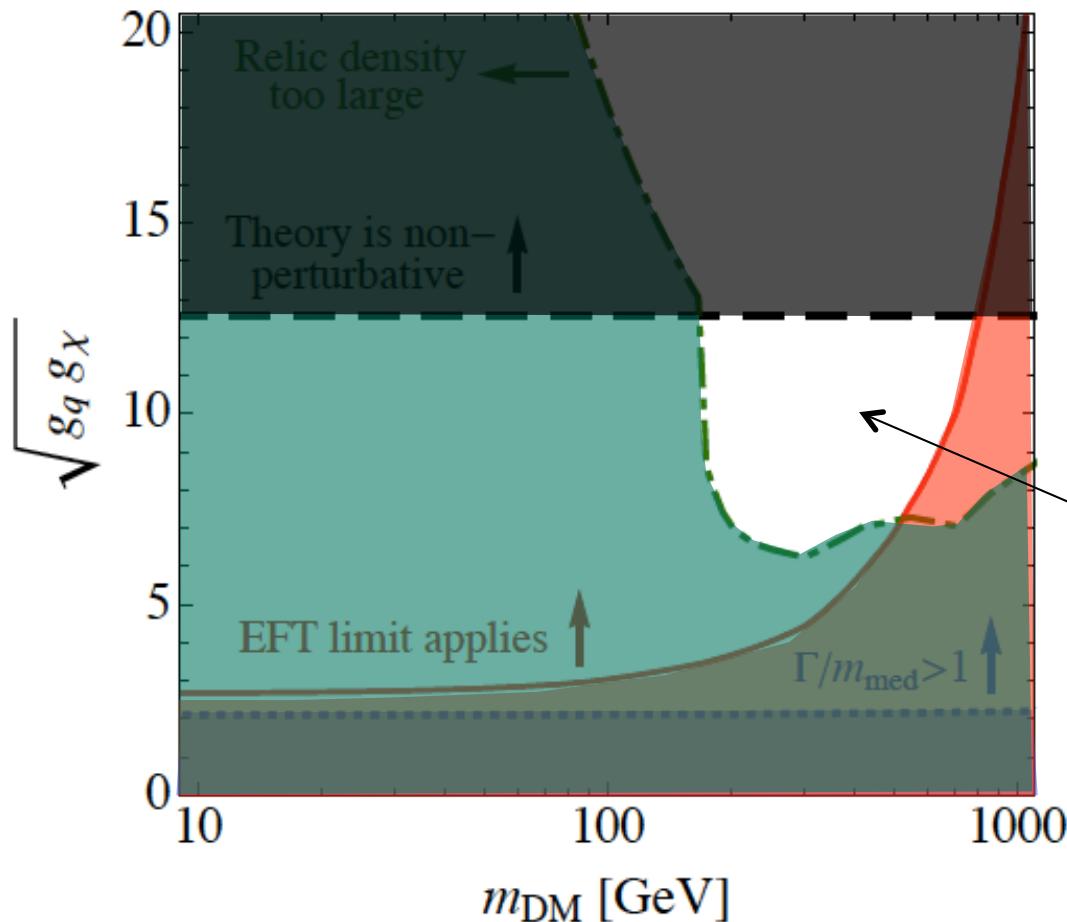
**Region II:** Medium  $m_{\text{med}}$  – Resonant enhancement

➤ EFT limits are too conservative!

**Region III:** Low  $m_{\text{med}}$

➤ EFT limits are too aggressive!

## What does this imply on model-dependencies of EFT limits?



Look at EFT validity in  $m_{\text{DM}}$  – coupling\* plane!

1. Must require  $m_{\text{med}} < \Gamma_{\text{med}}$
2. Region in which EFT is valid (20%)
3. Require compatibility with relic density
4. Require theory to be perturbative ( $< 4\pi$ )

When we also require that the region/theory must be perturbative:

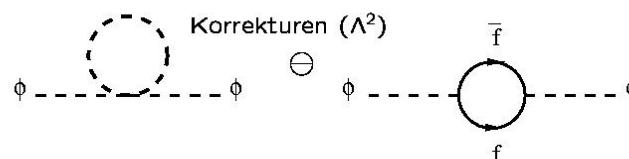
$$\sqrt{g_q g_\chi} < 4\pi$$

only a very small region is left!

This together with the observation that all DM theories for which the EFT is valid must have  $m_{\text{med}} < \Gamma_{\text{med}}$  leads to the conclusion the the EFT only applies to a very (as in VERY) small class of DM models.  
EFT limits of monojet searches are therefore highly model-dependent!

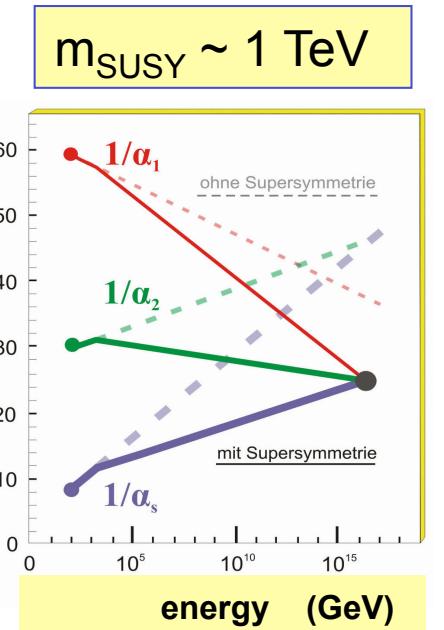
# Why is SUSY so attractive?

1. Quadratically divergent quantum corrections to the Higgs boson mass are avoided



(Hierarchy or naturalness problem)

2. Unification of coupling constants of the three interactions seems possible

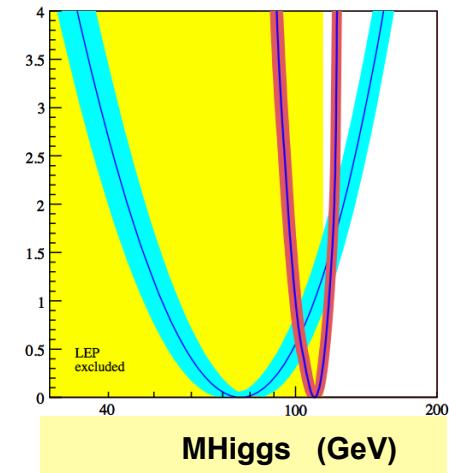


3. SUSY provides a candidate for dark matter,

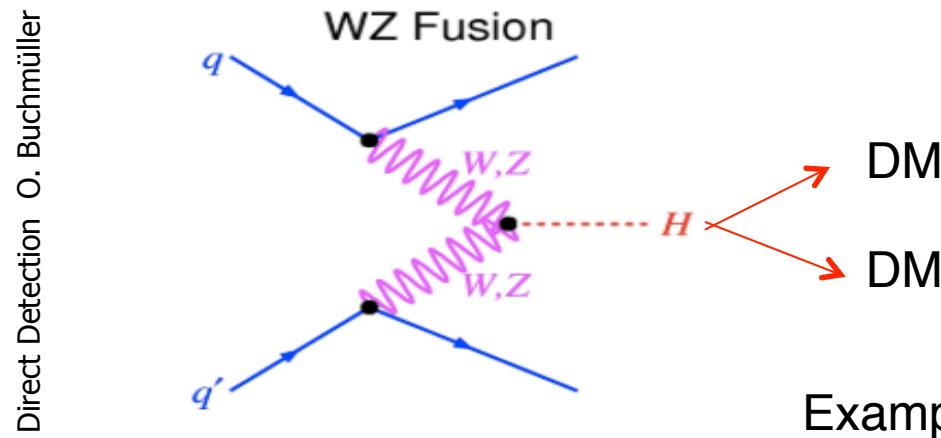


The lightest  
SUSY particle  
(LSP)

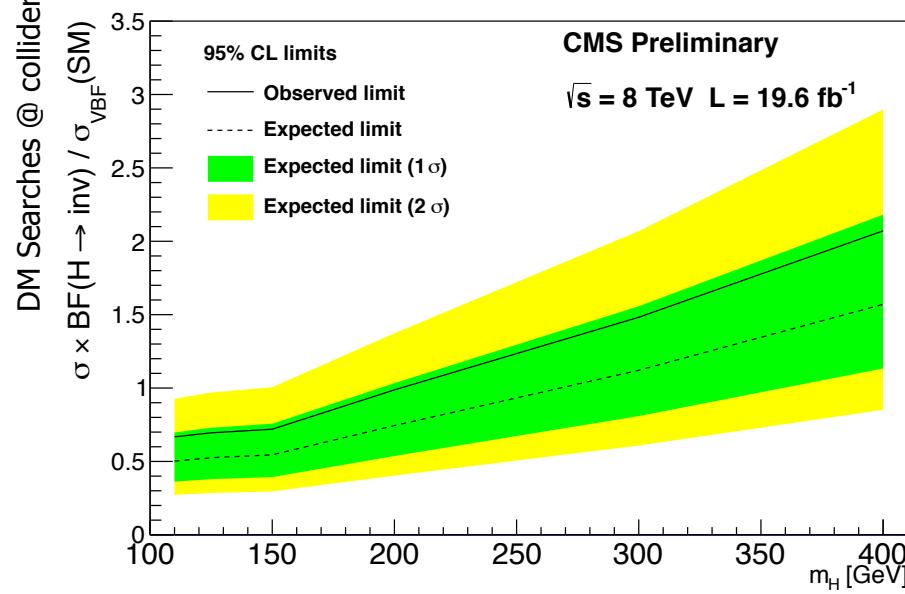
4. A SUSY extension is a small perturbation, consistent with the electroweak precision data



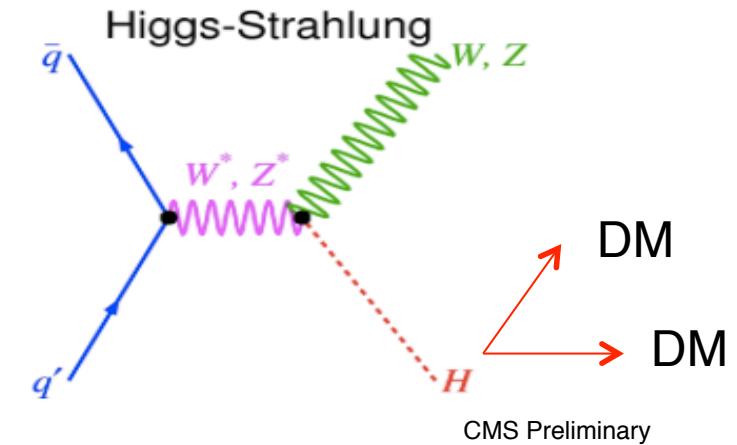
# Dark Matter from invisible Higgs searches



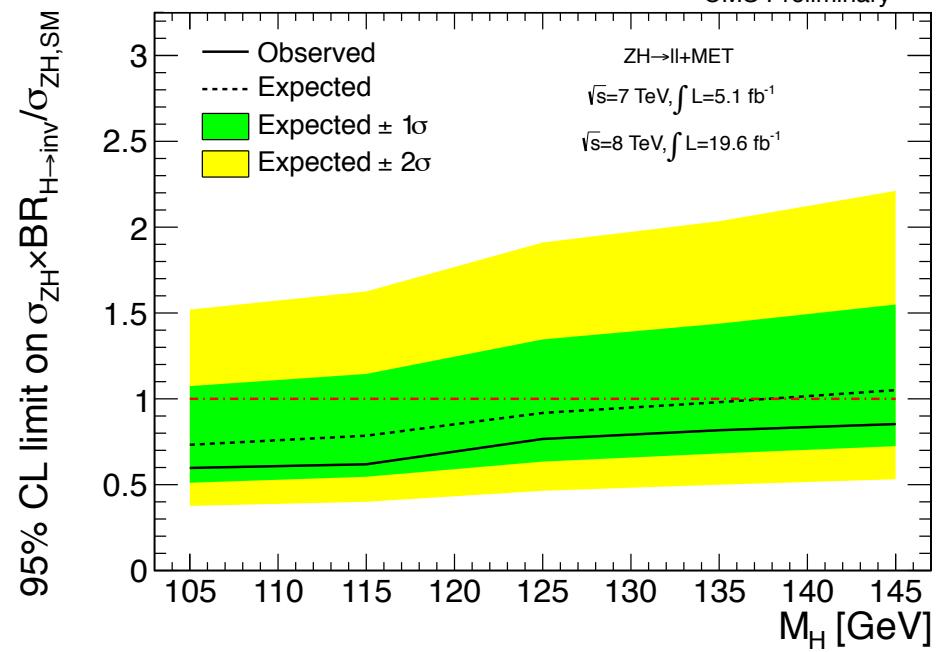
Example CMS



$\text{BR}(H \rightarrow \text{invisible}) < 68\% @ 95\%\text{CL}$



CMS Preliminary



$\text{BR}(H \rightarrow \text{invisible}) < 75\% @ 95\%\text{CL}$

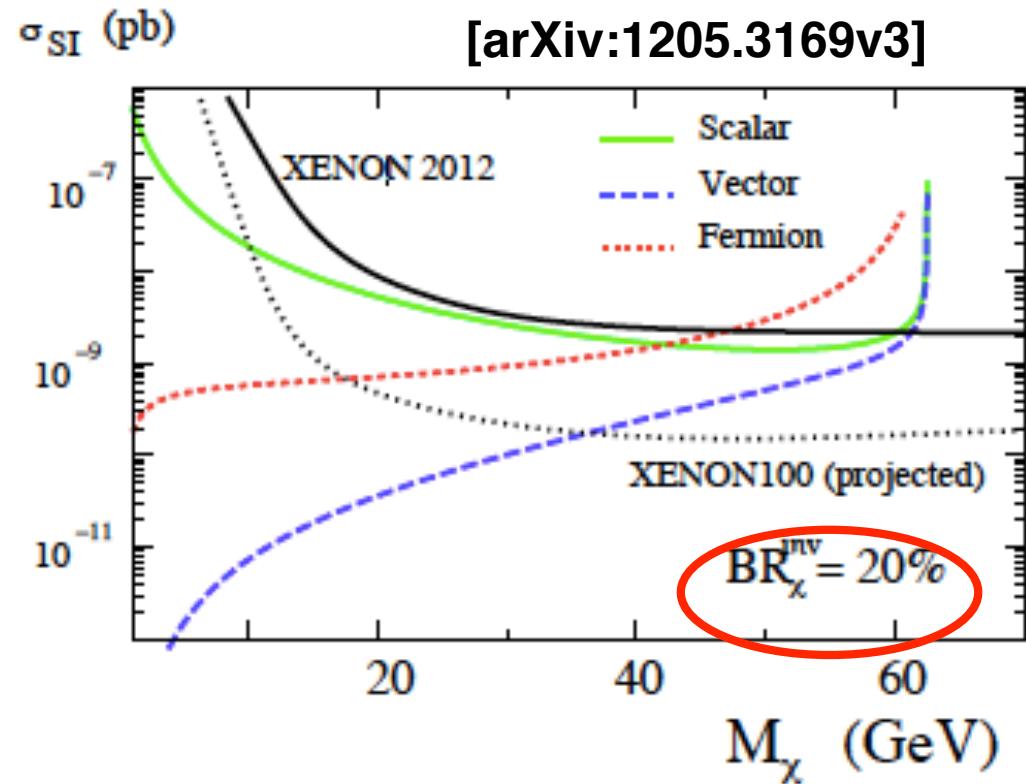
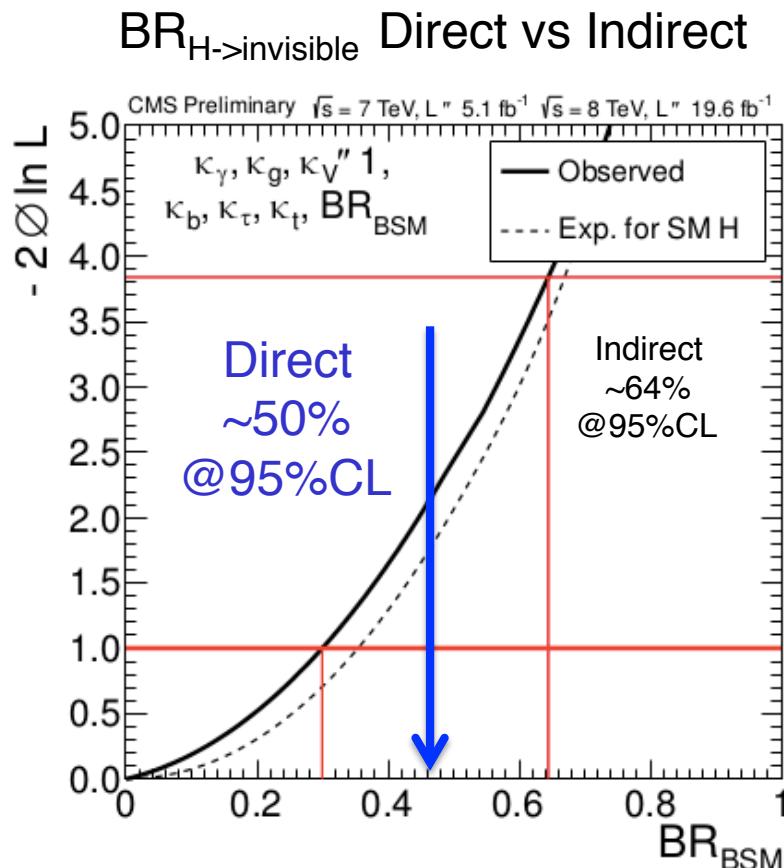
# Dark Matter from invisible Higgs searches

## Status 2012 CMS only:

VBF:  $\text{BR}_{H \rightarrow \text{invisible}} < 68\% @ 95\%\text{CL}$

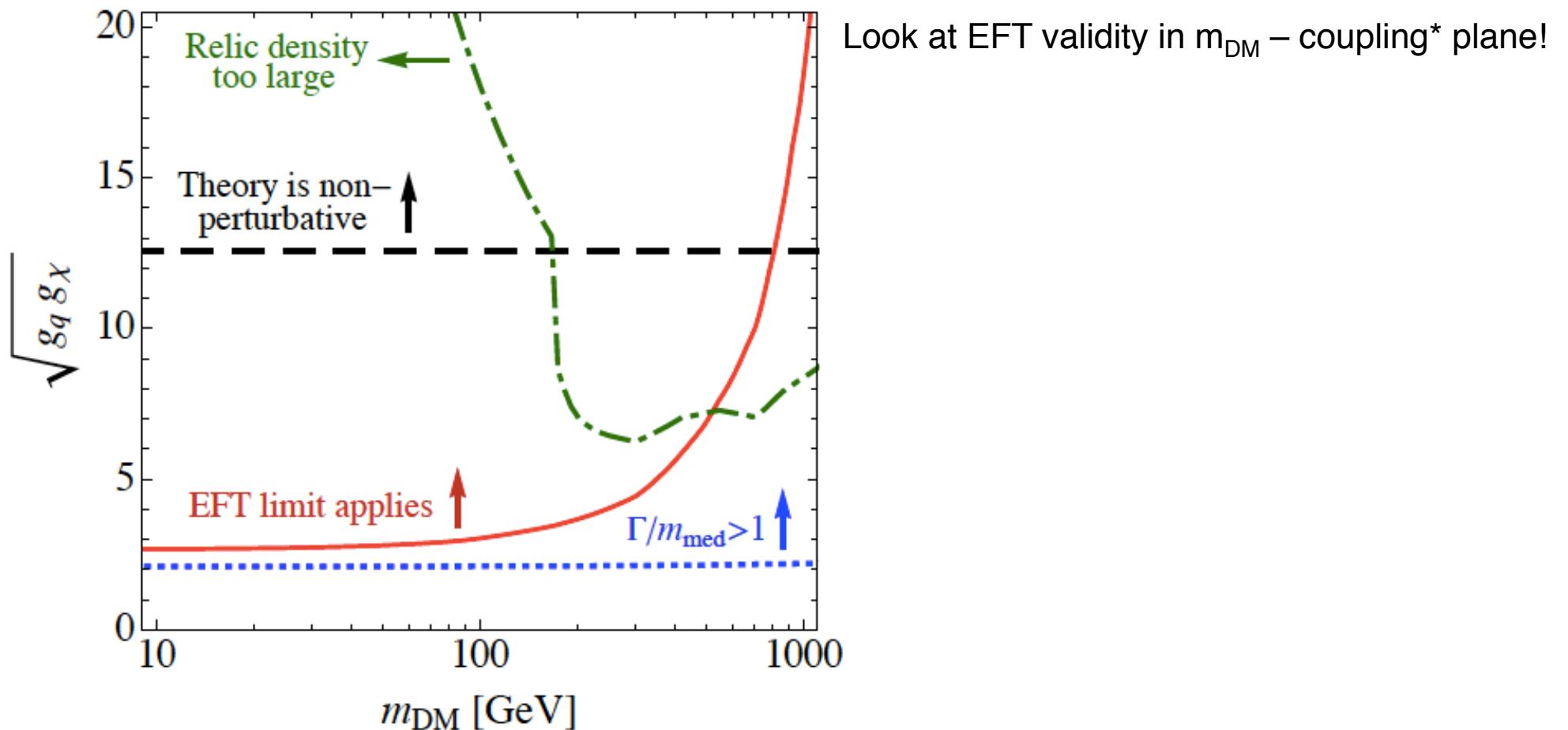
VH:  $\text{BR}_{H \rightarrow \text{invisible}} < 75\% @ 95\%\text{CL}$

Naïve combination:  $\sim 50\% @ 95\%\text{ CL}$



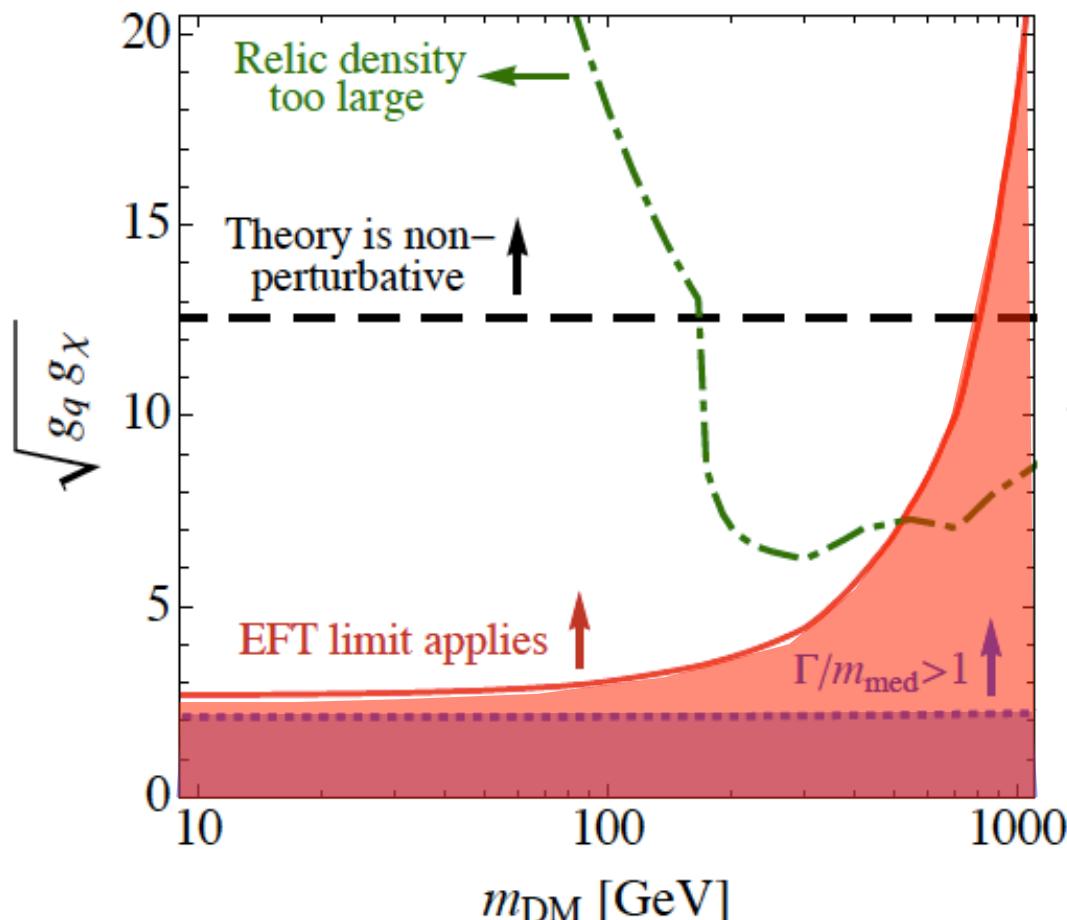
Assuming the experiments are able to maintain trigger and analysis acceptances, the LHC will provide a VERY powerful comparison of indirect & direct measurement of  $\Gamma_{H \rightarrow \text{invisible}}$ . In the (near) future this might provide a stringent constraint for  $M_{\text{DM}} < M_H/2$

## What does this imply on model-dependences of EFT limits?



\* Coupling chose such that CMS EFT limit on  $\Lambda$  applies to FT

## Model-dependences of EFT limits



Look at EFT validity in  $m_{\text{DM}} - \text{coupling}^*$  plane!

### 1. Region in which EFT is valid

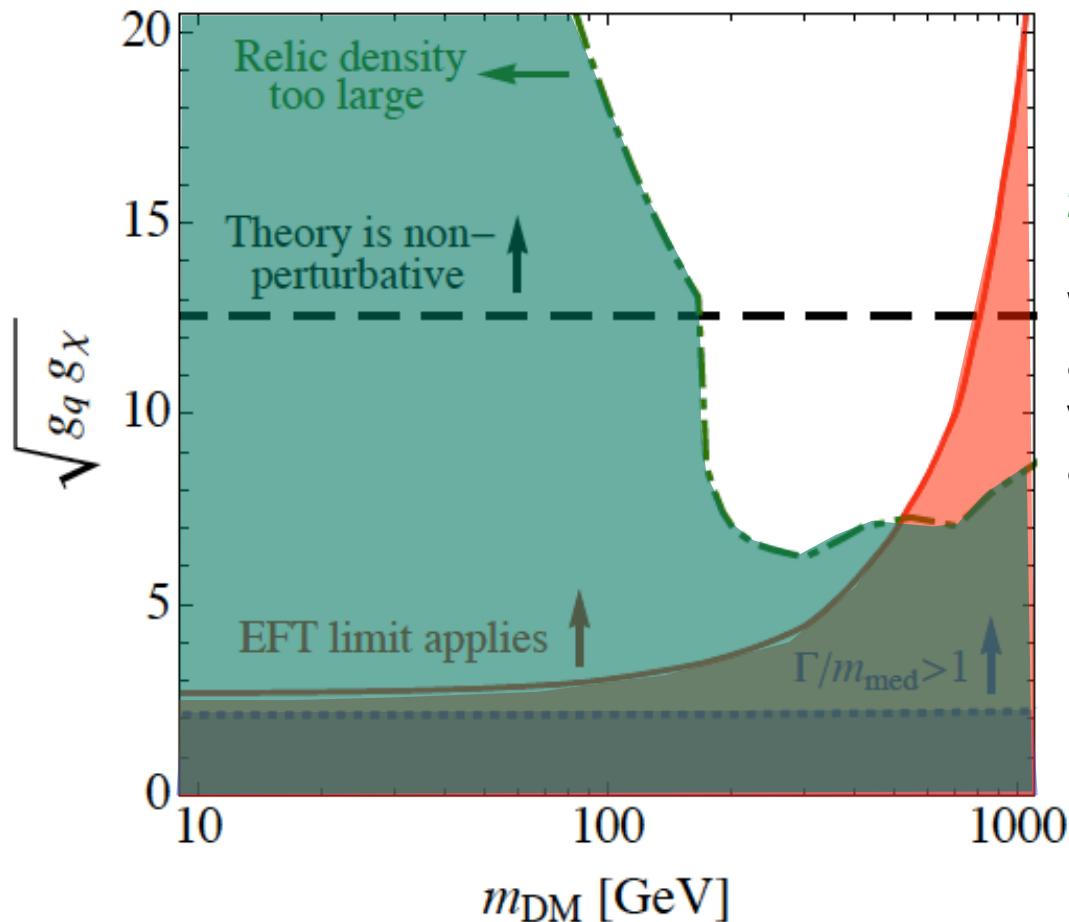
For this we calculate the minimum coupling

$$\sqrt{g_q g_\chi} = m_{\text{med}}/\Lambda_{\text{CMS}}$$

that the simplified model must have for the EFT limits to apply. This is defined by region I (i.e. better than 20% agreement of FT and EFT).

\* Coupling chose such that CMS EFT limit on  $\Lambda$  applies to FT

## Model-dependences of EFT limits



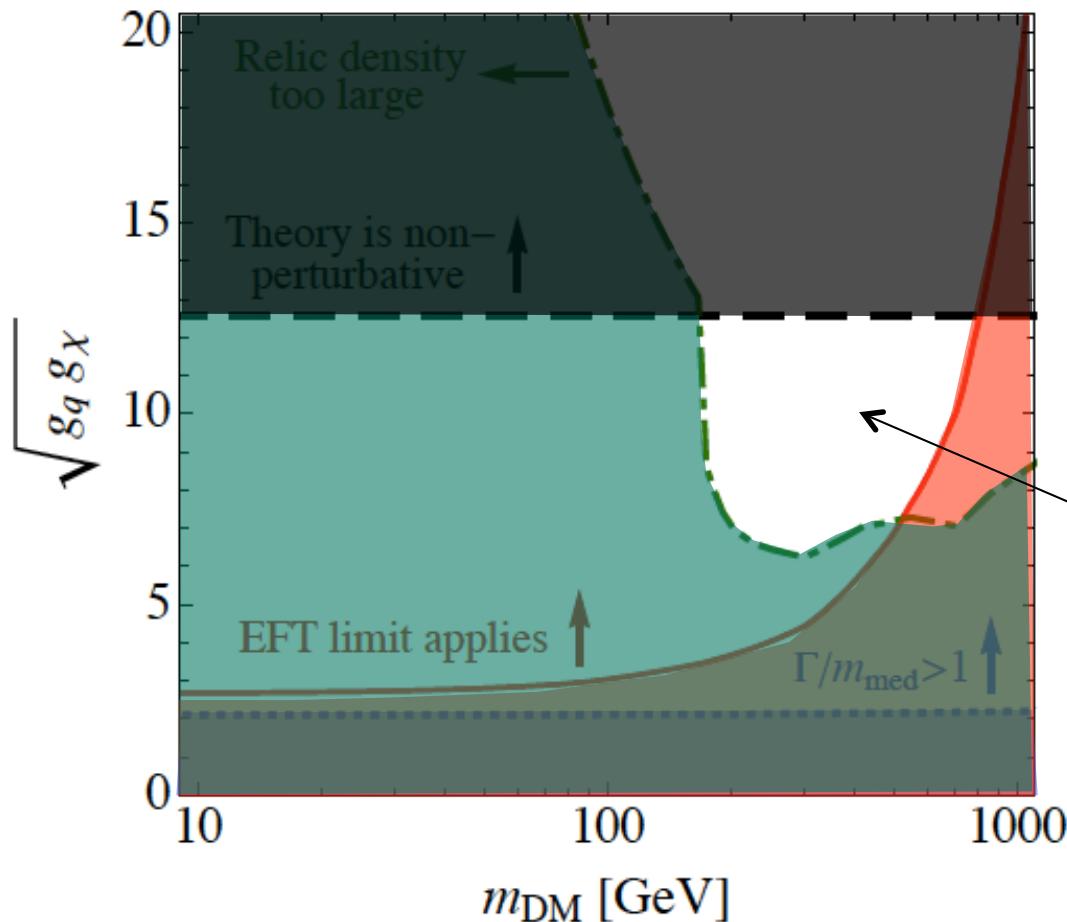
Look at EFT validity in  $m_{\text{DM}}$  – coupling\* plane!

1. Region in which EFT is valid (20%)
2. Require compatibility with relic density

When exclude the region in which relic abundance is larger than the observed value of  $\Omega_{\chi\chi} h^2 = 0.119$  only mediator masses above a few hundred GeV fulfill this.

\* Coupling chose such that CMS EFT limit on  $\Lambda$  applies to FT

## Model-dependences of EFT limits



Look at EFT validity in  $m_{\text{DM}} - \text{coupling}^*$  plane!

1. Must require  $m_{\text{med}} < \Gamma_{\text{med}}$
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4. Require theory to be perturbative ( $< 4\pi$ )

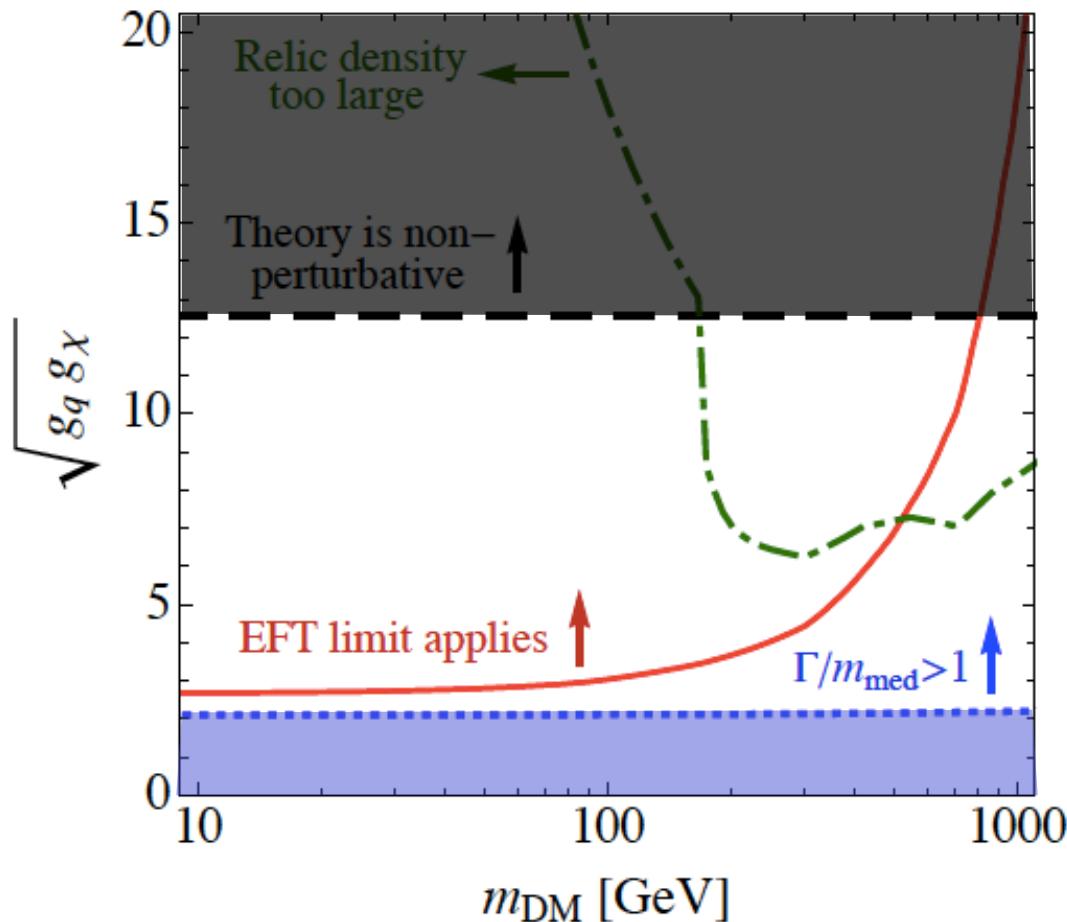
When we also require that the region/theory must be perturbative:

$$\sqrt{g_q g_\chi} < 4\pi$$

only a very small region is left!

EFT limits of monojet searches only apply to a very (as in VERY) small class of DM models!

## Model-dependences of EFT limits



Look at EFT validity in  $m_{\text{DM}} - \text{coupling}^*$  plane!

1. Region in which EFT is valid (20%)
2. Require compatibility with relic density
3. Require theory to be perturbative ( $< 4\pi$ )
4.  $m_{\text{med}} < \Gamma_{\text{med}}$  ALWAYS!

We also find that for all DM models the EFT is valid the mass of the mediator must be smaller than its width!

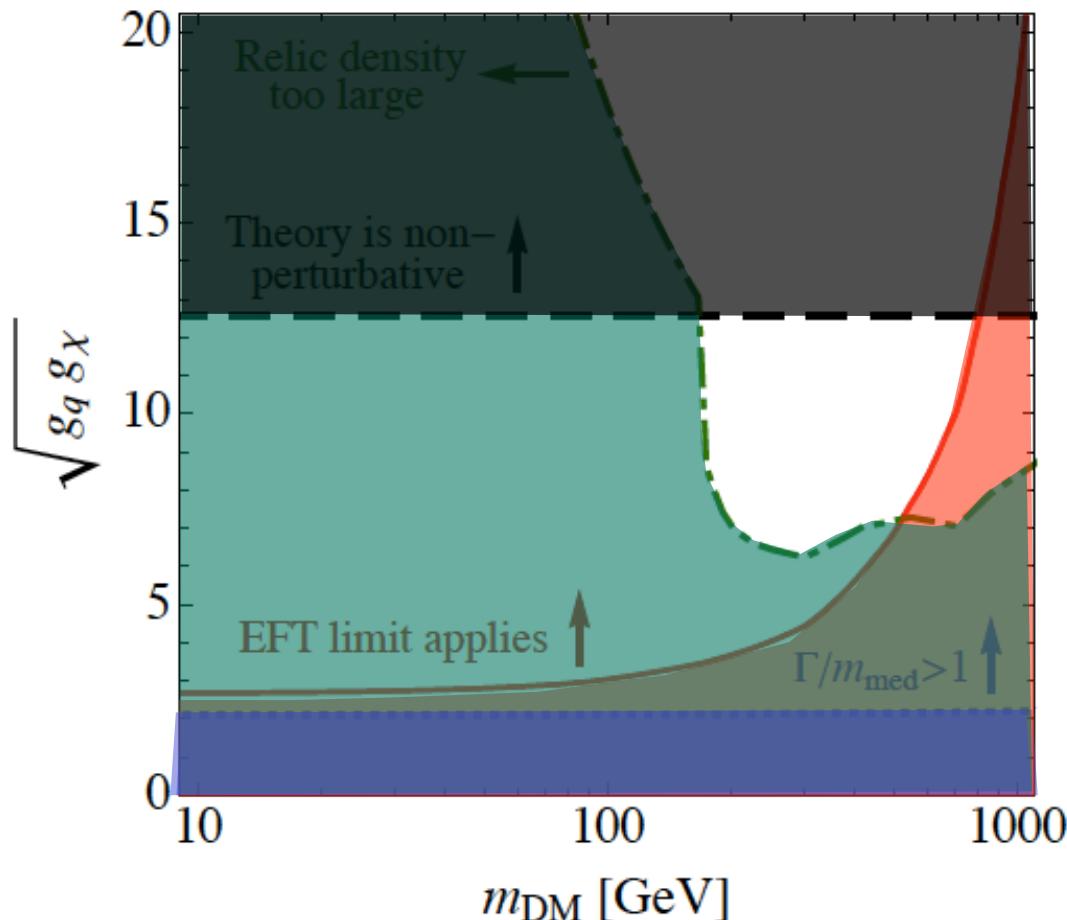
In the remaining part of the plot:

$$\sqrt{g_q g_\chi} > 2$$

a particle-like interpretation of the mediator is doubtful because of  $m_{\text{med}} < \Gamma_{\text{med}}$  !

See discussion about equation 3.5 in arXiv:1308.6799 for further details.

## What does this imply on model-dependences of EFT limits?



Look at EFT validity in  $m_{\text{DM}} - \text{coupling}^*$  plane!

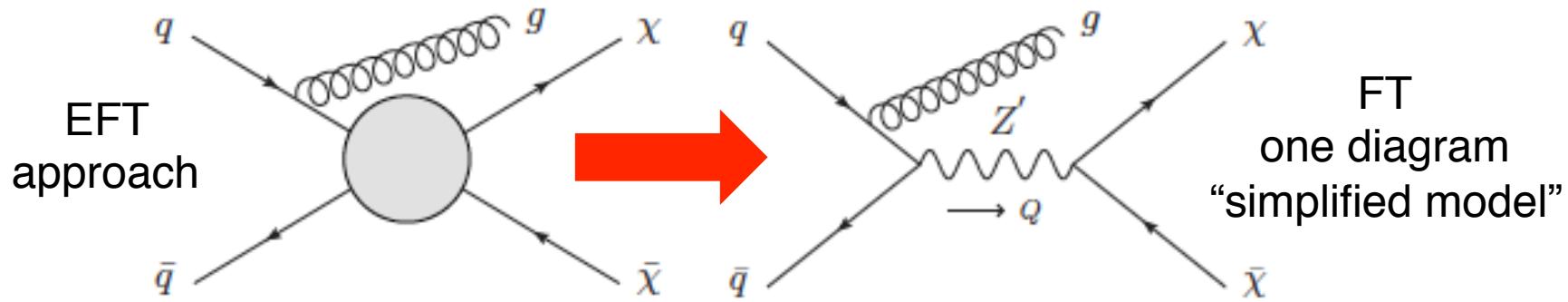
1. Region in which EFT is valid (20%)
2. Require compatibility with relic density
3. Require theory to be perturbative ( $< 4\pi$ )
4.  $m_{\text{med}} < \Gamma_{\text{med}}$  ALWAYS!

The observation that all DM theories for which the EFT is valid must have  $m_{\text{med}} < \Gamma_{\text{med}}$  and the small class to models it applies in any case leads to the conclusion the EFT only applies to a very small class of DM models.  
EFT limits of monojet searches are therefore highly model-dependent!

## Alternative Interpretation Ansatz: Simplified models

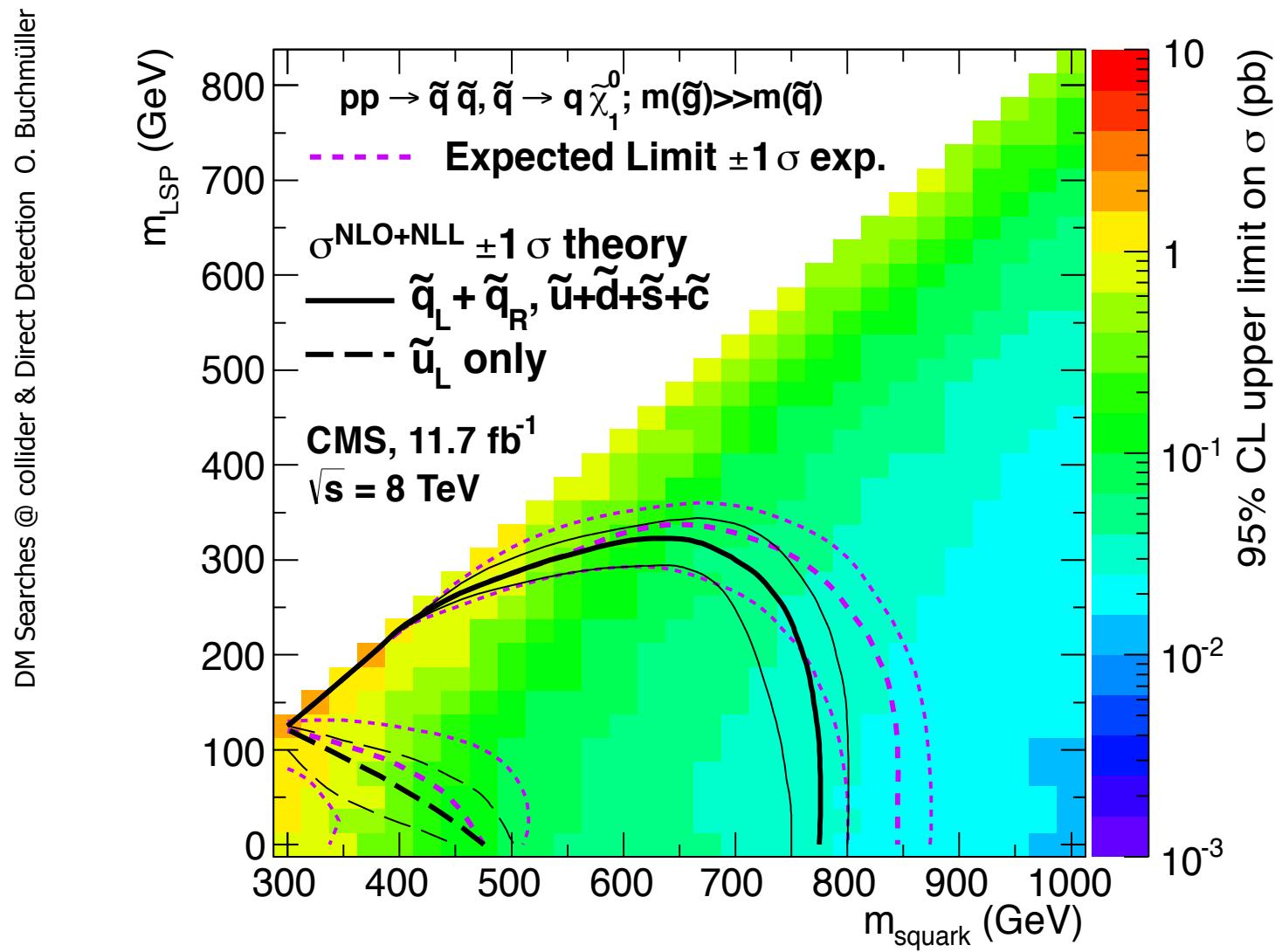
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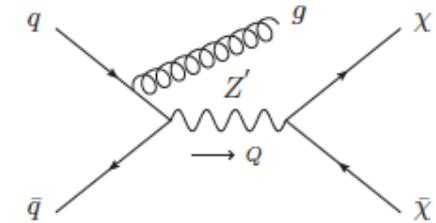
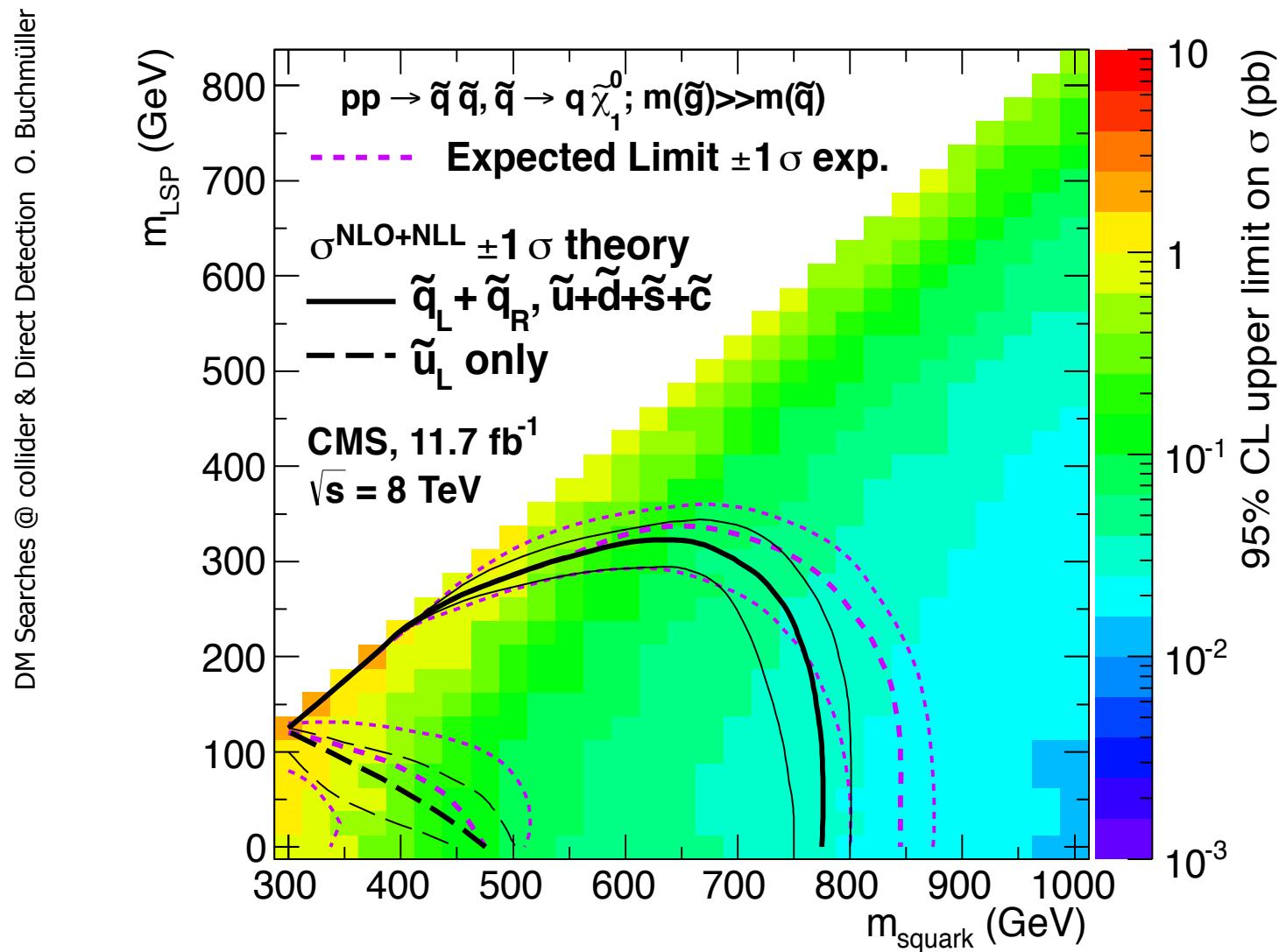


After three years of operation at the LHC the landscape for interpretation of searches has changed dramatically – new superior & modern approaches have replaced in many areas longstanding traditional ones (e.g. SUSY searches)

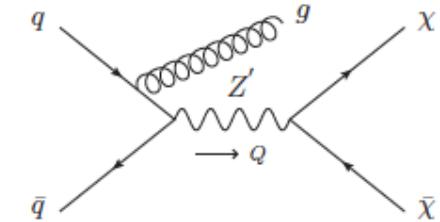
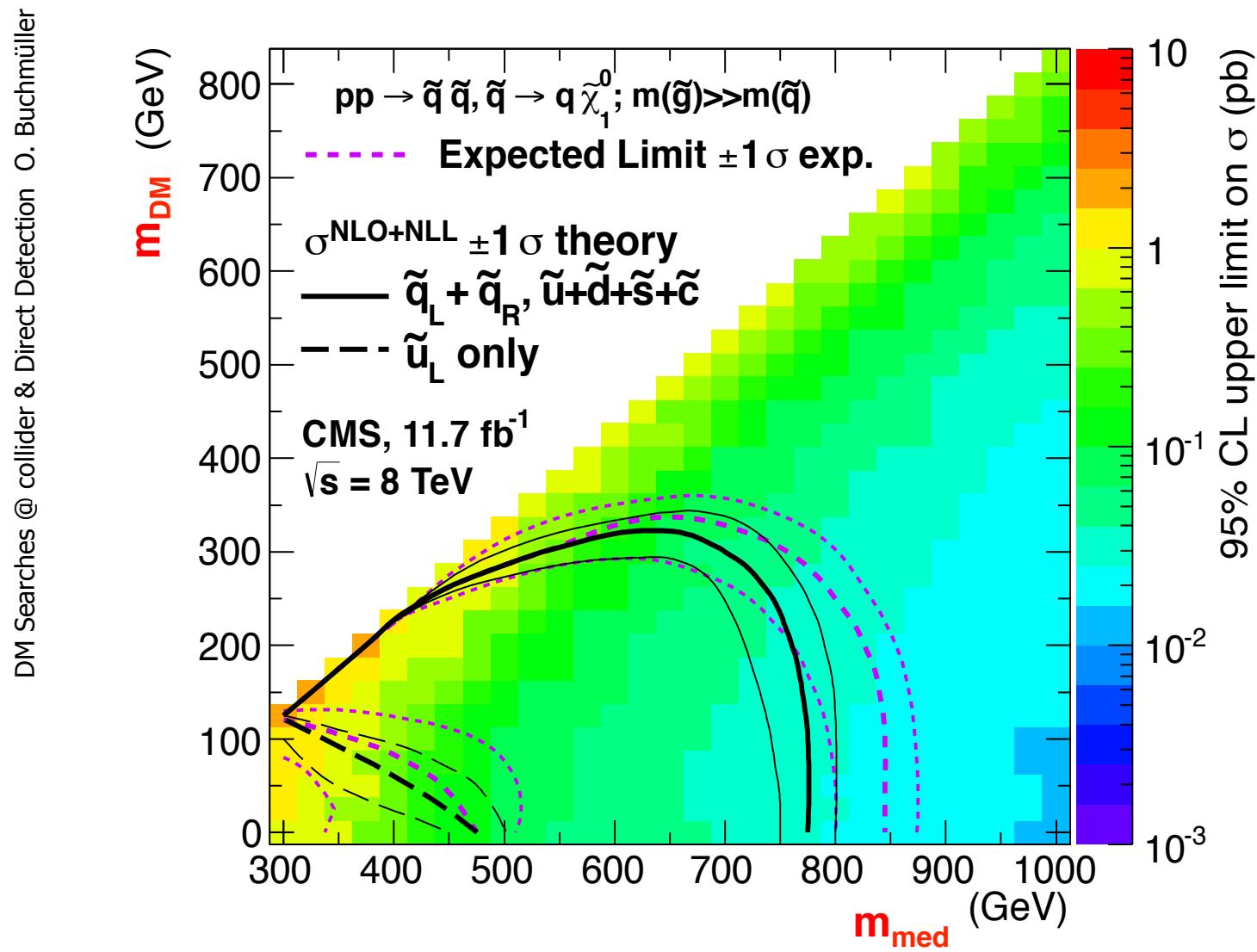
# The proposal



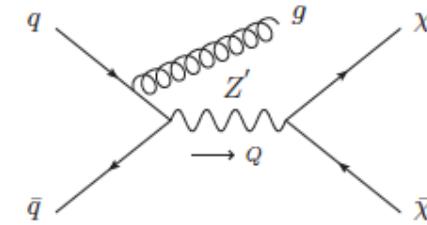
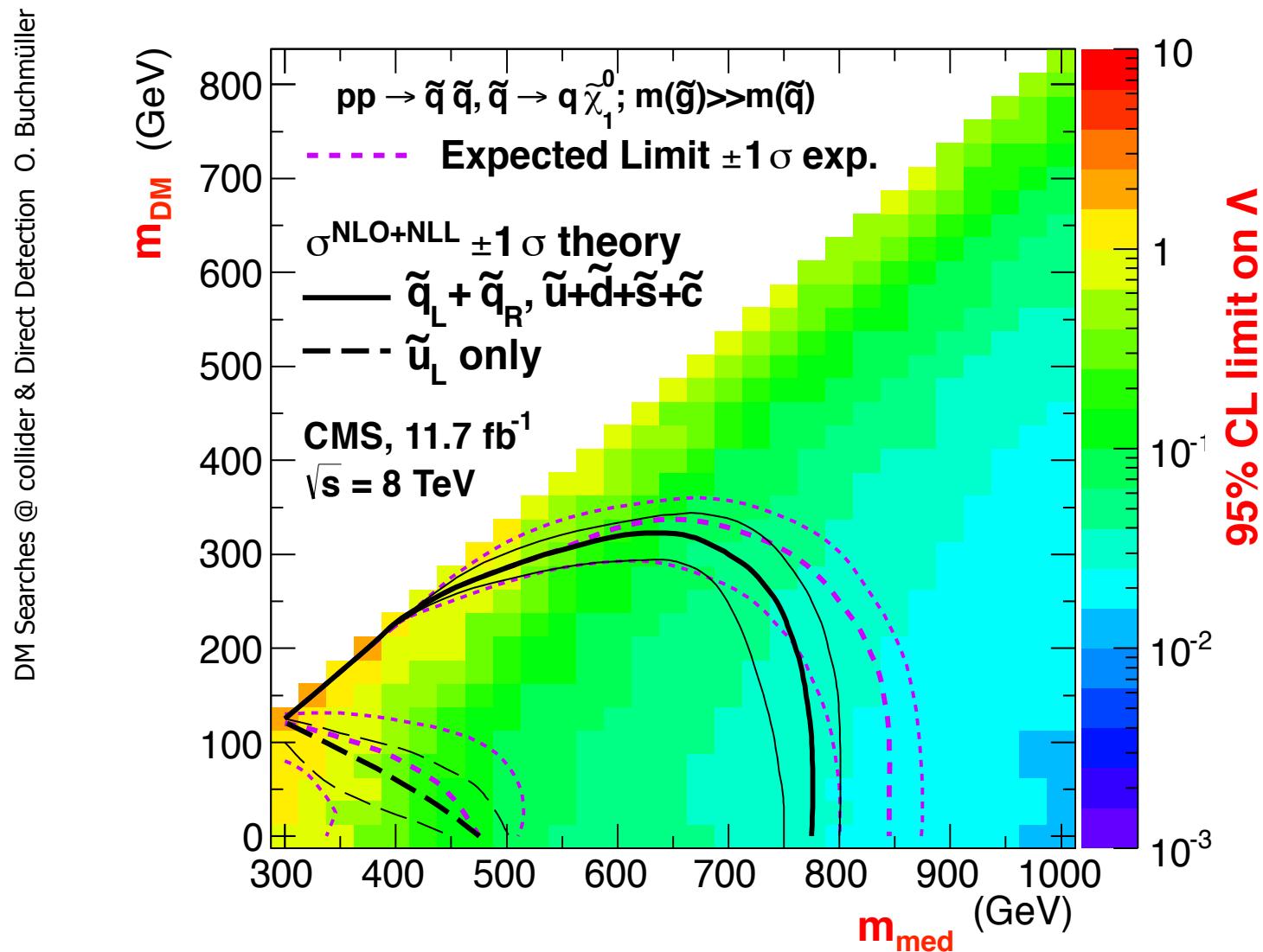
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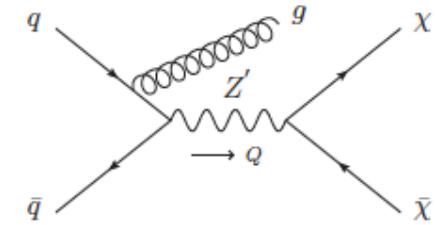
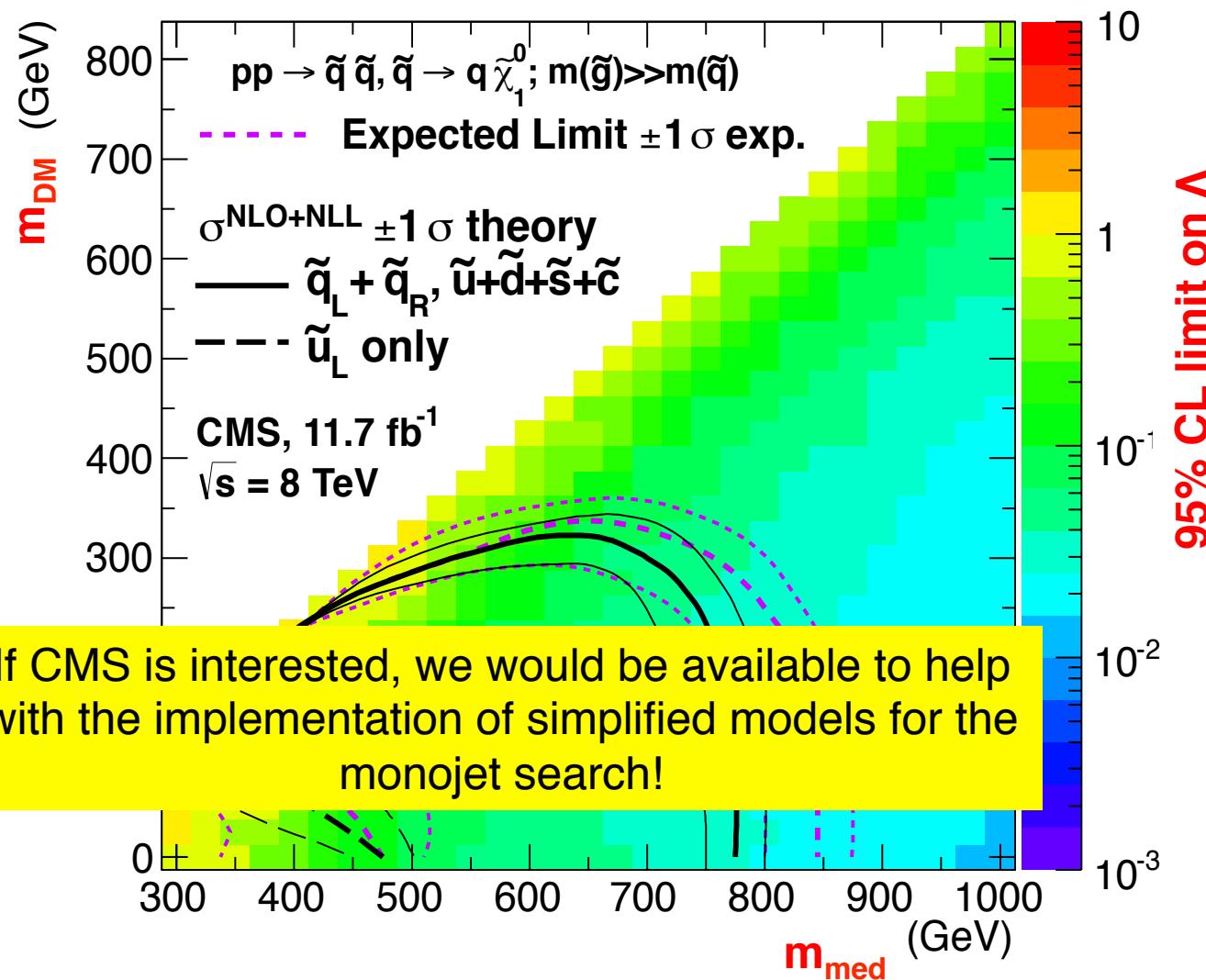
# The proposal



Very similar to limits in SUSY simplified models –  
 $m_{\text{med}}, m_{\text{DM}}, \Lambda$   
and possibly some variation of  $\Gamma_{\text{med}}$   
will cover the full problem!

# The proposal

DM Searches @ collider & Direct Detection O. Buchmüller

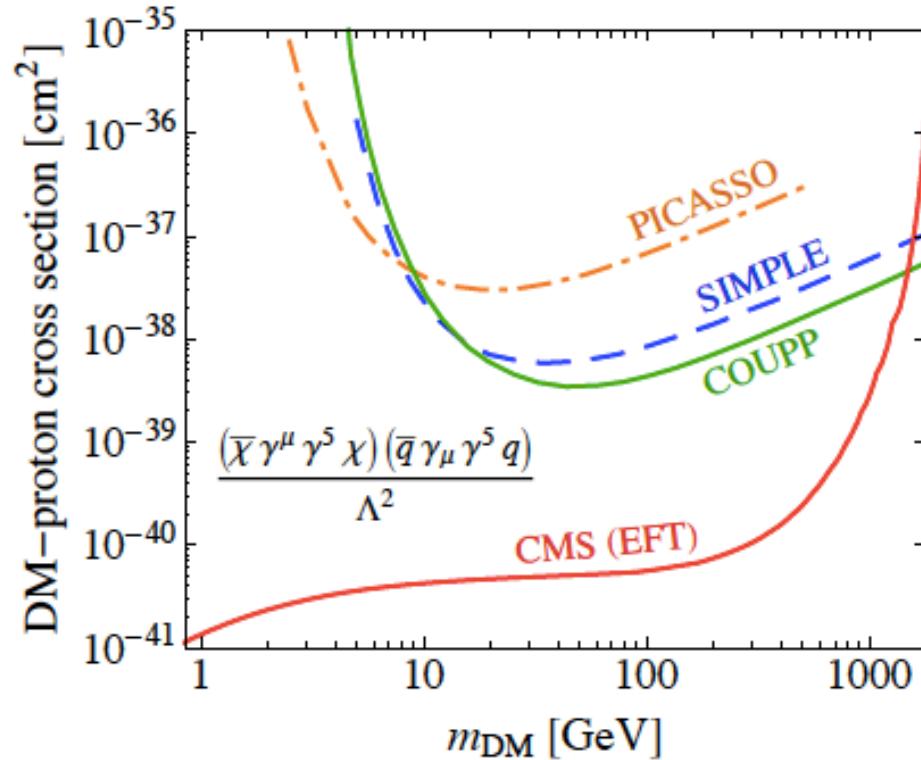


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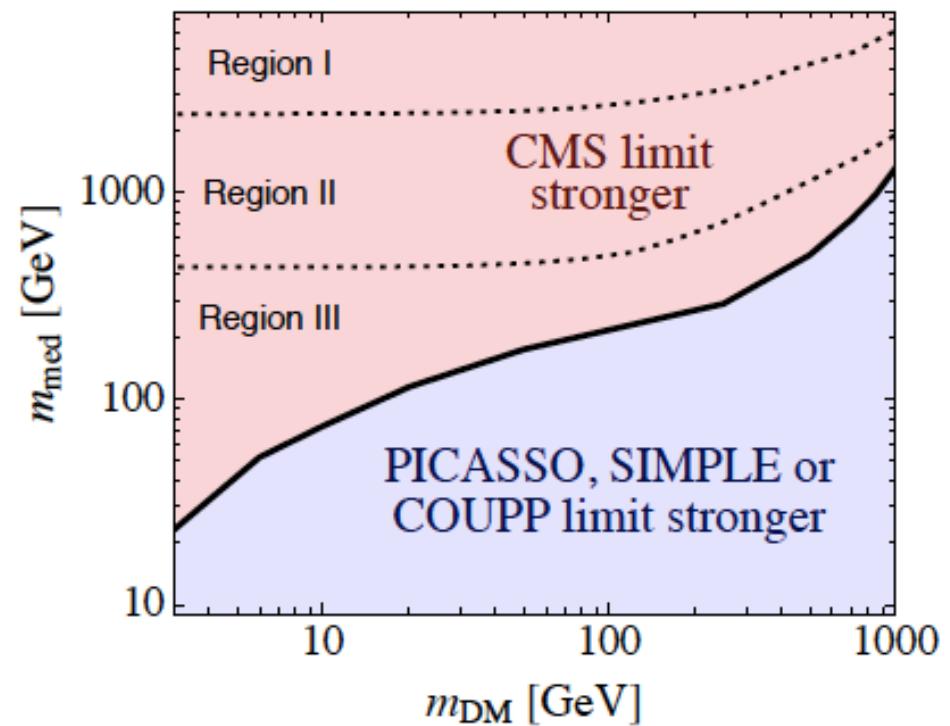
## Beyond EFT limits: Simplified models

Zuchmüller

Working out the complementarity between direct DM detection experiments and collider based DM searches!



EFT limits give the impression that monjet searches outperform direct detection BUT EFT only applies a VERY small class of DM models.

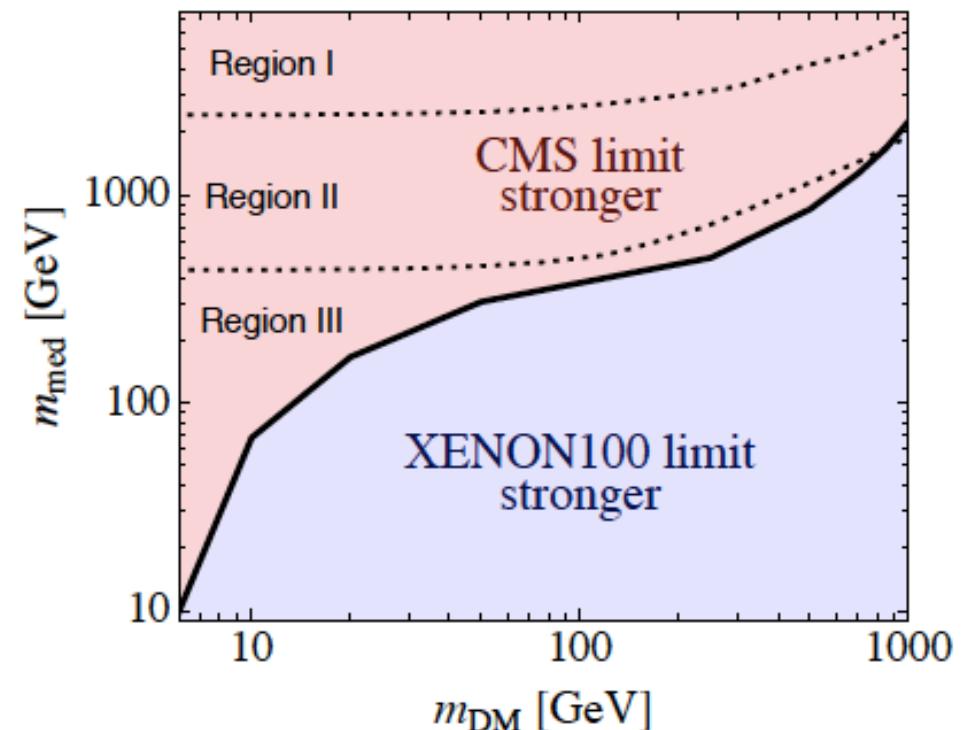
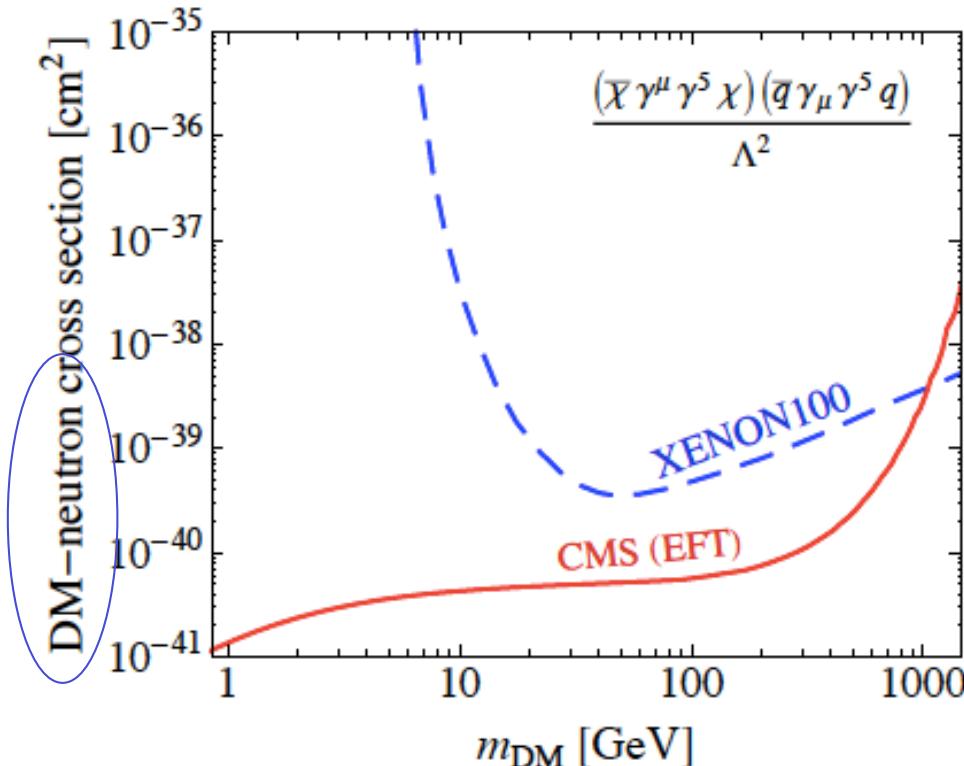


Simplified model limits give a much better Account of the REAL complementarity and thus seem superior for a comparison.

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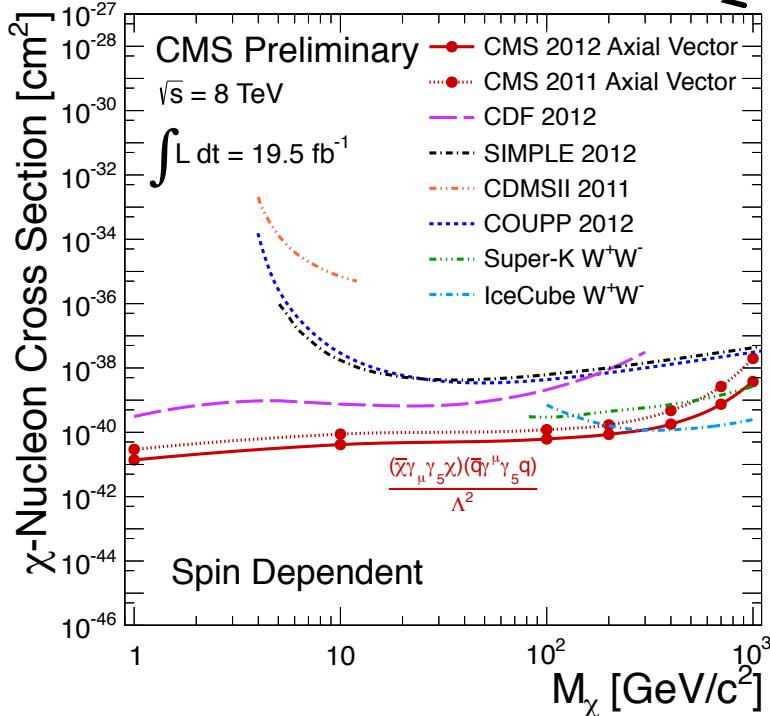
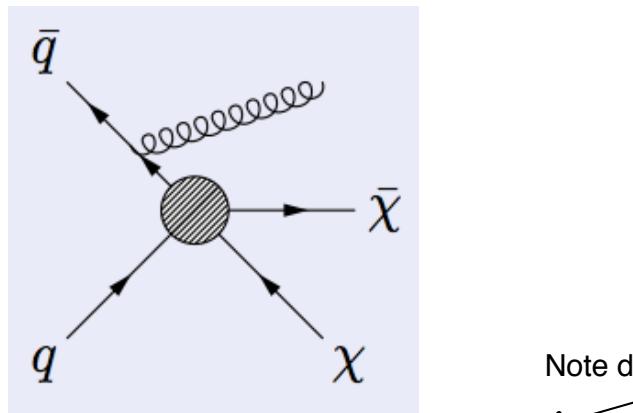
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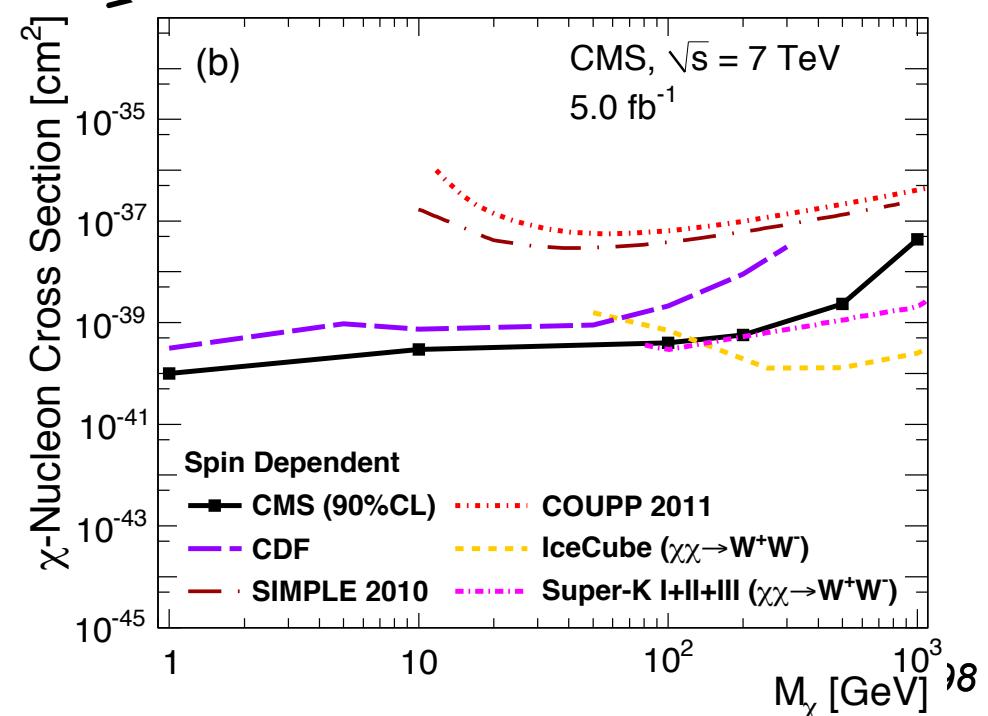
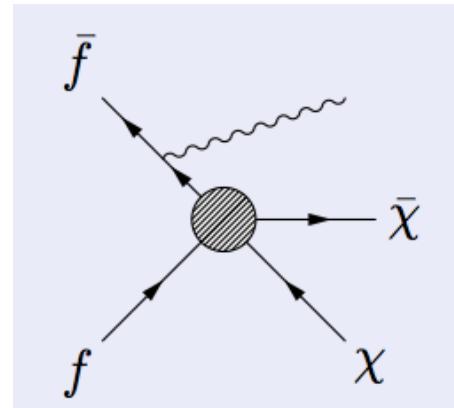
# Monojet and Monophoton (plus $E_T^{\text{miss}}$ )

DM Searches @ collider & Direct Detection O. Buchmüller

**Monojet: hard jet +  $E_T^{\text{miss}}$**



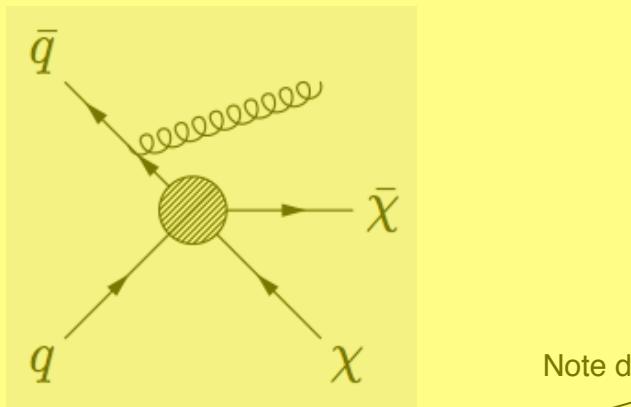
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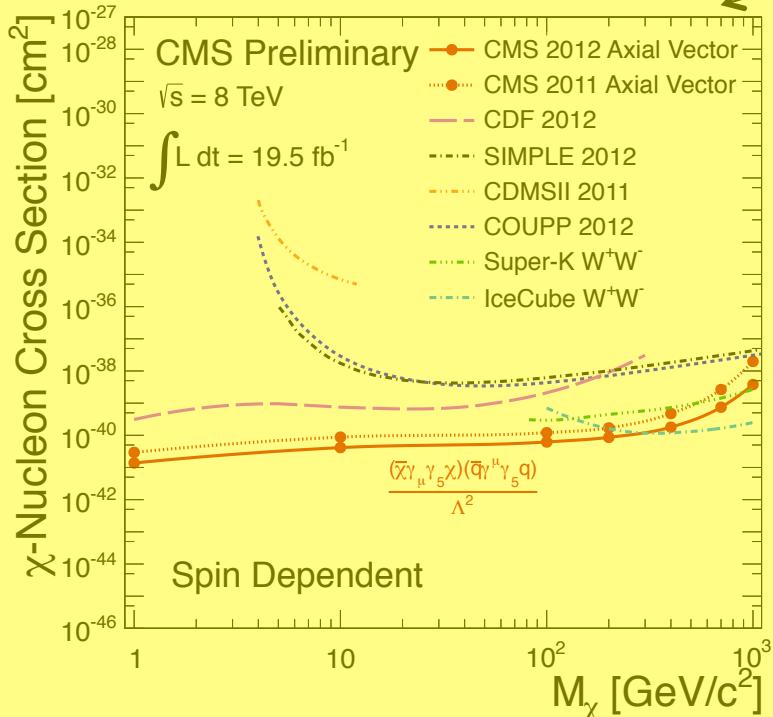
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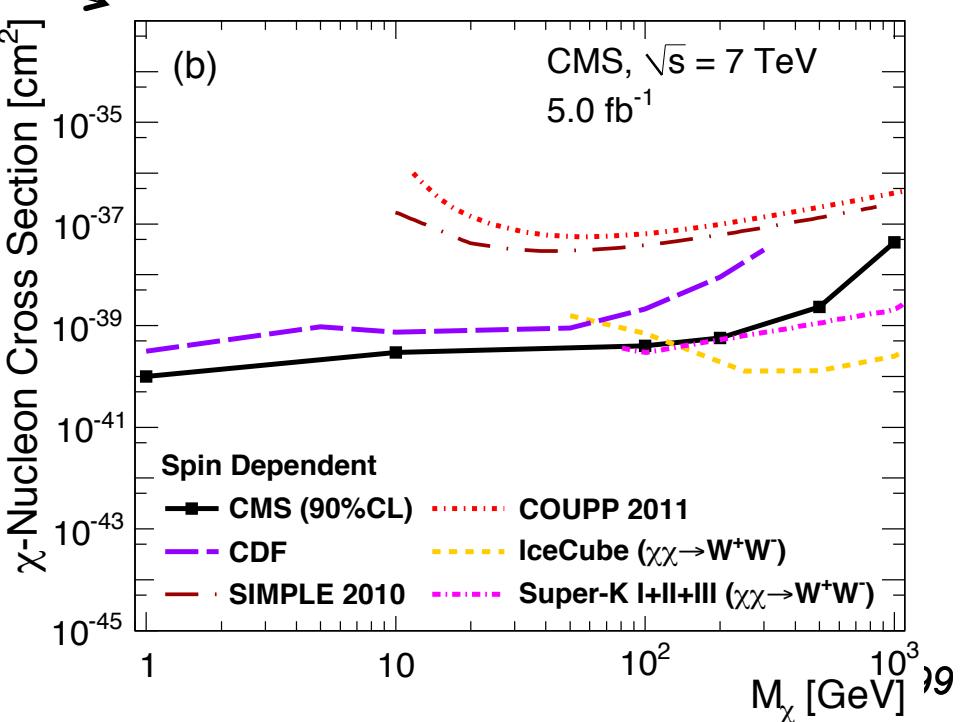
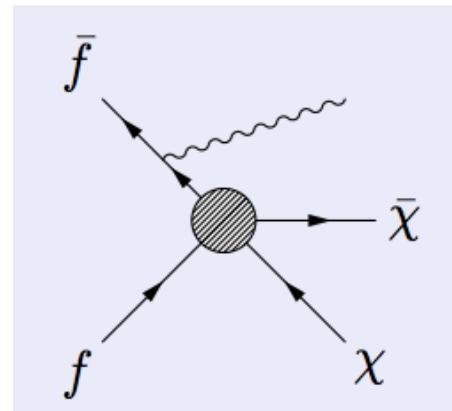
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Note different y scale

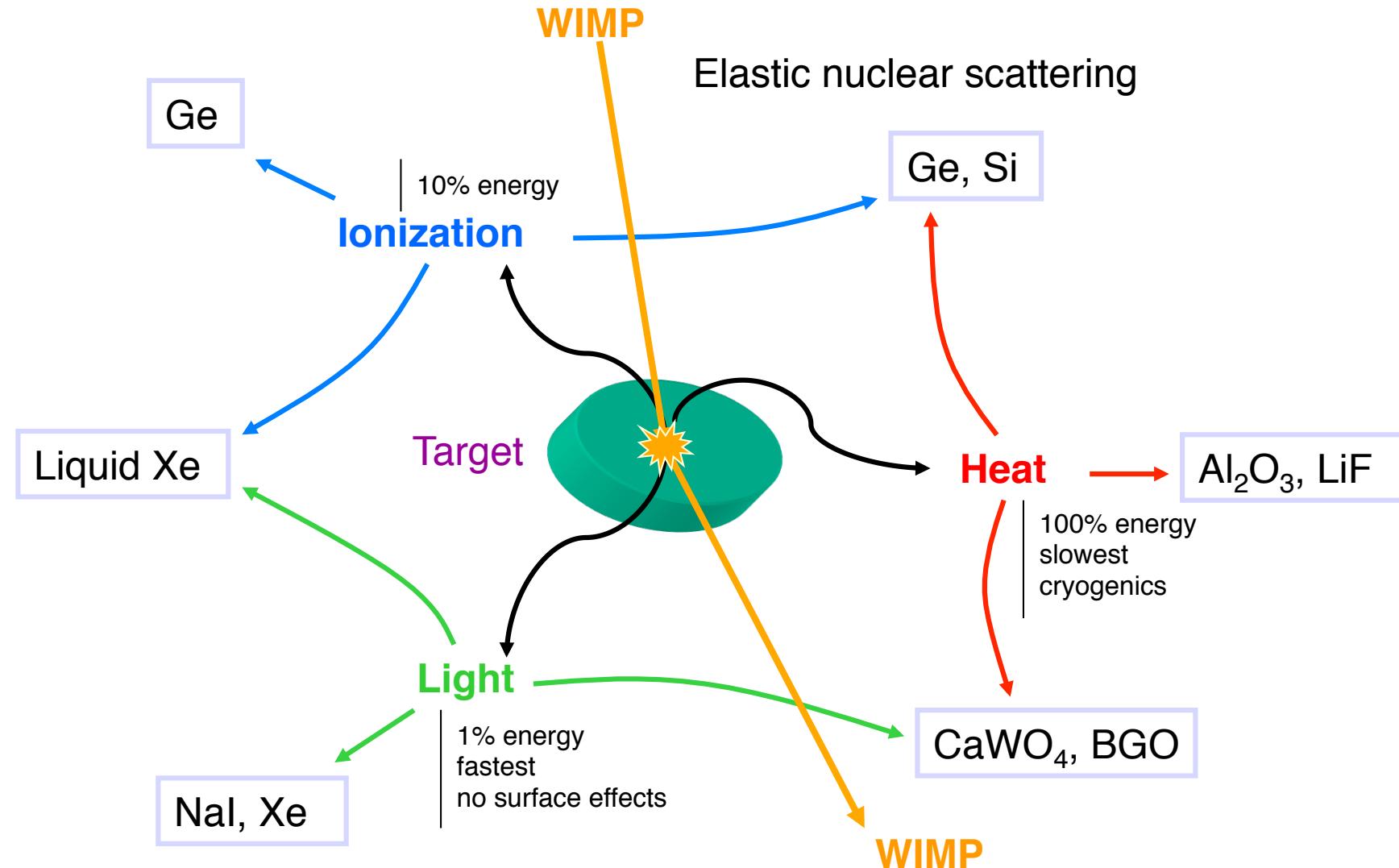


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# Direct Detection Techniques

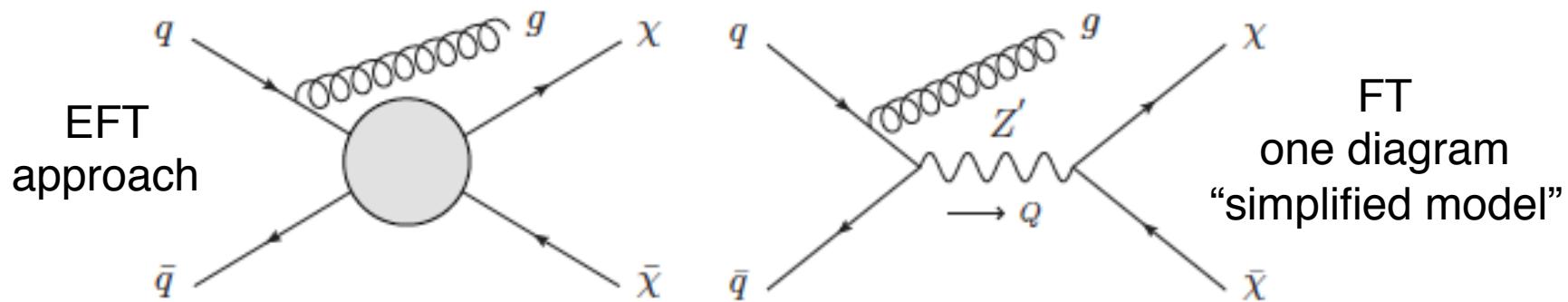
DM Searches @ collider & Direct Detection O. Buchmüller



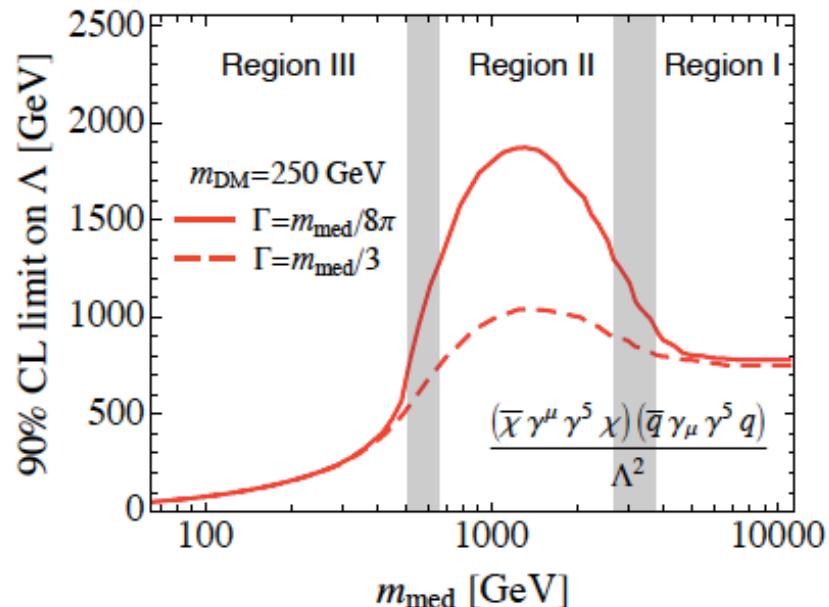
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